

## Introduction

Juyong Zhang School of Mathematics, USTC

#### The Course

- Learn how to solve math problems with tools
- Matlab, Mathematica, Python, Eigen, Ceres, etc...
- Grading policy:
  - Homework&Programming: 80%
  - Final Project: 20%





#### **Covered Topics**

- Image Processing, Image Filtering
- Face Detection, Face Recognition
- Image Stitching, Image Warping
- · Tracking, Optical Flow,
- Stereo Matching, Epipolar Geometry
- Structure From Motion, 3D Surface Reconstruction
- Neural Network, etc...





## TA & QQ Group

- 助教
  - 杨乐园 (ly 1207@mail.ustc.edu.cn)
  - 许子航 (ad123456@mail.ustc.edu.cn)
- 课程QQ群

#### USTC数学实验20...

群号: 827829466





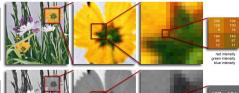


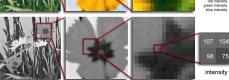
## **Digital Image**

Color images have 3 values per pixel; monochrome images have 1 value per pixel.

a grid of squares, each of which contains a single color

each square is called a pixel (for picture element)









- A digital image, I, is a mapping from a 2D grid of uniformly spaced discrete points,  $\{p = (r,c)\}$ , into a set of positive integer values,  $\{I(p)\}$ , or a set of vector values, e.g.,  $\{[R \ G \ B]^T(p)\}$ .
- At each column location in each row of I there is a value.
- The pair (p, I(p)) is called a "pixel" (for *picture element*).



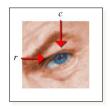


- p = (r,c) is the pixel location indexed by row, r, and column, c.
- I(p) = I(r,c) is the value of the pixel at location p.
- If I(p) is a single number then I is monochrome.
- If I(p) is a vector (ordered list of numbers) then I has multiple bands (e.g., a color image).











Pixel Location: p = (r, c)

Pixel Value: I(p) = I(r, c)

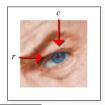
 $\mathsf{Pixel}:[p, \mathit{I}(p)]$ 





 $\mathsf{Pixel}:[p, \mathit{I}(p)]$ 







$$p = (r,c)$$
  
= (row #, col #)  
= (272, 277)







#### Read an Image into Matlab

```
** the chap being which this

To get started, select MATLAR Help or Demog from the Help menu.

>> cd 'Et\images\Animals\People\Famous'
>> I = imread('Les_Boingeoisie.jpg','jpg')!

>> class(I)

ans =

600 1200 3

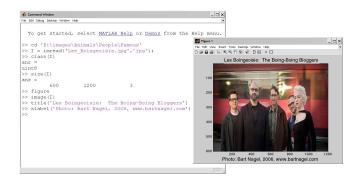
>> figure

>>
```





#### Read an Image into Matlab







### Read an Image into Matlab







## **Crop the Image**





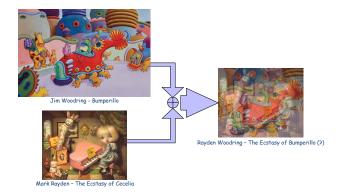


#### **Saving Images as Files**

Assuming that
'I' contains the image of
the correct class,
that
'cmap' is a colormap,
and that
'image\_name' is the
file-name that you want.











```
>> JW = imread('Jim Woodring - Bumperillo.jpg','jpg');
                                                                            Example
>> figure
                                                                            Matlab Code
>> image(JW)
>> truesize
>> title('Bumperillo')
>> xlabel('Jim Woodring')
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
>> image (MR)
>> truesize
>> title('The Ecstasy of Cecelia')
>> xlabel('Mark Ryden')
>> [RMR, CMR, DMR] = size(MR);
>> [RJW.CJW.DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-CMR)/2);
>> JWplusMR = uint8((double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:))+double(MR))/2);
>> figure
>> image(JWplusMR)
>> truesize
>> title('The Ecstasy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
```





```
>> JW = imread('Jim Woodring - Bumperillo.jpg','jpg');
                                                                             Example
>> figure
                                                                             Matlab Code
>> image(JW)
>> truesize
>> title('Bumperillo')
>> xlabel('Jim Woodring')
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
>> image (MR)
>> truesize
>> title('The Ecstasy of Cecelia')
>> xlabel('Mark Ryden')
                                       Cut a section out of the middle of the larger
>> [RMR, CMR, DMR] = size(MR);
                                       image the same size as the smaller image.
>> [RJW.CJW.DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-CMR)/2);
>> JWplusMR = uint8((double JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:))+double(MR))/2);
>> figure
>> image (JWplusMR)
>> truesize
>> title('The Ecstasy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
```





```
>> JW = imread('Jim Woodring - Bumperillo.jpg','jpg');
                                                                            Example
>> figure
                                                                            Matlab Code
>> image(JW)
>> truesize
>> title('Bumperillo')
>> xlabel('Jim Woodring')
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
>> image (MR)
>> truesize
>> title('The Ecstasy of Cecelia')
>> xlabel('Mark Ryden')
                                        Note that the images are averaged, pixelwise.
>> [RMR, CMR, DMR] = size(MR);
>> [RJW.CJW.DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-CMR)/2);
>> JWplusMR = uint8((double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:))+double(MR))/2
>> figure
>> image (JWplusMR)
>> truesize
>> title('The Ecstasy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
```



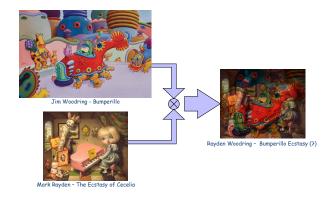


```
>> JW = imread('Jim Woodring - Bumperillo.jpg','jpg');
                                                                             Example
>> figure
                                                                             Matlab Code
>> image(JW)
>> truesize
>> title('Bumperillo')
>> xlabel('Jim Woodring')
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
>> image (MR)
>> truesize
>> title('The Ecstasy of Cecelia')
>> xlabel('Mark Ryden')
>> [RMR, CMR, DMR] = size(MR);
>> [RJW.CJW.DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-QMR)/2);
>> JWplusMR = uint8 ( double (JW (rb: (rb+RMR-1), cb: (cb+CMR-1),:)) - double (MR))/2);
>> figure
                              Note the data class conversions.
>> image(JWplu
>> truesize
>> title('The Ecstasy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
```





#### **Intensity Masking: Multiplying Two Images**







#### **Intensity Masking: Multiplying Two Images**

```
>> JW = imread('Jim Woodring - Bumperillo.ipg','ipg');
                                                                             Example
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
                                                                             Matlab Code
>> [RMR, CMR, DMR] = size(MR);
>> [RJW, CJW, DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-CMR)/2);
>> JWplusMR = uint8((double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:))+double(MR))/2);
>> figure
>> image(JWplusMR)
>> title('The Extacsy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
>> JWtimesMR = double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:)).*double(MR);
>> M = max(JWtimesMR(:));
>> m = min(JWtimesMR(:));
>> JWtimesMR = uint8(255*(double(JWtimesMR)-m)/(M-m));
>> figure
>> image(JWtimesMR)
>> truesize
>> title('EcstasyBumperillo')
```





#### **Intensity Masking: Multiplying Two Images**

```
>> JW = imread('Jim Woodring - Bumperillo.ipg','ipg');
                                                                              Example
>> MR = imread('Mark Ryden - The Ecstasy of Cecelia.jpg','jpg');
                                                                              Matlab Code
>> [RMR, CMR, DMR] = size(MR);
>> [RJW,CJW,DJW] = size(JW);
>> rb = round((RJW-RMR)/2);
>> cb = round((CJW-CMR)/2);
>> JWplusMR = uint8((double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:))+double(MR))/2);
>> figure
                                      Note that the images are multiplied, pixelwise,
>> image(JWplusMR)
>> title('The Extacsy of Bumperillo')
>> xlabel('Jim Woodring + Mark Ryden')
>> JWtimesMR = double(JW(rb:(rb+RMR-1),cb:(cb+CMR-1),:) .*double(MR);
>> M = max(JWtimesMR(:));
>> m = min(JWtimesMR(:));
>> JWtimesMR = uint8 (255*(double(JWtimesMR)-m)/(M-m));
>> figure
>> image(JWtimesMR)
>> truesize
>> title('EcstasyBumperillo')
                                      Note how the image intensities are
                                      scaled back into the range 0-255.
```





## **Pixel Indexing in Matlab**

```
>> I = imread('Lawraa - Flickr - 278635073 883bd891ec o.jpg','jpg');
>> size(I)
ans =
                                                File Edit View Insert Tools Desktop Window Help
   576 768
                                                D # B 4 | 4 | 4 | 4 | 7 | 8 | # | B | B | B | B |
>> r = randperm(576);
>> c = randperm(768);
                                                                               Scrambled Image
>> J = I(r,c,:);
>> figure
>> image(J)
>> truesize
>> title('Scrambled Image')
>> xlabel('What is it?')
```





#### **Point Processing of Images**

- In a digital image, point = pixel.
- Point processing transforms a pixel's value as function of its value alone;
- it does not depend on the values of the pixel's neighbors.





### **Point Processing of Images**

- Brightness and contrast adjustment
- Gamma correction
- Histogram equalization
- Histogram matching
- Color correction.





## **Point Processing**



- gamma



- brightness



original



+ brightness



+ gamma



histogram mod



contrast



original



+ contrast



histogram EQ





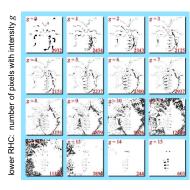
- Let I be a 1-band (grayscale) image.
- $\mathbf{I}(r,c)$  is an 8-bit integer between 0 and 255.
- Histogram,  $h_{\mathbf{I}}$ , of  $\mathbf{I}$ :
  - a 256-element array,  $h_{\rm I}$
  - $h_{\mathbf{I}}(g)$ , for g = 1, 2, 3, ..., 256, is an integer
  - $h_{\mathbf{I}}(g)$  = number of pixels in **I** that have value g-1.







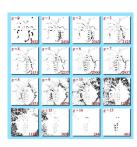
16-level (4-bit) image



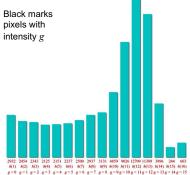
black marks pixels with intensity g





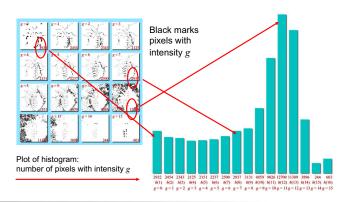


Plot of histogram: number of pixels with intensity *g* 



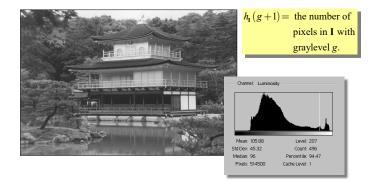
















#### The Histogram of a Color Image

- If I is a 3-band image (truecolor, 24-bit)
- then I(r,c,b) is an integer between 0 and 255.
- Either I has 3 histograms:
  - $h_{\mathbb{R}}(g+1) = \#$  of pixels in  $\mathbf{I}(:,:,1)$  with intensity value g
  - $h_G(g+1) = \#$  of pixels in  $\mathbf{I}(:,:,2)$  with intensity value g
  - $h_{\mathbf{B}}(g+1) = \#$  of pixels in  $\mathbf{I}(:,:,3)$  with intensity value g
- or 1 vector-valued histogram, h(g, 1, b) where
  - h(g+1,1,1) = # of pixels in I with red intensity value g
  - h(g+1,1,2) = # of pixels in **I** with green intensity value g
  - h(g+1,1,3) = # of pixels in **I** with blue intensity value g



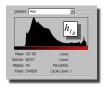


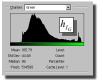
## The Histogram of a Color Image

There is one histogram per color band R, G, & B. Value histogram is from 1 band = (R+G+B)/3















#### **Value or Luminance Histograms**

The Value histogram of a 3-band (truecolor) image, **I**, is the histogram of the value image,

$$V(r,c) = \frac{1}{3} [\mathbf{R}(r,c) + \mathbf{G}(r,c) + \mathbf{B}(r,c)]$$

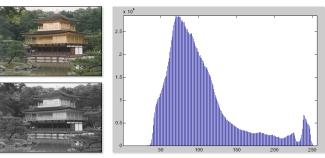
Where **R**, **G**, and **B** are the red, green, and blue bands of **I**. The luminance histogram of **I** is the histogram of the luminance image,

$$\mathbf{L}(r,c) = 0.299 \cdot \mathbf{R}(r,c) + 0.587 \cdot \mathbf{G}(r,c) + 0.114 \cdot \mathbf{B}(r,c)$$





# Value Histogram



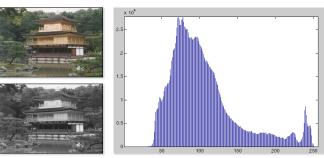


Histogram of the value image.





# **Luminance Histogram**



Luminance image, L.

Histogram of the luminance image.





#### Multi-Band Histogram Calculator in Matlab

```
% Multi-band histogram calculator
function h=histogram(I)

[R C B]=size(I);
% allocate the histogram
h=zeros(256,1,B);
% range through the intensity values
for g=0:255
   h(g+1,1,:) = sum(sum(I==g)); % accumulate
end
return;
```





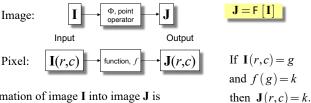
#### Multi-Band Histogram Calculator in Matlab

```
% Multi-band histogram calc
                                    Loop through all intensity levels (0-255)
function h=histogram(I)
                                    Tag the elements that have value g.
                                    The result is an RxCxB logical array that
[R C B]=size(I);
                                    has a 1 wherever I(r,c,b) = g and 0's
                                    evervwhere else.
                                    Compute the number of ones in each band of
% allocate the histogram
                                    the image for intensity g.
h=zeros(256,1,B);
                                    Store that value in the 256x1xB histogram
                                    at h(g+1,1,b).
% range through the intensi
for q=0:255
   h(q+1,1,:) = sum(sum(I==q)); % accumulate
end
                                      sum (sum (I==g)) computes one
If B==3, then h(g+1,1,:) contains
                                      number for each band in the image.
3 numbers: the number of pixels in
bands 1, 2, & 3 that have intensity q.
```





#### **Point Ops via Functional Mappings**



The transformation of image **I** into image **J** is accomplished by replacing each input intensity, g, with a specific output intensity, k, at every location (r,c) where  $\mathbf{I}(r,c) = g$ .

The rule that associates k with g is usually specified with a function, f, so that f(g) = k.





## **Point Ops via Functional Mappings**

#### One-band Image

 $\mathbf{J}(r,c) = f(\mathbf{I}(r,c)),$  for all pixels locations (r,c).

#### Three-band Image

$$\begin{split} \mathbf{J}(r,c,b) &= f\left(\mathbf{I}(r,c,b)\right), \text{ or } \\ \mathbf{J}(r,c,b) &= f_b\left(\mathbf{I}(r,c,b)\right), \\ \text{for } b &= 1,2,3 \text{ and all } (r,c). \end{split}$$





#### **Point Ops via Functional Mappings**

#### One-band Image

Either all 3 bands are mapped through the same function, f, or ...

Three-band Image

 $\mathbf{J}(r,c) = f(\mathbf{I}(r,c)),$ for all pixels locations (r,c).

$$\mathbf{J}(r,c,b) = f(\mathbf{I}(r,c,b)), \text{ or }$$

$$\mathbf{J}(r,c,b) = f_b(\mathbf{I}(r,c,b)),$$

for b=1,2,3 and all (r,c) mapped through a separate function,  $f_{\rm b}$ .





#### **Point Operations using Look-up Tables**

A look-up table (LUT) implements a functional mapping.

If k = f(g), for g = 0,...,255, and if k takes on values in  $\{0,...,255\}$ , ... ... then the LUT that implements f is a 256x1 array whose (g+1)<sup>th</sup> value is k = f(g).





#### **Point Operations using Look-up Tables**

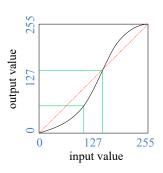
#### If I is 3-band, then

- a) each band is mapped separately using the same LUT for each band or
- b) each band is mapped using different LUTs one for each band.
- a) J = LUT(I+1), or
- b)  $\mathbf{J}(:,:,b) = \text{LUT}_b(\mathbf{I}(:,:,b)+1)$  for b = 1,2,3.





#### **Point Operations = Look-up Table Ops**



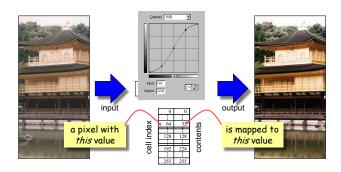
<i>E.g.</i> :	index	value
	101	64
	102	68
	103	69
	104	70
	105	70
	106	71

input output





## **Look-Up Tables**







## **Point Processes: Original Image**



Luminance Histogram

Kinkaku-ji (金閣寺 Temple of the Golden Pavilion), also known as Rokuon-ji (鹿苑寺 Deer Garden Temple), is a Zen Buddhist temple in Kyoto, Japan.

Photo by Richard Alan Peters II, August 1993.

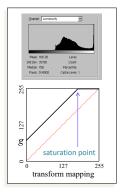




#### **Point Processes: Increase Brightness**



$$\mathbf{J}(r,c,b) = \begin{cases} \mathbf{I}(r,c,b) + g, & \text{if } \mathbf{I}(r,c,b) + g < 256 \\ 255, & \text{if } \mathbf{I}(r,c,b) + g > 255 \end{cases}$$
  $g \ge 0 \text{ and } b \in \{1,2,3\} \text{ is the band index.}$ 



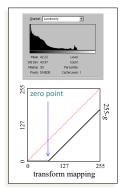




#### **Point Processes: Decrease Brightness**



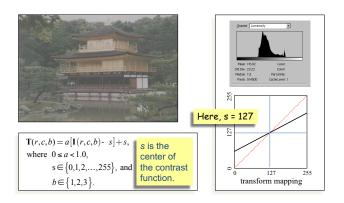
$$\mathbf{J}(r,c,b) = \begin{cases} 0, & \text{if } \mathbf{I}(r,c,b) - g < 0 \\ \mathbf{I}(r,c,b) - g, & \text{if } \mathbf{I}(r,c,b) - g > 0 \end{cases}$$
 
$$g \ge 0 \text{ and } b \in \left\{1,2,3\right\} \text{ is the band index.}$$







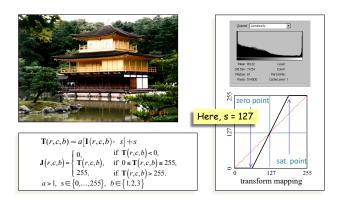
#### **Point Processes: Decrease Contrast**







#### **Point Processes: Increase Contrast**



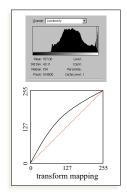




#### **Point Processes: Increased Gamma**



$$\mathbf{J}(r,c) = 255 \cdot \left[ \frac{\mathbf{I}(r,c)}{255} \right]^{\frac{1}{\gamma}} \text{ for } \gamma > 1.0$$



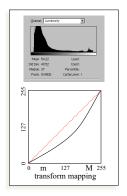




#### **Point Processes: Decreased Gamma**



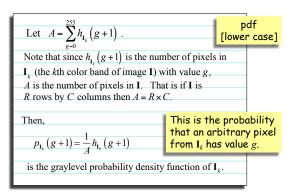
$$\mathbf{J}(r,c) = 255 \cdot \left[ \frac{\mathbf{I}(r,c)}{255} \right]^{\frac{1}{\gamma}} \text{ for } \gamma < 1.0$$







## The Probability Density Function of an Image







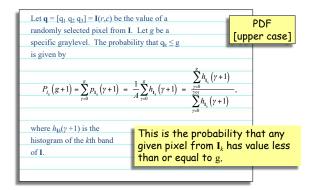
## The Probability Density Function of an Image

- p<sub>band</sub>(g+1) is the fraction of pixels in (a specific band of) an image that have intensity value g.
- p<sub>band</sub>(g+1) is the probability that a pixel randomly selected from the given band has intensity value g.
- Whereas the sum of the histogram  $h_{band}(g+1)$  over all g from 1 to 256 is equal to the number of pixels in the image, the sum of  $p_{band}(g+1)$  over all g is 1.
- $p_{\text{band}}$  is the normalized histogram of the band.





#### The Probability Distribution Function of an Image







#### **Point Processes: Histogram Equalization**

Task: remap image I so that its histogram is as close to constant as possible

Let  $P_{\mathbf{I}}(g+1)$  be the probability distribution function of  $\mathbf{I}$ .

Then J has, as closely as possible, the correct histogram if

$$\mathbf{J}(r,c,b) = 255 \cdot P_{\mathbf{I}} [\mathbf{I}(r,c,b) + 1].$$

The PDF itself is used as the LUT.

all bands processed similarly

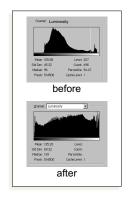




#### **Point Processes: Histogram Equalization**



$$J(r,c,b) = 255 \cdot P_I(g+1),$$
  
 $g = I(r,c,b), b \in \{1,2,3\}.$ 



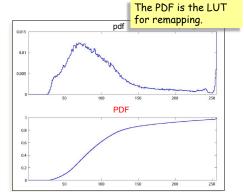




# **Histogram EQ**







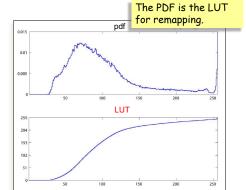




# **Histogram EQ**



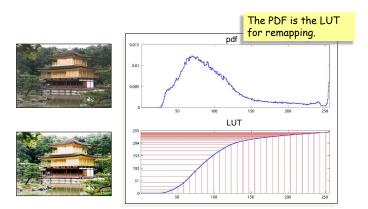








# **Histogram EQ**







#### **Point Processes: Histogram Equalization**

Task: remap image I with min =  $m_I$  and max =  $M_I$  so that its histogram is as close to constant as possible and has min =  $m_{\rm I}$ and max =  $M_{\rm J}$ .

Let  $P_{\mathbf{I}}(\mathbf{g}+1)$  be the probability distribution function of **I**.

Then **J** has, as closely as possible, the correct histogram if





## **Point Processes: Histogram Matching**

Task: remap image **I** so that it has, as closely as possible, the same histogram as image **J**.

Because the images are digital it is not, in general, possible to make  $h_I \equiv h_I$ . Therefore,  $p_I \neq p_I$ .

Q: How, then, can the matching be done?

A: By matching percentiles.





#### **Matching Percentiles**

... assuming a 1-band image or a single band of a color image.

#### Recall:

- The PDF of image I is such that  $0 \le P_{I}(g_{I}) \le 1$ .
- $P_{\mathbf{I}}(g_{\mathbf{I}}+1) = c$  means that c is the fraction of pixels in  $\mathbf{I}$  that have a value less than or equal to  $g_{\mathbf{I}}$ .
- 100c is the percentile of pixels in I that are less than or equal to  $g_{I}$ .

To match percentiles, replace all occurrences of value  $g_I$  in image I with the value,  $g_J$ , from image J whose percentile in J most closely matches the percentile of  $g_I$  in image I.





#### **Matching Percentiles**

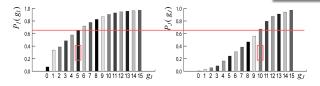
... assuming a 1-band image or a single band of a color image.

So, to create an image, K, from image I such that K has nearly the same PDF as image J do the following:

If 
$$\mathbf{I}(r,c) = g_{\mathbf{I}}$$
 then let  $K(r,c) = g_{\mathbf{J}}$  where  $g_{\mathbf{J}}$  is such that

$$P_{\mathbf{I}}(g_{\mathbf{I}}) > P_{\mathbf{J}}(g_{\mathbf{J}} - I) \text{ AND } P_{\mathbf{I}}(g_{\mathbf{I}}) \leq P_{\mathbf{J}}(g_{\mathbf{J}}).$$









#### **Histogram Matching Algorithm**

```
\begin{split} [R,C] &= \text{size}\left(\mathbf{I}\right); \\ \mathbf{K} &= \text{zeros}\left(R,C\right); \\ g_{\mathbf{J}} &= m_{\mathbf{J}}; \\ \text{for } g_{\mathbf{I}} &= m_{\mathbf{I}} \text{ to } M_{\mathbf{I}} \\ &\quad \text{while } g_{\mathbf{J}} < 255 \text{ AND } P_{\mathbf{I}}\left(g_{\mathbf{I}}+\mathbf{I}\right) < 1 \text{ AND } \\ &\quad P_{\mathbf{J}}\left(g_{\mathbf{J}}+\mathbf{I}\right) < P_{\mathbf{I}}\left(g_{\mathbf{I}}+\mathbf{I}\right) \\ g_{\mathbf{J}} &= g_{\mathbf{J}}+\mathbf{I}; \\ &\quad \text{end} \\ &\quad \mathbf{K} = \mathbf{K} + \left[g_{\mathbf{J}} \cdot \left(\mathbf{I} == g_{\mathbf{I}}\right)\right] \\ \text{end} \end{split}
```

Assuming a 1-band image or a single band of a color image.

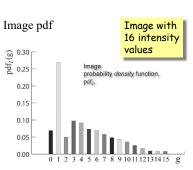
# This directly matches image I to image J.

```
P_{\mathbf{I}}(g_{\mathbf{I}}+1): PDF of \mathbf{I}
P_{\mathbf{J}}(g_{\mathbf{J}}+1): PDF of \mathbf{J}.
m_{\mathbf{J}} = \min \mathbf{J},
m_{\mathbf{I}} = \max \mathbf{J},
m_{\mathbf{I}} = \min \mathbf{I},
M_{\mathbf{I}} = \max \mathbf{I}.
```

Better to use a LUT.





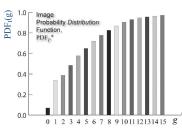








## **Image PDF**

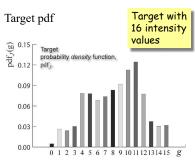










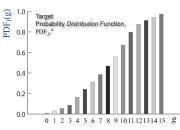








## **Target PDF**



\*a.k.a Cumulative Distribution Function, CDF<sub>1</sub>.







## **Histogram Matching with a Lookup Table**

Often it is faster or more versatile to use a lookup table (LUT). Rather than remapping each pixel in the image separately, one can create a table that indicates to which target value each input value should be mapped. Then

$$\mathbf{K} = \text{LUT}[\mathbf{I} + 1]$$

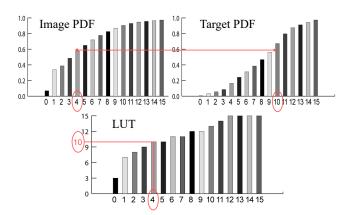
In *Matlab* if the LUT is a 256  $\times$  1 matrix with values from 0 to 255 and if image **I** is of type **uint8**, it can be remapped with the following code:

$$K = uint8(LUT(I+1));$$





#### **LUT Creation**







## **Look Up Table for Histogram Matching**

```
LUT = zeros(256.1);
g_{\mathbf{I}} = 0;
for g_{\rm I} = 0 to 255
    while P_{I}(g_{I}+1) < P_{I}(g_{I}+1) AND g_{I} < 255
         g_1 = g_1 + 1;
    end
    LUT(g_I + 1) = g_I;
end
```

This creates a look-up table which can then be used to remap the image.

 $P_{\mathbf{I}}(g_{\mathbf{I}}+1)$ : PDF of  $\mathbf{I}$ ,

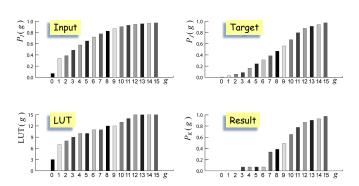
 $P_{\mathbf{J}}(g_{\mathbf{J}}+1)$ : PDF of  $\mathbf{J}$ ,

 $LUT(g_1 + 1)$ : Look- Up Table





#### Input & Target PDFs, LUT and Resultant PDF







### **Example: Histogram Matching**







target

ginal



remapped







# **Image Convolution**

Juyong Zhang School of Mathematics, USTC

### **Spatial Filtering**

Let **I** and **J** be images such that J = T[I].

 $T[\cdot]$  represents a transformation, such that,

$$\mathbf{J}(r,c) = \mathbf{T}[\mathbf{I}](r,c) = f(\{\mathbf{I}(\rho,\chi) | \rho \in \{r-s,...,r,...r+s\}, \chi \in \{c-d,...,c,...c+d\}\}).$$

That is, the value of the transformed image, **J**, at pixel location (r,c) is a function of the values of the original image, **I**, in a  $2s+1 \times 2d + 1$  rectangular neighborhood centered on pixel location (r,c).





### **Moving Windows**

- The value, J(r,c) = T[I](r,c), is a function of a rectangular neighborhood centered on pixel location (r,c) in I.
- There is a different neighborhood for each pixel location, but if the dimensions of the neighbor-hood are the same for each location, then transform T is sometimes called a moving window transform.







We'll take a section of this image to demonstrate the

photo: R.A.Peters II, 1999

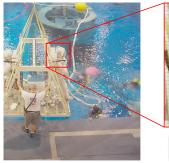








Pixelize the section to better see the effects.



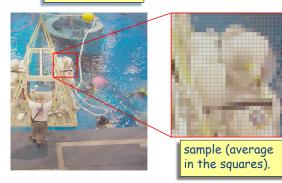


apply a pixel grid





Pixelize the section to better see the effects.





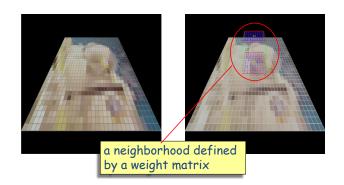






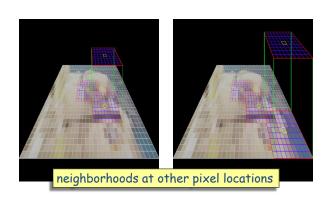










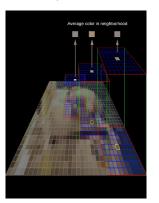






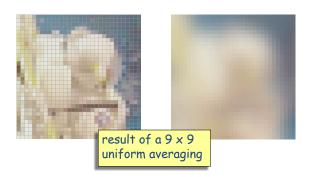
# Linear Moving-Window Transformations (i.e. convolution)

The output of the transform at each pixel is the (weighted) average of the pixels in the neighborhood.













### **Convolution: Mathematical Representation**

If a MW transformation is *linear* then it is a *convolution*:

$$\mathbf{J}(r,c) = \left[\mathbf{I} * \mathbf{h}\right](r,c) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathbf{I}(r-\rho,c-\chi)\mathbf{h}(\rho,\chi) d\rho d\chi,$$

for a real image  $(I: R \times R \rightarrow R)$ , or for a digital image  $(I: Z \times Z \rightarrow Z)$ :

$$\mathbf{J}(r,c) = \left[\mathbf{I} * \mathbf{h}\right](r,c) = \sum_{\rho = -s}^{s} \sum_{\chi = -d}^{d} \mathbf{I}(r - \rho, c - \chi) \mathbf{h}(\rho, \chi)$$





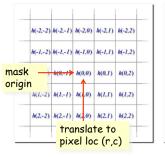
### Convolution Mask (Weight Matrix)

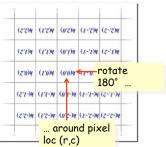
- The object,  $\mathbf{h}(\rho, \chi)$ , in the equation is a weighting function, or in the discrete case, a rectangular matrix of numbers.
- The matrix is the moving window.
- Pixel (r,c) in the output image is the weighted sum of pixels from the original image in the neighborhood of (r,c) traced by the matrix.
- Each pixel in the neighborhood of (r,c) is multiplied by the corresponding matrix value after the matrix is rotated by 180°.
- The sum of those products is the value of pixel (r,c) in the output image





### **Convolution Masks: Moving Window**

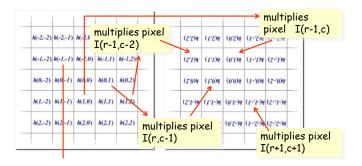








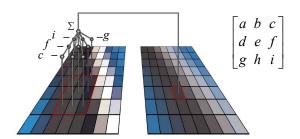
### **Convolution Masks: Moving Window**







### **Convolution by Moving Window**







#### Another example



original



3x3 average

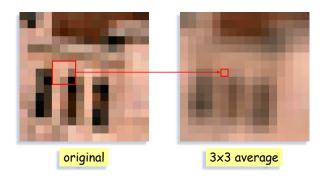






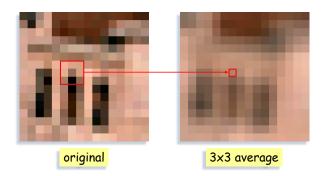


















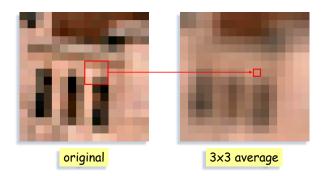








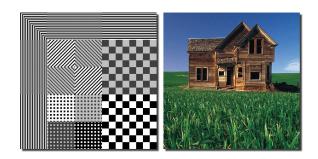








# **Convolution Examples: Original Images**

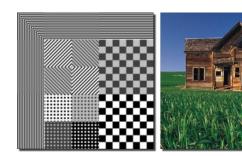






# Convolution Examples: $3 \times 3$ Blur

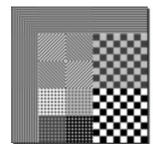
 $\frac{1}{9} \begin{bmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{bmatrix}$ 







### Convolution Examples: $5 \times 5$ Blur

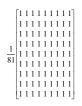


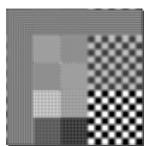


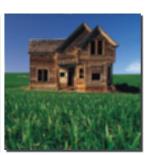




### **Convolution Examples: 9×9 Blur**



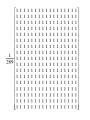


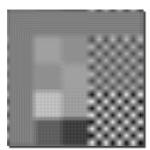


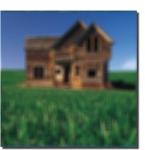




# **Convolution Examples: 17 × 17 Blur**



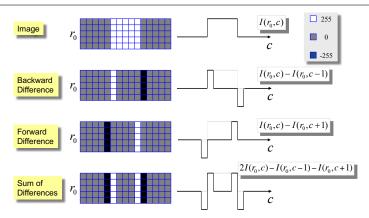








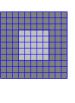
### **Vertical Edge Detection**







### **Symmetric Edge Detection**









$$2I(r,c) - I(r,c-1)$$
$$-I(r,c+1)$$

$$2I(r,c) - I(r-1,c)$$
  
-  $I(r+1,c)$ 

$$4I(r,c) - I(r-1,c) - I(r+1,c) - I(r,c-1) - I(r,c+1)$$

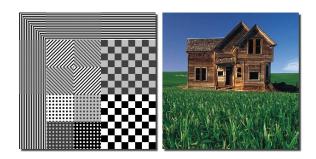








# **Convolution Examples: Original Images**

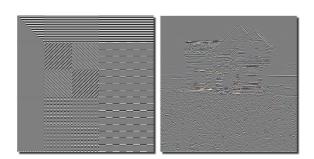






### **Convolution Examples: Vertical Difference**

 $\begin{bmatrix} -1\\2\\-1 \end{bmatrix}$ 

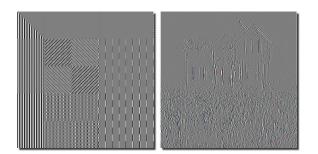






### **Convolution Examples: Horizontal Difference**

 $\begin{bmatrix} -1 & 2 & -1 \end{bmatrix}$ 

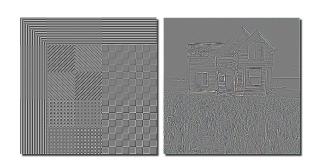






## **Convolution Examples: H + V Diff.**

$$\begin{bmatrix} 0 & -1 & 0 \\ -1 & 4 & -1 \\ 0 & -1 & 0 \end{bmatrix}$$

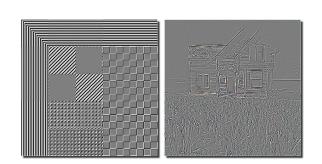






## **Convolution Examples: Diagonal Difference**

$$\begin{bmatrix} -1 & 0 & 0 \\ 0 & 2 & 0 \\ 0 & 0 & -1 \end{bmatrix}$$

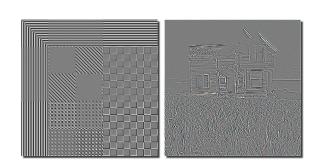






## **Convolution Examples: Diagonal Difference**

$$\begin{bmatrix} 0 & 0 & -1 \\ 0 & 2 & 0 \\ -1 & 0 & 0 \end{bmatrix}$$

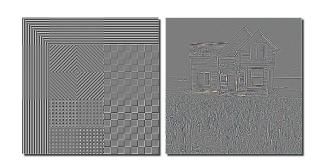






# **Convolution Examples: D + D Difference**

$$\begin{bmatrix} -1 & 0 & -1 \\ 0 & 4 & 0 \\ -1 & 0 & -1 \end{bmatrix}$$

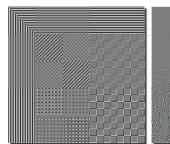






## Convolution Examples: H + V + D Diff.

$$\begin{bmatrix} -1 & -1 & -1 \\ -1 & 8 & -1 \\ -1 & -1 & -1 \end{bmatrix}$$









#### The Median Filter

- Returns the median value of the pixels in a neighborhood
- Is non-linear
- Is similar to a uniform blurring filter which returns the mean value of the pixels in a neighborhood of a pixel
- Unlike a mean value filter the median tends to preserve step edges







#### **Median Filter: General Definition**

$$\text{med} \big\{ I, Z \big\} \big( p \big) \! = \! \underset{q \in \text{supp}(Z + p)}{\text{median}} \big\{ I \big( q \big) \big\}$$

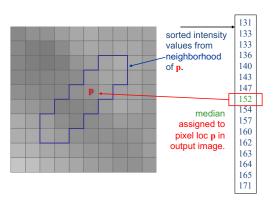
#### This can be computed as follows:

- 1. Let I be a monochrome (1-band) image.
- 2. Let **Z** define a neighborhood of arbitrary shape.
- 3. At each pixel location,  $\mathbf{p} = (r,c)$ , in  $\mathbf{I}$  ...
- 4. ... select the n pixels in the **Z**-neighborhood of **p**,
- 5. ... sort the *n* pixels in the neighborhood of **p**, by value, into a list L(j) for j = 1,...,n.
- 6. The output value at **p** is L(n), where  $m = \lfloor \frac{n}{2} \rfloor + 1$





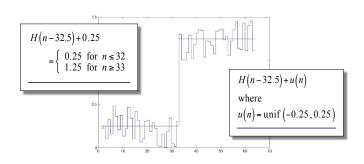
#### **Median Filter: General Definition**







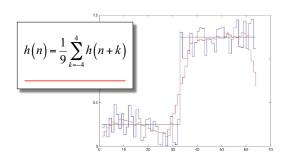
### A Noisy Step Edge







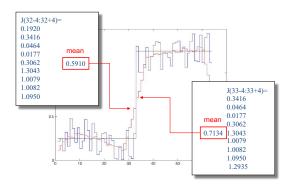
### **Blurred Noisy 1D Step Edge**







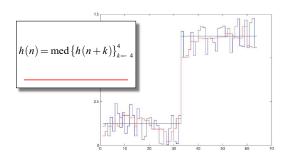
## **Blurred Noisy 1D Step Edge**







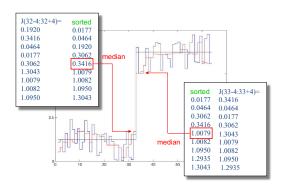
### Median Filtered Noisy 1D Step Edge







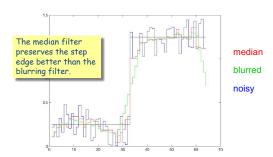
#### **Median Filtered Noisy 1D Step Edge**







#### Median vs. Blurred







## **Median Filtering of Binary Images**







Original





## **Median Filtering of Binary Images**



Median Filtered Noisy



Original











Original











Noisy







3x3-blur x 1



3x3-median x 1







3x3-blur x 2



3x3-median x 2







3x3-blur x 3



3x3-median x 3







3x3-blur x 4



3x3-median x 4







3x3-blur x 5



3x3-median x 5







3x3-blur x 10



3x3-median x 10





### **Limit and Root Images**

Fact: if you repeatedly filter an image with the same blurring filter or median filter, eventually the output does not change. That is, let

$$\mathbf{I}[*\mathbf{h}]^{\sharp} \equiv (((\mathbf{I}*\mathbf{h})*\mathbf{h})\cdots *\mathbf{h}), \quad k \text{ times, and}$$

$$\mathbf{I}[\operatorname{med} \mathbf{Z}]^{\sharp} \equiv (((\mathbf{I} \operatorname{med} \mathbf{Z}) \operatorname{med} \mathbf{Z})\cdots \operatorname{med} \mathbf{Z}), \quad k \text{ times.}$$

Then

$$\lim_{k \to \infty} \mathbf{I} \left[ *\mathbf{h} \right]^k = \mathbf{I} \left[ *\mathbf{h} \right]^n = \mathbf{I}_0, \text{ and}$$
$$\lim_{k \to \infty} \mathbf{I} \left[ \text{med } \mathbf{Z} \right]^k = \mathbf{I} \left[ \text{med } \mathbf{Z} \right]^m = \mathbf{I}_r,$$

where n and m are integers  $(< \infty)$ ,  $I_0$  is a single-valued image and  $I_r$  is called the *median root* of I.





## **Limit and Root Images**



3x3-blur x 10



3x3-median x 10





## **Limit and Root Images**



 $3 \times 3$ -blur  $\times n \to \infty$ 



3x3-median root





### Median Filter Algorithm in Matlab

```
function D = median filt(I,SE,origy,origx)
[R,C] = size(I); % assumes 1-band image
[SER.SEC] = size(SE); % SE < 0 not in nbhd
N = sum(sum(SE>=0)); % no. of pixels in nbhd
A = -ones(R+SER-1,C+SEC-1,N); % accumulator
n=1; % copy I into band n of A for nbhd pix n
for j = 1 : SER % neighborhood is def'd in SE
   for i = 1 : SEC
      if SE(j,i) >= 0 % then is a nbhd pixel
         A(j:(R+j-1),i:(C+i-1),n) = I;
         n=n+1; % next accumulator band
      end
   end
% pixel-wise median across the bands of A
A = shiftdim(median(shiftdim(A, 2)), 1);
D = A(origy:(R+origy-1), origx:(C+origx-1));
return:
```





#### **Vector Median Filter**

A vector median filter selects from among a set of vectors, the one vector that is closest to all the others.

That is, if *S* is a set of vectors, in  $\mathbb{F}^n$  the median,  $\overline{\mathbf{v}}$ , is

$$\overline{\mathbf{v}} = \underset{k \neq j}{\operatorname{arg\,min}} \Big\{ \left\| \left. \mathbf{v}_k - \mathbf{v}_j \right\| \; \middle| \; \mathbf{v}_k, \mathbf{v}_j \in S \; \right\}.$$

 $riangleleft \mathbb{F}^n$  is an n-dimensional linear vector space over the field,  $\mathbb{F}$ .)

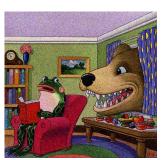








Jim Woodring - A Warm Shoulder



Sparse noise, 32% coverage in each band







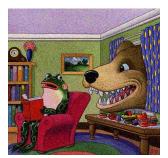
3 × 3 color median filter applied once



3 × 3 color median filter applied twice







Sparse noise, 32% coverage in each band

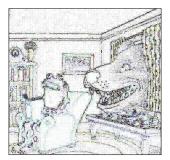


Jim Woodring – A Warm Shoulder





Absolute differences displayed as negatives to enhance visibility



(3 × 3 CMF2 of noisy) - original



(3×3 CMF2 of noisy) - (3×3 CMF2 of original)





#### CMF vs. Standard Median on Individual Bands

A color median filter has to compute the distances between all the color vectors in the neighborhood of each pixel. That's expensive computationally.

- **Q:** Why not simply take the 1-band median of each color band individually?
- A: The result at a pixel could be a color that did not exist in the pixel's neighborhood in the input image. The result is not the median of the colors it is the median of the intensities of each color band treated independently.
- Q: Is that a problem?
- A: Maybe. Maybe not. It depends on the application. It may make little difference visually. If the colors need to be preserved, it could be problematic.

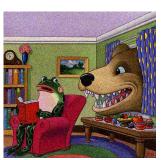




#### CMF vs. Standard Median on Individual Bands



Jim Woodring - A Warm Shoulder



Sparse noise, 32% coverage in each band





#### CMF vs. Standard Median on Individual Bands



3 × 3 color median filter applied once



3 × 3 color median filter applied twice





#### CMF vs. Standard Median on Individual Bands



 $3 \times 3$  median filter applied to each band once



3 × 3 median filter applied to each band twice







# **Gradient Image Processing**

Juyong Zhang School of Mathematics, USTC

### **Today: Gradient manipulation**

#### Idea:

- Human visual system is very sensitive to gradient
- Gradient encode edges and local contrast quite well
- Do your editing in the gradient domain
- Reconstruct image from gradient



 Various instances of this idea, 1'll mostly follow Perez et al. Siggraph 2003 http://research.microsoft.com/vision/cambridge/papers/perez siggraph03.pdf





### **Problems with direct cloning**





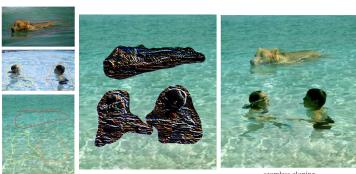
cloning

From Perez et al. 2003





## Solution: clone gradient









### Gradients and grayscale images

- Grayscale image: n × n scalars
- Gradient: n × n 2D vectors
- Overcomplete!
- What's up with this?
- Not all vector fields are the gradient of an image!
- Only if they are curl-free (a.k.a. conservative)
  - But it does not matter for us





### Today message I

• Manipulating the gradient is powerful





#### **Today message II**

- · Optimization is powerful
  - In particular least square
- Good Least square optimization reduces to a big linear system
- Linear algebra is your friend
  - Big sparse linear systems can be solved efficiently





#### **Today message III**

- Toy examples are good to further understanding
- 1D can however be overly simplifying, n-D is much more complicated





#### **Seamless Poisson cloning**

• Given vector field v (pasted gradient), find the value of f in unknown region that optimize:

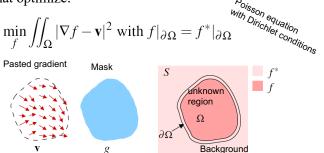


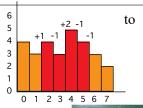
Figure 1: Guided interpolation notations. Unknown function f interpolates in domain  $\Omega$  the destination function  $f^*$ , under guidance of vector field  $\mathbf{v}$ , which might be or not the gradient field of a source function g.

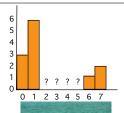




### Discrete 1D example: minimization

• Copy





- Min  $((f_2-f_1)-1)^2$
- Min  $((f_3-f_2)-(-1))^2$
- Min  $((f_4-f_3)-2)^2$
- Min  $((f_5-f_4)-(-1))^2$
- Min  $((f_6-f_5)-(-1))^2$



With

$$f_1 = 6$$

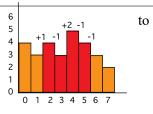
$$f_6 = 1$$

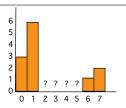




### 1D example: minimization

Copy





• Min 
$$((f_2-6)-1)^2$$

$$=> f_2^2 + 49 - 14f_2$$

• Min 
$$((f_3-f_2)-(-1))^2$$

$$Min ((f_3-f_2)-(-1))^2 = f_3^2+f_2^2+1-2f_3f_2+2f_3-2f_2$$

• Min 
$$((f_4-f_3)-2)^2$$

$$==> f_4^2 + f_3^2 + 4 - 2f_3f_4 - 4f_4 + 4f_3$$

• Min 
$$((f_5-f_4)-(-1))^2$$

$$Min ((f_5-f_4)-(-1))^2 ==> f_5^2+f_4^2+1-2f_5f_4+2f_5-2f_4$$

• Min 
$$((1-f_5)-(-1))^2$$

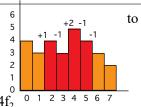
$$==> f_5^2 + 4 - 4f_5$$

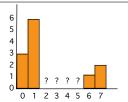




## 1D example: big quadratic

Copy





• Min  $(f_2^2+49-14f_2^2)$ 

$$+ f_3^2 + f_2^2 + 1 - 2f_3f_2 + 2f_3 - 2f_2$$

$$+ f_4^2 + f_3^2 + 4 - 2f_3f_4 - 4f_4 + 4f_3$$

$$+f_5^2+f_4^2+1-2f_5f_4+2f_5-2f_4$$

$$+ f_5^2 + 4 - 4f_5$$

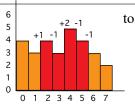
Denote it Q

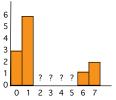




### 1D example: derivatives

Copy





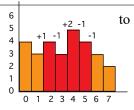
$$\begin{array}{l} \text{Min } \textbf{(f}_2\textbf{^2+49-14f}_2 & \frac{dQ}{df_2} = 2f_2 + 2f_2 - 2f_3 - 16 \\ & \textbf{+f}_3\textbf{^2+f}_2\textbf{^2+1-2f}_3\textbf{f}_2 + \textbf{2f}_3\textbf{-2f}_2 & \frac{dQ}{df_3} = 2f_3 - 2f_2 + 2 + 2f_3 - 2f_4 + 4 \\ & \textbf{+f}_4\textbf{^2+f}_3\textbf{^2+4-2f}_3\textbf{f}_4 - \textbf{4f}_4\textbf{+4f}_3 & \frac{dQ}{df_3} = 2f_3 - 2f_2 + 2 + 2f_3 - 2f_4 + 4 \\ & \textbf{+f}_5\textbf{^2+f}_4\textbf{^2+1-2f}_5\textbf{f}_4 + \textbf{2f}_5\textbf{-2f}_4 & \frac{dQ}{df_3} = 2f_4 - 2f_3 - 4 + 2f_4 - 2f_5 - 2 \\ \textbf{Denote it Q} & \frac{dQ}{df_5} = 2f_5 - 2f_4 + 2 + 2f_5 - 4 \end{array}$$

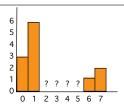




### 1D example: set derivatives to zero

#### Copy





$$\frac{dQ}{df_2} = 2f_2 + 2f_2 - 2f_3 - 16$$

$$\frac{dQ}{df_3} = 2f_3 - 2f_2 + 2 + 2f_3 - 2f_4 + 4$$

$$\frac{dQ}{df_4} = 2f_4 - 2f_3 - 4 + 2f_4 - 2f_5 - 2$$

$$\frac{dQ}{df_5} = 2f_5 - 2f_4 + 2 + 2f_5 - 4$$

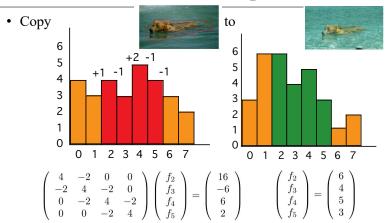
$$\begin{array}{c} \frac{dQ}{df_5} = 2f_5 - 2f_4 + 2 + 2f_5 - 4 \\ = > \begin{pmatrix} 4 & -2 & 0 & 0 \\ -2 & 4 & -2 & 0 \\ 0 & -2 & 4 & -2 \\ 0 & 0 & -2 & 4 \end{pmatrix} \begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 16 \\ -6 \\ 6 \\ 2 \end{pmatrix}$$

$$\begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 16 \\ -6 \\ 6 \\ 2 \end{pmatrix}$$





#### 1D example







### 1D example: remarks





$$\begin{pmatrix} 4 & -2 & 0 & 0 \\ -2 & 4 & -2 & 0 \\ 0 & -2 & 4 & -2 \\ 0 & 0 & -2 & 4 \end{pmatrix} \begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 16 \\ -6 \\ 6 \\ 2 \end{pmatrix}$$

- Matrix is sparse
- · Matrix is symmetric
- Everything is a multiple of 2
  - because square and derivative of square
- Matrix is a convolution (kernel -2 4 -2)
- Matrix is independent of gradient field. Only RHS is
- · Matrix is a second derivative





## Let's try to further analyze

• What is a simple case?





#### **Membrane interpolation**

- What if *v* is null?
- Laplace equation (a.k.a. membrane equation )

$$\min_{f} \iint_{\Omega} |\nabla f|^2 \text{ with } f|_{\partial\Omega} = f^*|_{\partial\Omega}$$

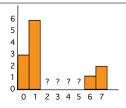






### 1D example: minimization

Minimize derivatives to interpolate



- Min  $(f_2-f_1)^2$
- Min  $(f_3-f_2)^2$
- Min  $(f_4-f_3)^2$
- Min  $(f_5-f_4)^2$
- Min  $(f_6-f_5)^2$

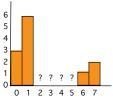
With f<sub>1</sub>=6 f<sub>6</sub>=1





### 1D example: derivatives

· Minimize derivatives to interpolate



$$\begin{array}{l} \text{Min } (f_2^2 + 36 - 12f_2 \\ \qquad \qquad + f_3^2 + f_2^2 - 2f_3f_2 \\ \qquad \qquad + f_4^2 + f_3^2 - 2f_3f_4 \\ \qquad \qquad + f_5^2 + f_4^2 - 2f_5f_4 \\ \qquad \qquad + f_5^2 + 1 - 2f_5) \\ \text{Denote it Q} \end{array}$$

$$\frac{dQ}{df_2} = 2f_2 + 2f_2 - 2f_3 - 12$$

$$\frac{dQ}{df_3} = 2f_3 - 2f_2 + 2f_3 - 2f_4$$

$$\frac{dQ}{df_4} = 2f_4 - 2f_3 + 2f_4 - 2f_5$$

$$\frac{dQ}{df_5} = 2f_5 - 2f_4 + 2f_5 - 2$$





### 1D example: set derivatives to zero

· Minimize derivatives to interpolate

$$\frac{dQ}{df_2} = 2f_2 + 2f_2 - 2f_3 - 12$$

$$\frac{dQ}{df_3} = 2f_3 - 2f_2 + 2f_3 - 2f_4$$

$$\frac{dQ}{df_4} = 2f_4 - 2f_3 + 2f_4 - 2f_5$$

$$\frac{dQ}{df_5} = 2f_5 - 2f_4 + 2f_5 - 2$$

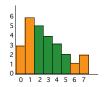
$$= > \begin{pmatrix}
4 & -2 & 0 & 0 \\
-2 & 4 & -2 & 0 \\
0 & -2 & 4 & -2 \\
0 & 0 & -2 & 4
\end{pmatrix}
\begin{pmatrix}
f_2 \\
f_3 \\
f_4 \\
f_5
\end{pmatrix} = \begin{pmatrix}
12 \\
0 \\
0 \\
2
\end{pmatrix}$$





#### 1D example

- Minimize derivatives to interpolate
- Pretty much says that second derivative should be zero  $(-1\ 2\ -1)$ is a second derivative filter



$$\begin{pmatrix} 4 & -2 & 0 & 0 \\ -2 & 4 & -2 & 0 \\ 0 & -2 & 4 & -2 \\ 0 & 0 & -2 & 4 \end{pmatrix} \begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 12 \\ 0 \\ 0 \\ 2 \end{pmatrix} \qquad \begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 5 \\ 4 \\ 3 \\ 2 \end{pmatrix}$$

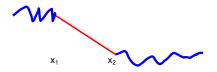
$$\begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 5 \\ 4 \\ 3 \\ 2 \end{pmatrix}$$





#### Intuition

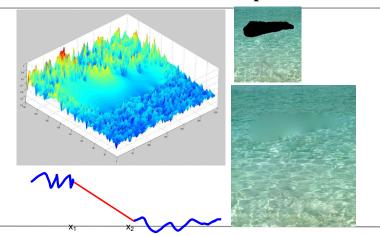
- In 1D; just linear interpolation!
  - The min of  $\int f$  is the slope integrated over the interval
- Locally, if the second derivative was not zero, this would mean that the first derivative is varying, which is bad since we want \( \int f \) to be minimized
- Note that, in 1D: by setting f', we leave two degrees of freedom. This is exactly what we need to control the boundary condition at x<sub>1</sub> and x<sub>2</sub>







### In 2D: membrane interpolation





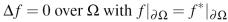
### Membrane interpolation

- What if v is null?
- Laplace equation (a.k.a. membrane equation )

$$\min_{f} \iint_{\Omega} |\nabla f|^2 \text{ with } f|_{\partial \Omega} = f^*|_{\partial \Omega}$$

 Mathematicians will tell you there is an Associated Euler-Lagrange equation:

$$\Delta f = 0$$
 over  $\Omega$  with  $f|_{\partial\Omega} = f^*|_{\partial\Omega}$ 



- Where the Laplacian  $\Delta$  is similar to -1 2 -1 in 1D
- Kind of the idea that we want a minimum, so we kind of derive and get a simpler equation







#### What is *v* is not null?





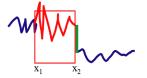


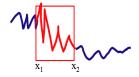
#### What if v is not null?

#### • 1D case

Seamlessly paste onto

Just add a linear function so that the boundary condition is respected









### (Review) Seamless Poisson cloning

• Given vector field *v* (pasted gradient), find the value of *f* in unknown region that optimize:

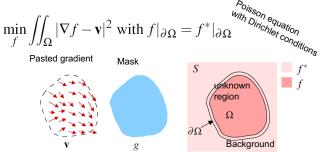


Figure 1: Guided interpolation notations. Unknown function f interpolates in domain  $\Omega$  the destination function  $f^*$ , under guidance of vector field  $\mathbf{v}$ , which might be or not the gradient field of a source function e.





#### What if v is not null: 2D

Variational minimization (integral of a functional) with boundary condition

$$\min_{f}\iint_{\Omega}|\nabla f-\mathbf{v}|^{2}\text{ with }f|_{\partial\Omega}=f^{*}|_{\partial\Omega},$$

• Euler-Lagrange equation:

$$\Delta f = {
m div} {f v} \ {
m over} \ \Omega, \ {
m with} \ f|_{\partial\Omega} = f^*|_{\partial\Omega}$$
 where  ${
m div} {f v} = rac{\partial u}{\partial x} + rac{\partial v}{\partial v}$  is the divergence of  ${f v} = (u,v)$ 

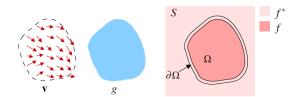




#### In 2D, if v is conservative

- If v is the gradient of an image g
- Correction function  $\hat{f}$  so that  $f = g + \hat{f}$
- $\widehat{f}$  performs membrane interpolation over $\Omega$ :

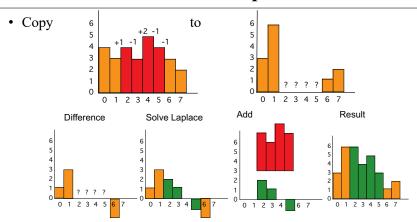
$$\Delta \tilde{f} = 0 \text{ over } \Omega, \ \tilde{f}|_{\partial \Omega} = (f^* - g)|_{\partial \Omega}$$







#### 1D example

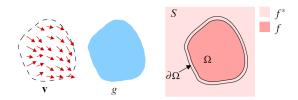






#### In 2D, if v is NOT conservative

- Also need to project the vector field v to a conservative field
- And do the membrane thing
- Of course, we do not need to worry about it, it's all handled naturally by the least square approach







#### Recap

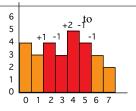
- Find image whose gradient best approximates the input gradient
  - least square Minimization
- Discrete case: turns into linear equation
  - Set derivatives to zero
  - Derivatives of quadratic ==> linear
- Continuous: turns into Euler-Lagrange form
  - $-\Delta f = div v$
- When gradient is null, membrane interpolation
  - Linear interpolation in 1D

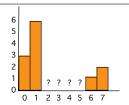




#### Discrete solver: Recall 1D

• Copy





$$\frac{dQ}{df_2} = 2f_2 + 2f_2 - 2f_3 - 16$$

$$\frac{dQ}{df_0} = 2f_3 - 2f_2 + 2 + 2f_3 - 2f_4 + 4$$

$$\frac{dQ}{df_4} = 2f_4 - 2f_3 - 4 + 2f_4 - 2f_5 - 2$$

$$\frac{dQ}{df_5} = 2f_5 - 2f_4 + 2 + 2f_5 - 4$$

$$= \begin{pmatrix} 4 & -2 & 0 & 0 \\ -2 & 4 & -2 & 0 \\ 0 & -2 & 4 & -2 \\ 0 & 0 & -2 & 4 \end{pmatrix} \begin{pmatrix} f_2 \\ f_3 \\ f_4 \\ f_5 \end{pmatrix} = \begin{pmatrix} 16 \\ -6 \\ 6 \\ 2 \end{pmatrix}$$





#### **Discrete Poisson solver**

- Two approaches:
  - Minimize variational problem
  - Solve Euler-Lagrange equation
     In practice, variational is best

$$\begin{split} & \min_{f} \iint_{\Omega} |\nabla f - \mathbf{v}|^2 \text{ with } f|_{\partial \Omega} = f^*|_{\partial \Omega}, \\ & \Delta f = \operatorname{div} \mathbf{v} \text{ over } \Omega, \text{ with } f|_{\partial \Omega} = f^*|_{\partial \Omega}. \end{split}$$

- · In both cases, need to discretize derivatives
  - Finite differences over 4 pixel neighbors
  - We are going to work using pairs
    - · Partial derivatives are easy on pairs
    - · Same for the discretization of v







### Discrete Poisson solver

Minimize variational problem  $\min_{f} \iint_{\Omega} |\nabla f - \mathbf{v}|^2 \operatorname{with} f|_{\partial\Omega} = f^*|_{\partial\Omega}$ Discretized gradient  $\min_{\substack{f\mid_{\Omega}\\ (\text{all pairs that})}} \sum_{\substack{(f_p-f_q-\nu_{pq})^2, \text{ with } f_p=f_p^*, \text{for all } p\in\partial\Omega}} \sup_{\substack{\text{Discretized}\\ \text{v: g(p)-g(q)}}} \sup_{\substack{\text{Boundary condition}}}$ are in  $\Omega$ )

Rearrange and call N<sub>p</sub> the neighbors of p

$$\text{Big yet sparse linear system} \begin{array}{c} \text{for all } p \in \Omega, \quad |N_p|f_p - \sum_{q \in N_p \cap \Omega} f_q = \sum_{q \in N_p \cap \partial} f_q^* + \sum_{q \in N_p} v_{pq} \\ \end{array}$$



Only for boundary pixels





# Result (eye candy)











cloning

seamless cloning



## Recap

- Find image whose gradient best approximates the input gradient
  - least square Minimization
- Discrete case: turns into big sparse linear equation
  - Set derivatives to zero
  - Derivatives of quadratic ==> linear





# Solving big matrix systems

- Ax=b
- You can use Matlab's \
  - (Gaussian elimination)
  - But not very scalable





### Iterative solvers

#### Important ideas

- Do not inverse matrix
- Maintain a vector x' that progresses towards the solution
- Updates mostly require to apply the matrix.
  - In many cases, it means you do no even need to store the matrix (e.g. for a convolution matrix you only need the kernel)
- Usually, you don't even wait until convergence
- Big questions: in which direction do you walk?
  - Yes, very similar to gradient descent





### Solving big matrix systems

- Ax=b, where A is sparse (many zero entries)
- In Pset 3, we ask you to use conjugate gradient
  - http://www.cs.cmu.edu/~quake-papers/painless-conjugate-gradient.pdf
  - $\ \underline{http://www.library.cornell.edu/nr/bookcpdf/c10-6.pdf}$

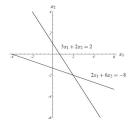




### Ax=b

- A is square, symmetric and positive-definite
- When A is dense, you're stuck, use backsubstitution
- When A is sparse, iterative techniques (such as Conjugate Gradient) are faster and more memory efficient
- Simple example:

$$\begin{bmatrix} 3 & 2 \\ 2 & 6 \end{bmatrix} x = \begin{bmatrix} 2 \\ -8 \end{bmatrix}$$





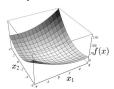


## Turn Ax=b into a minimization problem

- Minimization is more logical to analyze iteration (gradient ascent/descent)
- Quadratic form

$$f(x) = \frac{1}{2}x^T A x - b^T x + c$$

- c can be ignored because we want to minimize
- Intuition:
  - the solution of a linear system is always the intersection of n hyperplanes
  - Take the square distance to them
  - A needs to be positive-definite so that we have a nice parabola



Graph of quadratic form  $f(x) = \frac{1}{2}x^TAx - b^Tx + c$ . The minimum point of this surface is the solution to Ax = b.



Contours of the quadratic form. Each ellipsoidal curve has constant f(x).



# **Conjugate gradient**

- Smarter choice of direction
  - Ideally, step directions should be orthogonal to one another (no redundancy)
  - But tough to achieve
  - Next best thing: make them A-orthogonal (conjugate)
    That is, orthogonal when transformed by A:  $d_{(i)}^T A d_{(i)} = 0$

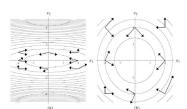


Figure 22: These pairs of vectors are A-orthogonal . . . because these pairs of vectors are orthogonal





### Recap

- Poisson image cloning: paste gradient, enforce boundary condition
- Variational formulation
- Also Euler-Lagrange formulation
- Discretize variational version, leads to big but sparse linear system
- $\min_{f} \iint_{\Omega} |\nabla f \mathbf{v}|^2 \text{ with } f|_{\partial \Omega} = f^*|_{\partial \Omega},$

$$\Delta f = \operatorname{div} \mathbf{v} \text{ over } \Omega, \text{ with } f|_{\partial\Omega} = f^*|_{\partial\Omega}$$

• Conjugate gradient is a smart iterative technique to solve it





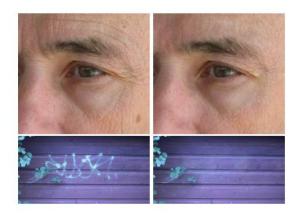
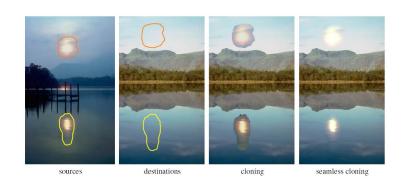


Figure 2: **Concealment**. By importing seamlessly a piece of the background, complete objects, parts of objects, and undesirable artifacts can easily be hidden. In both examples, multiple strokes (not shown) were used.







### Manipulate the gradient

• Mix gradients of g & f: take the max



Figure 8: **Inserting one object close to another**. With seamless cloning, an object in the destination image touching the selected region  $\Omega$  bleeds into it. Bleeding is inhibited by using mixed gradients as the guidance field.





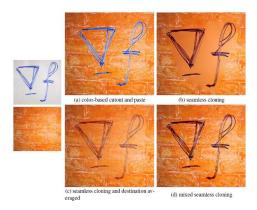


Figure 6: **Inserting objects with holes**. (a) The classic method, color-based selection and alpha masking might be time consuming and often leaves an undesirable halo; (b-c) seamless cloning, even averaged with the original image, is not effective; (d) mixed seamless cloning based on a loose selection proves effective.





swapped textures





Figure 7: **Inserting transparent objects**. Mixed seamless cloning facilitates the transfer of partly transparent objects, such as the rainbow in this example. The non-linear mixing of gradient fields picks out whichever of source or destination structure is the more salient at each location.



### **Reduce big gradients**

- Dynamic range compression
- See Fattal et al. 2002



Figure 10: Local illumination changes. Applying an appropriate non-linear transformation to the gradient field inside the selection and then integrating back with a Poisson solver, modifies locally the apparent illumination of an image. This is useful to highlight under-exposed foreground objects or to reduce specular reflections.





## **Seamless Image Stitching in the Gradient Domain**

- Anat Levin, Assaf Zomet, Shmuel Peleg, and Yair Weiss http://www.cs.huji.ac.il/~alevin/papers/eccv04-blending.pdf http://eprints.pascal-network.org/archive/00001062/01/tips05-blending.pdf
- Various strategies (optimal cut, feathering)

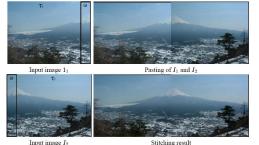


Fig. 1. Image stitching. On the left are the input images.  $\omega$  is the overlap region. On top right is a simple pasting of the input images. On the bottom right is the result of the GIST1 algorithm.





## **Poisson Matting**

- Sun et al. Siggraph 2004
- Assume gradient of F & B is negligible
- Plus various image-editing tools to refine matte

$$\begin{split} I &= \alpha F + (1 - \alpha)B \\ \nabla I &= (F - B)\nabla\alpha + \alpha\nabla F + (1 - \alpha)\nabla B \\ \nabla\alpha &\approx \frac{1}{F - B}\nabla I \end{split}$$



Figure 1: Pulling of matte from a complex scene. From left to right: a complex natural image for existing matting techniques where the color background is complex, a high quality matte generated by Poisson matting, a composite image with the extracted koala and a constant-color background, and a composite image with the extracted koala and a different background.





## Poisson-ish mesh editing

- http://portal.acm.org/citation.cfm?id=1057432.1057456
- http://www.cad.zju.edu.cn/home/xudong/Projects/mesh\_ec
- http://people.csail.mit.edu/sumner/research/deftransfer/



Figure 1: An unknown mythical creature. Left: mesh components for merging and deformation (the arm), Right: final editing result.

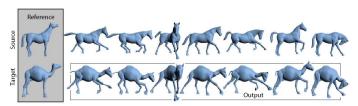


Figure 1: Deformation transfer copies the deformations exhibited by a source mesh onto a different target mesh. In this example, deformations of the reference horse mesh are transfered to the reference camel, generating seven new camel poses. Both gross skeletal changes as well as more subtle skin deformations are successfully reproduced.



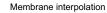


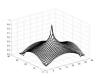
### Alternative to membrane

• Thin plate: minimize *second* derivative

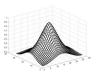
$$\min_f \int \int f_{xx}^2 + 2f_{xy}^2 + f_{yy}^2 \mathrm{d} \mathrm{x} \mathrm{d} \mathrm{y}$$







#### Thin-plate interpolation







# **Inpainting**

- More elaborate energy functional/PDEs
- <a href="http://www-mount.ee.umn.edu/~guille/inpainting.htm">http://www-mount.ee.umn.edu/~guille/inpainting.htm</a>







### **Key references**

- Socolinsky, D. <u>Dynamic Range Constraints in Image Fusion and Visualization</u> 2000. http://www.equinoxsensors.com/news.html
- Elder, Image editing in the contour domain, 2001 http://elderlab.yorku.ca/~elder/publications/journals/ElderPAMI01.pdf
- Fattal et al. 2002
   Gradient Domain HDR Compression <a href="http://www.cs.huji.ac.il/%7Edanix/hdr/">http://www.cs.huji.ac.il/%7Edanix/hdr/</a>
- Poisson Image Editing Perez et al. <a href="http://research.microsoft.com/vision/cambridge/papers/perez\_siggraph03.pdf">http://research.microsoft.com/vision/cambridge/papers/perez\_siggraph03.pdf</a>
- Covariant Derivatives and Vision, Todor Georgiev (Adobe Systems) ECCV 2006





A Gentle Introduction to Bilateral Filtering and its Applications

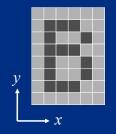


# Naïve Image Smoothing: Gaussian Blur

Sylvain Paris - MIT CSAIL

# **Notation and Definitions**

Image = 2D array of pixels



Pixel = intensity (scalar) or color (3D vector)

•  $I_{\mathbf{p}}$  = value of image I at position:  $\mathbf{p} = (p_x, p_y)$ 

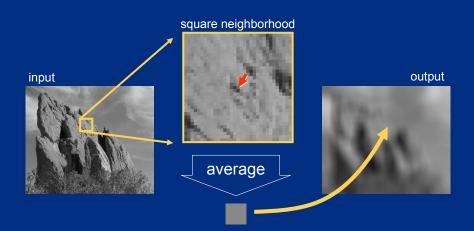
F[I] = output of filter F applied to image I



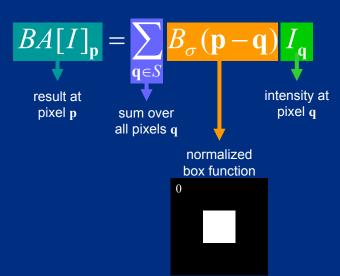
# Strategy for Smoothing Images

- Images are not smooth because adjacent pixels are different.
- Smoothing = making adjacent pixels look more similar.
- Smoothing strategy
   pixel → average of its neighbors

# **Box Average**



# **Equation of Box Average**



# **Square Box Generates Defects**

Axis-aligned streaks

Blocky results

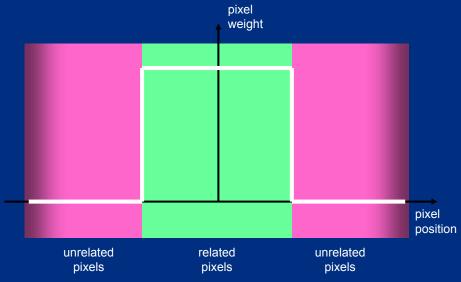
output







# **Box Profile**



# Strategy to Solve these Problems

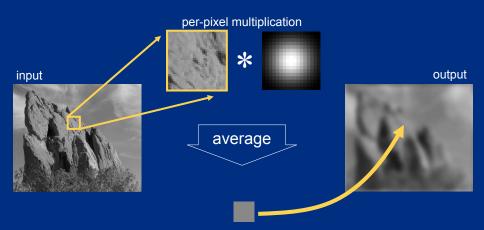
- Use an isotropic (i.e. circular) window.
- Use a window with a smooth falloff.





Gaussian window

# **Gaussian Blur**





# box average

# Gaussian blur

# **Equation of Gaussian Blur**

Same idea: weighted average of pixels.

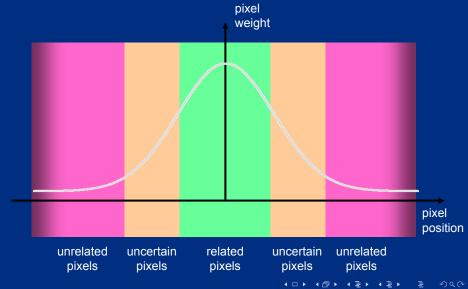
$$GB[I]_{\mathbf{p}} = \sum_{\mathbf{q} \in S} G_{\sigma}(\|\mathbf{p} - \mathbf{q}\|) I_{\mathbf{q}}$$

$$\text{normalized}$$

$$\text{Gaussian function}$$

#### **Gaussian Profile**

$$G_{\sigma}(x) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{x^2}{2\sigma^2}\right)$$



## **Spatial Parameter**

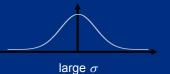


$$GB[I]_{\mathbf{p}} = \sum_{\mathbf{q} \in S} G_{\mathbf{q}}(\|\mathbf{p} - \mathbf{q}\|) I_{\mathbf{q}}$$

size of the window

input









strong smoothing

#### How to set $\sigma$

Depends on the application.

- Common strategy: proportional to image size
  - e.g. 2% of the image diagonal
  - property: independent of image resolution

## **Properties of Gaussian Blur**

Weights independent of spatial location

linear convolution

well-known operation

efficient computation (recursive algorithm, FFT )

- Does smooth images
- But smoothes too much: edges are blurred.
  - Only spatial distance matters
  - No edge term

$$GB[I]_{\mathbf{p}} = \sum_{\mathbf{q} \in S} G_{\sigma} (||\mathbf{p} - \mathbf{q}||) I_{\mathbf{q}}$$
space







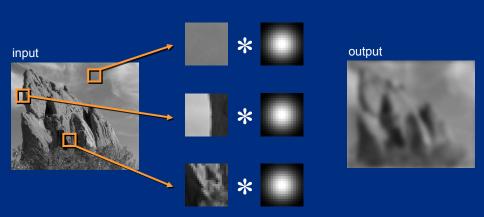
A Gentle Introduction to Bilateral Filtering and its Applications



# "Fixing the Gaussian Blur": the Bilateral Filter

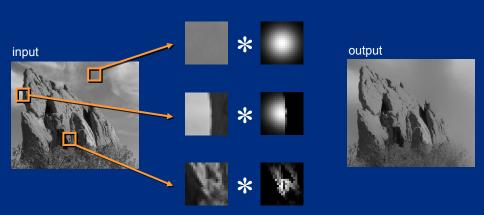
Sylvain Paris - MIT CSAIL

# Blur Comes from Averaging across Edges



Same Gaussian kernel everywhere.

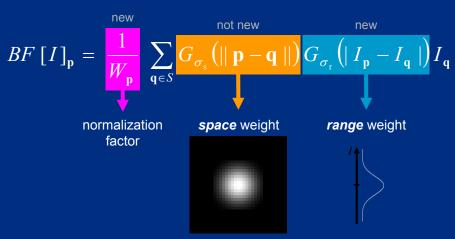
# Bilateral Filter [Aurich 95, Smith 97, Tomasi 98] No Averaging across Edges



The kernel shape depends on the image content.

# Bilateral Filter Definition: an Additional Edge Term

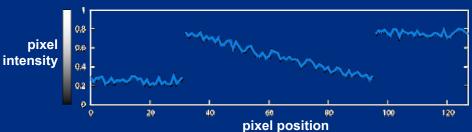
Same idea: weighted average of pixels.



#### Illustration a 1D Image

1D image = line of pixels

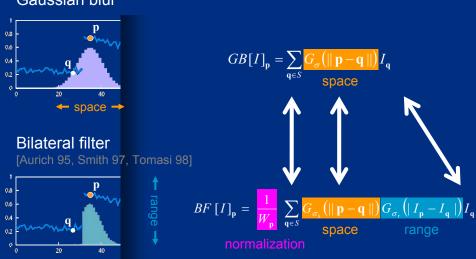




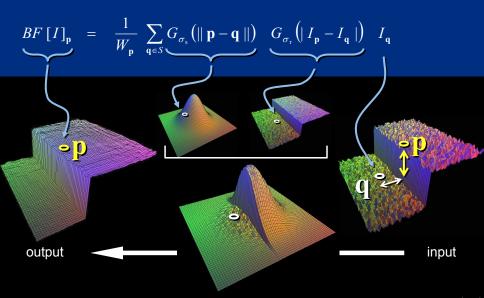
#### Gaussian Blur and Bilateral Filter

#### Gaussian blur

space



## Bilateral Filter on a Height Field



## **Space and Range Parameters**

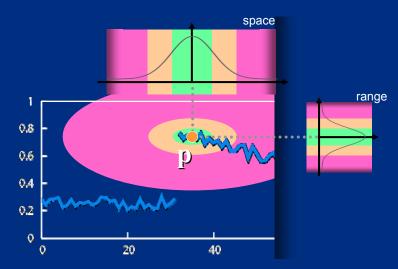
$$BF\left[I\right]_{\mathbf{p}} = \frac{1}{W_{\mathbf{p}}} \sum_{\mathbf{q} \in S} G_{\sigma_{s}} (||\mathbf{p} - \mathbf{q}||) G_{\sigma_{r}} (|I_{\mathbf{p}} - I_{\mathbf{q}}|) I_{\mathbf{q}}$$

• space  $\sigma_s$ : spatial extent of the kernel, size of the considered neighborhood.

• range  $\sigma_{\rm r}$ : "minimum" amplitude of an edge

#### Influence of Pixels

Only pixels close in space and in range are considered.





#### **Exploring the Parameter Space**

$$\sigma_{\rm r}$$
 = 0.1

$$\sigma_{\rm r}$$
 = 0.25

$$\sigma_{\rm r}^{\,=\,\infty}$$
 (Gaussian blur)

input 
$$\sigma_{
m s}\!=\!2$$





$$\sigma_{\rm s} = 6$$







 $\sigma_{\rm s}$  = 18









input

$$\sigma_{\rm r} = 0.1$$

$$\sigma_{\rm r} = 0.25$$

$$\sigma_{\rm r} = \infty$$
 (Gaussian blur)

$$\sigma_{\rm r} = 0.25$$

 $\sigma_{\rm s}$  = 6

 $\sigma_{\rm s} = 2$ 







 $\sigma_{\rm s} = 18$ 













$$\sigma_{\rm r} = \infty$$
 (Gaussian blur)



input

$$\sigma_{\rm r} = 0.1$$

$$\sigma_{\rm r} = 0.25$$

 $\sigma_{\rm r}^{\,=\,\infty}$  (Gaussian blur)







 $\sigma_{\rm s} = 2$ 













$$\sigma_{\rm s}$$
 = 18







**4** 🗗 ▶











#### **How to Set the Parameters**

Depends on the application. For instance:

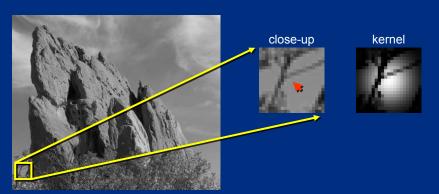
- space parameter: proportional to image size
  - e.g., 2% of image diagonal
- range parameter: proportional to edge amplitude
  - e.g., mean or median of image gradients
- independent of resolution and exposure



# A Few More Advanced Remarks

#### **Bilateral Filter Crosses Thin Lines**

- Bilateral filter averages across features thinner than ~2σ<sub>s</sub>
- Desirable for smoothing: more pixels = more robust
- Different from diffusion that stops at thin lines



#### **Iterating the Bilateral Filter**

$$I_{(n+1)} = BF\left[I_{(n)}\right]$$

- Generate more piecewise-flat images
- Often not needed in computational photo.









#### Bilateral Filtering Color Images

For gray-level images

$$BF\left[I\right]_{\mathbf{p}} = \frac{1}{W_{\mathbf{p}}} \sum_{\mathbf{q} \in S} G_{\sigma_{s}} \left( \| \mathbf{p} - \mathbf{q} \| \right) G_{\sigma_{r}} \left( \frac{I_{\mathbf{p}} - I_{\mathbf{q}}}{I_{\mathbf{p}}} \right) \frac{I_{\mathbf{q}}}{scala}$$



For color images

For color images 
$$BF\left[I\right]_{\mathbf{p}} = \frac{1}{W_{\mathbf{p}}} \sum_{\mathbf{q} \in S} G_{\sigma_{s}} \left( \| \mathbf{p} - \mathbf{q} \| \right) G_{\sigma_{r}} \left( \| \mathbf{C}_{\mathbf{p}} - \mathbf{C}_{\mathbf{q}} \| \right) \mathbf{C}_{\mathbf{q}}$$
 3D vector



The bilateral filter is extremely easy to adapt to your need.

#### **Hard to Compute**

- Nonlinear  $BF[I]_{\mathbf{p}} = \frac{1}{W_{\mathbf{p}}} \sum_{\mathbf{q} \in S} G_{\sigma_s} (||\mathbf{p} \mathbf{q}||) \frac{G_{\sigma_r} (|I_{\mathbf{p}} I_{\mathbf{q}}|)}{I_{\mathbf{q}}}$
- Complex, spatially varying kernels
  - Cannot be precomputed, no FFT









Brute-force implementation is slow > 10min

# **Questions?**



### **Image Smoothing**



A fundamentally important tool





[Kass and Solomon 10]

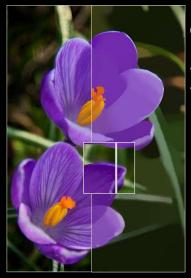


[Fattal et al. 06]



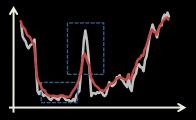
[Farbman et al. 08]

### Image Smoothing



#### General goals:

- Suppress insignificant details
- · Maintain major edges

















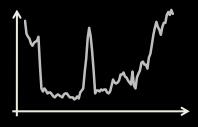




#### **Our New Smoothing Method**

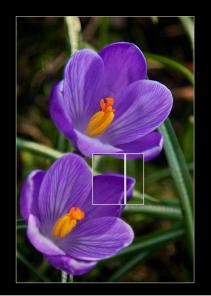
A general and effective global smoothing strategy based on a sparsity measure

$$c(f) := \#\{p \mid \left| \nabla f_p \right| \neq 0\}$$



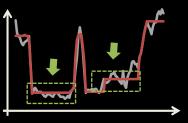
which corresponds to the LO-norm of gradient

#### **Two Features**

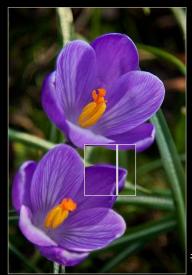


#### 1. Flattening insignificant details

By removing small non-zero gradients

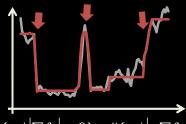


#### **Two Features**



#### 2. Enhancing prominent edges

Because large gradients receive the same penalty as small ones



$$\|\#\{p \mid \left| \nabla f_p \right| \neq 0\} = \#\{p \mid \left| \alpha \nabla f_p \right| \neq 0\}$$

#### D

#### Our Framework in 1D

• Constrain # of non-zero gradients

$$c(f) = \#\{p \mid |f_p - f_{p+1}| \neq 0\} = k$$

• Make the result similar to the input g

$$\min_{f} \sum_{p} (f_p - g_p)^2$$

Objective function

$$\min_{f} \sum (f_p - g_p)^2 \quad \text{s.t.} \quad c(f) = k$$



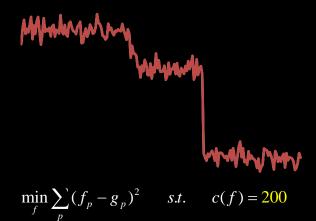
$$\min_{f} \sum_{p} (f_p - g_p)^2 \quad \text{s.t.} \quad c(f) = 1$$



$$\min_{f} \sum_{p} (f_p - g_p)^2 \qquad s.t. \qquad c(f) = 2$$



$$\min_{f} \sum_{p} (f_p - g_p)^2 \qquad s.t. \qquad c(f) = 5$$

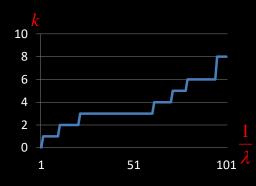


### **Transformation**

$$\min_{f} \sum_{p} (f_p - g_p)^2 \quad \text{s.t.} \quad c(f) = k$$



$$\min_{f} \sum_{p} (f_{p} - g_{p})^{2} + \lambda \cdot c(f)$$



#### 2D Image

$$\min_{f} \sum_{p} (f_{p} - g_{p})^{2} + \lambda \cdot c(\partial_{x} f, \partial_{y} f)$$

$$c(\partial_{x} f, \partial_{y} f) = \#\{p \mid |\partial_{x} f_{p}| + |\partial_{y} f_{p}| \neq 0\}$$

Finding the global optimum is NP hard

#### **Approximation**

$$\min_{f} \sum_{p} (f_{p} - g_{p})^{2} + \lambda \cdot c(\boldsymbol{h}, \boldsymbol{v})$$

$$+ \beta \cdot \sum_{p} \left( (\partial_{x} f_{p} - h_{p})^{2} + (\partial_{y} f_{p} - v_{p})^{2} \right)$$

Separately estimate f and (h, v)

#### **Iterative Optimization**

• Compute f given h, v

$$E(f) = \sum (f_p - g_p)^2 + \beta \cdot ((\partial_x f_p - h_p)^2 + (\partial_y f_p - v_p)^2)$$

Both the sub-problems are with closed-form solutions

$$E(h, v) = \sum_{p} \left( \left( \partial_{x} f_{p} - h_{p} \right)^{2} + \left( \partial_{y} f_{p} - v_{p} \right)^{2} \right) + \frac{\kappa}{\beta} c(h, v)$$

Gradually approximate the original problem

### One Example



Iteration #89



Input



 $\lambda = 0.01$ 



 $\lambda = 0.02$ 



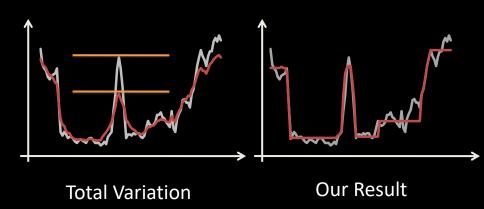
 $\lambda = 0.03$ 

# Comparison

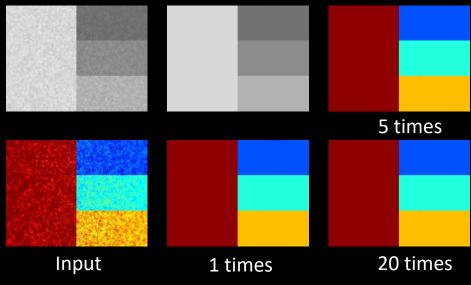


Total BABIes saut to n

### Comparison



# **Another Example**



# **Applications**





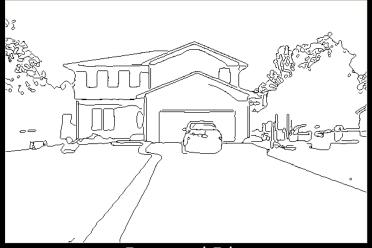
**Gradient Map** 



**Extracted Edge** 



Smoothing result



**Extracted Edge** 









Without smoothing

With smoothing

### **Clip-Art JPEG Artifact Removal**









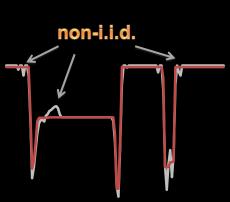




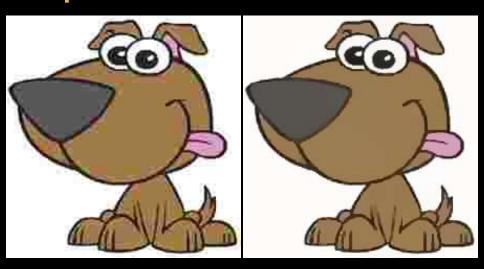


# **Clip-Art JPEG Artifact Removal**

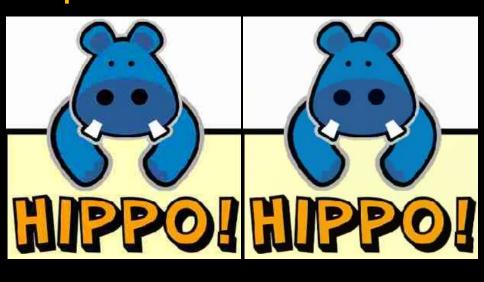




#### **Clip-Art JPEG Artifact Removal**



## **Clip-Art JPEG Artifact Removal**



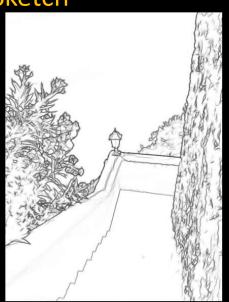
# Image Abstraction



# Image Abstraction



## Pencil Sketch



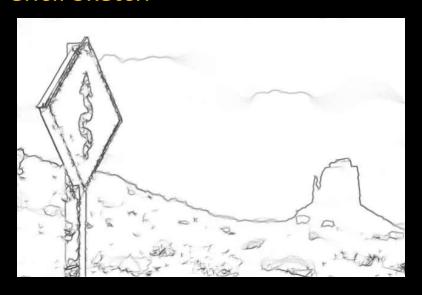
# Image Abstraction

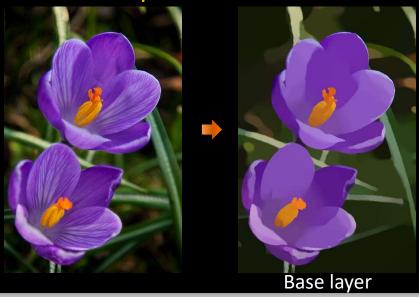


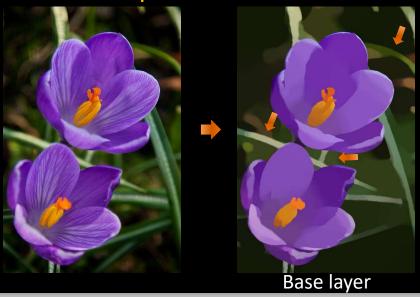
## Image Abstraction



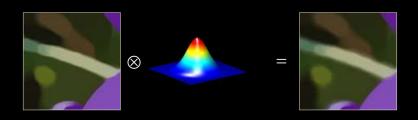
## Pencil Sketch







#### Edge Adjustment



Spatially varying Gaussian blur in an optimization procedure

## Edge Adjustment

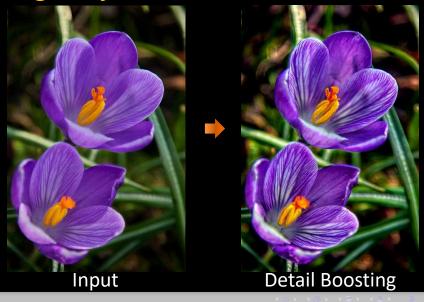
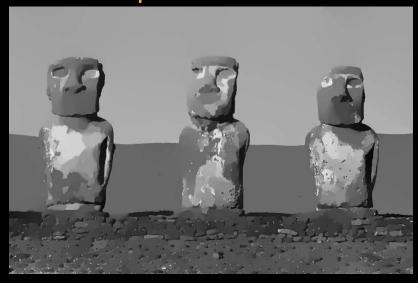




Image of [Farbman et al. 08]







HDR Input (gamma adjusted)



Log-base layer to be compressed



Detail layer





HDR Input (gamma adjusted)









Strong texture will be preserved



Our smoothing result



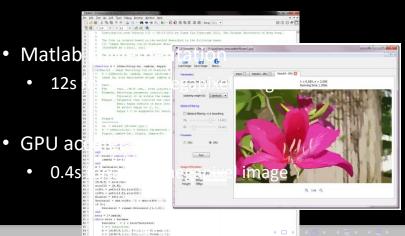
**Bilateral Filter** 



Bilateral Filter + Ours

#### **Implementation**

 Matlab source code and Windows software are available



#### Conclusion

- A simple and general smoothing framework
  - Approximate L0-norm gradient measure
  - Flatten low-amplitude details
  - Enhance prominent structures
- Possible extensions in graphics and vision
  - Video
  - 3D surface (modeling)
  - Depth

#### We wish to thank

- Michael S. Brown for narrating the video.
- The *anonymous reviewers* for constructive comments.
- Flicker Users: John McCormick, conner395, cyber-seb, T-KONI, Remi Longva, dms\_a\_jem for allowing us to use their pictures.



The End



张举勇 中国科学技术大学

#### Overview

- Image stitching is to combine multiple photos to create a larger photo.
- This technology is now widely available. It's on pretty much all smart phones that are in market today.
- It's also used in other domains, such as medical imaging, and remote sensing.











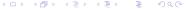
Image 1

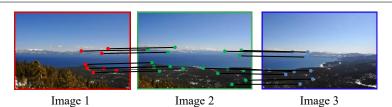
Image 2

Image 3

How would you align these images?



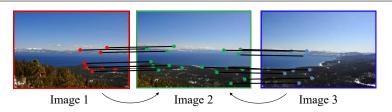




Find corresponding points (using feature detectors like SIFT)





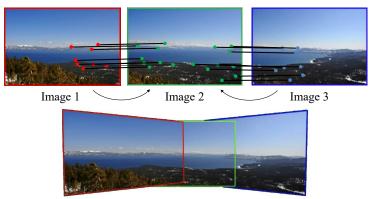


Find geometric relationship between the images





# **Image Stitching**



Warp images so that corresponding points align





## Image Stitching





Blend images to remove hard seams





## Image Stitching

### Topics:

- 2x2 Image Transformations
- 3x3 Image Transformations
- Computing Homography
- Dealing with Outliers:RANSAC
- Warping and Blending Images





## 2x2 Image Transformations

Image Filtering: Change range (brightness)

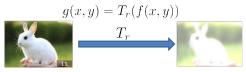
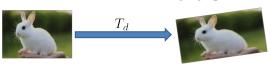


Image Warping: Change domain (location)

$$g(x,y) = f(T_d(x,y))$$

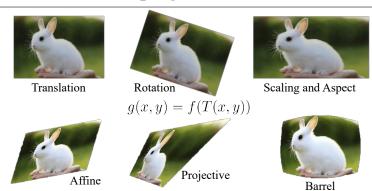
Transformation  $T_d$  is a coordinate changing operator







### Global Warping/Transformation

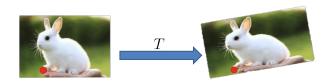


Transformation T is the same over entire domain often can be described by just a few parameters





### 2x2 Linear Transformations



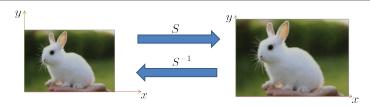
T can be represented by a matrix.

$$\mathbf{p}_2 = T\mathbf{p}_1 \qquad \begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = T \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} \qquad \begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = \begin{bmatrix} a_{11} & a_{12} \\ a_{21} & a_{22} \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$





## Scaling(Stretching and Squishing)



#### Forward:

$$x_2 = ax_1 \qquad y_2 = by_1$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = S \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} a & 0 \\ 0 & b \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$

#### Inverse:

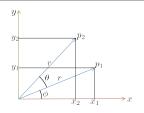
$$x_1 = \frac{1}{a}x_2 \qquad y_1 = \frac{1}{b}y_2$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = S \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} a & 0 \\ 0 & b \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} \qquad \qquad \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = S^{-1} \begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = \begin{bmatrix} 1/a & 0 \\ 0 & 1/b \end{bmatrix} \begin{bmatrix} x_2 \\ y_2 \end{bmatrix}$$





### 2D Rotation



$$x_1 = r\cos(\phi)$$

$$y_1 = r\sin(\phi)$$

$$x_2 = r\cos(\phi + \theta)$$

$$y_2 = r\sin(\phi + \theta)$$

$$x_2 = r\cos(\phi)\cos(\theta) - r\sin(\phi)\sin(\theta)$$

$$y_2 = r\cos(\phi)\sin(\theta) + r\sin(\phi)\cos(\theta)$$

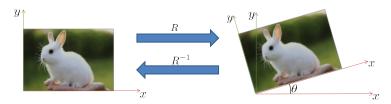
$$x_2 = x_1 \cos(\theta) - y_1 \sin(\theta)$$

$$y_2 = x_1 \sin(\theta) + y_1 \cos(\theta)$$





### Rotation



#### Forward:

$$x_2 = x_1 \cos(\theta) - y_1 \sin(\theta)$$
$$y_2 = x_1 \sin(\theta) + y_1 \cos(\theta)$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = R \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} \cos(\theta) & -\sin(\theta) \\ \sin(\theta) & \cos(\theta) \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$

#### Inverse:

$$x_1 = x_2 \cos(\theta) + y_2 \sin(\theta)$$

$$y_1 = -x_2\sin(\theta) + y_2\cos(\theta)$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = R \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} \cos(\theta) & -\sin(\theta) \\ \sin(\theta) & \cos(\theta) \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} \qquad \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = R^{-1} \begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = \begin{bmatrix} \cos(\theta) & \sin(\theta) \\ -\sin(\theta) & \cos(\theta) \end{bmatrix} \begin{bmatrix} x_2 \\ y_2 \end{bmatrix}$$





### Skew

*y*↑

#### Horizontal Skew:

$$x_2 = x_1 + m_x y_1$$
$$y_2 = y_1$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = S_x \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} 1 & m_x \\ 0 & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$



#### Vertical Skew:

$$x_2 = x_1$$
$$y_2 = m_y x_1 + y_1$$

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = S_y \begin{bmatrix} x_1 \\ y_1 \end{bmatrix} = \begin{bmatrix} 1 & 0 \\ m_y & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$





### Mirrow

 $\hat{x}$ 



Mirrow about Y-axis:

$$x_2 = -x_1$$
$$y_2 = y_1$$

$$M_y = \begin{bmatrix} -1 & 0 \\ 0 & 1 \end{bmatrix}$$



Mirrow about line y = x:

$$x_2 = y_1$$

$$y_2 = x_1$$

$$M_{xy} = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}$$





### 2x2 Matrix Transformations

Any transformation of the form:

$$\begin{bmatrix} x_2 \\ y_2 \end{bmatrix} = \begin{bmatrix} a_{11} & a_{12} \\ a_{21} & a_{22} \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \end{bmatrix}$$

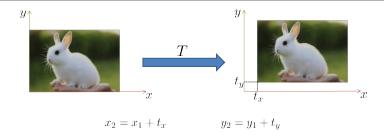
- Origin maps to the origin
- Lines map to lines
- Parallel lines remain parallel
- Closed under composition

$$\left. \begin{array}{l} \mathbf{p}_2 = T_{21} \mathbf{p}_1 \\ \mathbf{p}_3 = T_{32} \mathbf{p}_2 \\ \mathbf{p}_3 = T_{31} \mathbf{p}_1 \end{array} \right\} \mathbf{p}_3 = T_{32} \mathbf{p}_2 = T_{32} T_{21} \mathbf{p}_1 \Rightarrow T_{31} = T_{32} T_{21}$$





### **Translation**



Can translation be expressed as a 2x2 matrix? No





## Homogenous Coordinates

The homogenous representation of a 2D point  $\mathbf{p} = (x, y)$  is a 3D point  $\tilde{\mathbf{p}} = (\tilde{x}, \tilde{y}, \tilde{z})$ . The third coordinate  $\tilde{z} \neq 0$  is fictitious such that:

$$x = \frac{\tilde{x}}{\tilde{z}} \qquad y = \frac{\tilde{y}}{\tilde{z}}$$

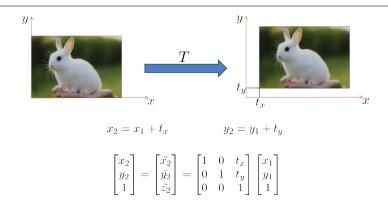
$$\mathbf{p} = \begin{bmatrix} x \\ y \\ 1 \end{bmatrix} = \begin{bmatrix} \tilde{z}x \\ \tilde{z}y \\ \tilde{z} \end{bmatrix} = \begin{bmatrix} \tilde{x} \\ \tilde{y} \\ \tilde{z} \end{bmatrix} = \tilde{\mathbf{p}}$$

Every point on line **L**(except origin) represent the homogenous coordinate of  $\mathbf{p}(x, y)$ 





### **Translation**







## Scaling, Rotation, Skew, Translation

$$\begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} 1 & m_x & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \\ 1 \end{bmatrix}$$

Skew

$$\begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} 1 & 0 & t_x \\ 0 & 1 & t_y \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \\ 1 \end{bmatrix}$$

$$\begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} 1 & 0 & t_x \\ 0 & 1 & t_y \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \\ 1 \end{bmatrix} \qquad \qquad \begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} \cos \theta & -\sin \theta & 0 \\ \sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_1 \\ y_1 \\ 1 \end{bmatrix}$$

Translation

**Rotation** 

Composition of these transformations?





### Affine Transformation

### Any transformation of the form:

$$\begin{bmatrix} x_2 \\ y_2 \\ 1 \end{bmatrix} = \begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \tilde{x_1} \\ \tilde{y_1} \\ \tilde{z_1} \end{bmatrix}$$











### **Affine Transformation**

### Any transformation of the form:

$$\begin{bmatrix} x_2 \\ y_2 \\ 1 \end{bmatrix} = \begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \tilde{x_1} \\ \tilde{y_1} \\ \tilde{z_1} \end{bmatrix}$$

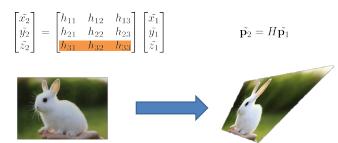
- Origin does not necessarily map to the origin
- · Lines map to lines
- Parallel lines remain parallel
- Closed under composition





## **Projective Transformation**

### Any transformation of the form:



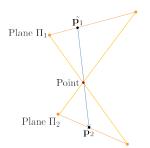
Also called Homography





## **Projective Transformation**

### Mapping of one plane to another through a point



$$\tilde{\mathbf{p}_2} = H\tilde{\mathbf{p}_1}$$

$$\begin{bmatrix} \tilde{x_2} \\ \tilde{y_2} \\ \tilde{z_2} \end{bmatrix} = \begin{bmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{bmatrix} \begin{bmatrix} \tilde{x_1} \\ \tilde{y_1} \\ \tilde{z_1} \end{bmatrix}$$

Same as imaging a plane through a pinhole





## **Projective Transformation**

Homography can only be defined up to a scale.

$$\begin{bmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{bmatrix} \begin{bmatrix} \vec{x_1} \\ \vec{y_1} \\ \vec{z_1} \end{bmatrix} = \begin{bmatrix} \vec{x_2} \\ \vec{y_2} \\ \vec{z_2} \end{bmatrix} = k \begin{bmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{bmatrix} \begin{bmatrix} \vec{x_1} \\ \vec{y_1} \\ \vec{z_1} \end{bmatrix}$$

If we fix scale such that  $\sqrt{\sum (h_{ij})^2}$  then 8 free parameters

- · Origin does not necessarily map to the origin
- Lines map to lines
- Parallel lines do not necessarily remain parallel
- Closed under composition



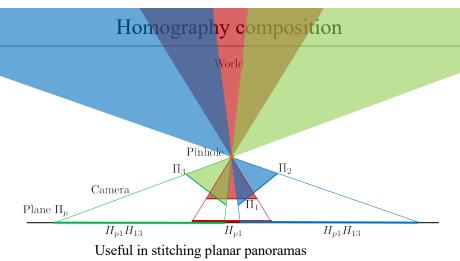


# Remember Vanishing Points?













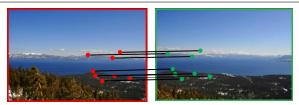


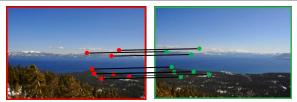
Image 1

Image 2

Given a set of matching features/points between image images 1 and 2, find the homography *H* that best "agrees" with the matches.







Source Image

Destination Image

$$\begin{bmatrix} x_d \\ y_d \\ 1 \end{bmatrix} = \begin{bmatrix} \tilde{x_d} \\ \tilde{y_d} \\ \tilde{z_d} \end{bmatrix} = \begin{bmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{bmatrix} \begin{bmatrix} x_s \\ y_s \\ 1 \end{bmatrix}$$

How many unknows? 9 ...But 8 degrees of freedom How many minimum pairs of matching points? 4





### For a given pair *i* of corresponding points:



$$x_d^{(i)} = \frac{\tilde{x}_d^{(i)}}{\tilde{z}_d^{(i)}} = \frac{h_{11}x_s^{(i)} + h_{12}y_s^{(i)} + h_{13}}{h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}}$$
$$y_d^{(i)} = \frac{\tilde{y}_d^{(i)}}{\tilde{z}_d^{(i)}} = \frac{h_{21}x_s^{(i)} + h_{22}y_s^{(i)} + h_{23}}{h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}}$$

#### Rearranging the terms:



$$\begin{aligned} x_d^{(i)}(h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}) &= h_{11}x_s^{(i)} + h_{12}y_s^{(i)} + h_{13} \\ y_d^{(i)}(h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}) &= h_{21}x_s^{(i)} + h_{22}y_s^{(i)} + h_{23} \end{aligned}$$





$$\begin{aligned} x_d^{(i)}(h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}) &= h_{11}x_s^{(i)} + h_{12}y_s^{(i)} + h_{13} \\ y_d^{(i)}(h_{31}x_s^{(i)} + h_{32}y_s^{(i)} + h_{33}) &= h_{21}x_s^{(i)} + h_{22}y_s^{(i)} + h_{23} \end{aligned}$$

Rearranging the terms and writing as linear equation:







### Combining the equation for all corresponding points:

Solve for  $\mathbf{h}$ :  $\mathbf{A}\mathbf{h} = 0$  such that  $\|\mathbf{h}\|^2 = 1$ 





## **Constrained Least Squares**

Solve for  $\mathbf{h}$ :  $\mathbf{A}\mathbf{h} = 0$  such that  $\|\mathbf{h}\|^2 = 1$ 

### Define least squares problem:

$$\min_{\mathbf{h}} \|\mathbf{A}\mathbf{h}\|^2 \text{ such that } \|\mathbf{h}\|^2 = 1$$

We know that:

$$\|\mathbf{A}\mathbf{h}\|^2 = (\mathbf{A}\mathbf{h})^T \mathbf{A}\mathbf{h} = \mathbf{h}^T \mathbf{A}^T \mathbf{A}\mathbf{h}$$
 and  $\|\mathbf{h}\|^2 = \mathbf{h}^T \mathbf{h} = 1$ 

$$\min_{\mathbf{h}}(\mathbf{h}^T \boldsymbol{A}^T \boldsymbol{A} \mathbf{h}) \text{ such that } \|\mathbf{h}\|^2 = 1$$





### Constrained Least Squares

$$\min_{\mathbf{h}}(\mathbf{h}^T \mathbf{A}^T \mathbf{A} \mathbf{h})$$
 such that  $\|\mathbf{h}\|^2 = 1$ 

Define Loss function  $L(\mathbf{h}, \lambda)$ :

$$L(\mathbf{h}, \lambda) = \mathbf{h}^T \mathbf{A}^T \mathbf{A} \mathbf{h} - \lambda (\mathbf{h}^T \mathbf{h} - 1)$$

Taking derivatives of  $L(\mathbf{h}, \lambda)$  w.r.t  $\mathbf{h}$ :  $2\mathbf{A}^T \mathbf{A} \mathbf{h} - 2\lambda \mathbf{h} = 0$ 

$$A^T A h = \lambda h$$
 Eigenvalue Problem

Eigenvector **h** with smallest eigenvalue  $\lambda$  of matrix  $A^T A$  minimizes the loss function  $L(\mathbf{h})$ .

Matlab:  $eig(A^*A)$  returns eigenvalues and vectors of  $A^TA$ 





# What could go wrong?

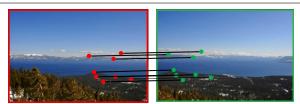


Image 1 Image 2





## What could go wrong?

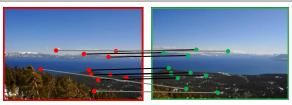


Image 1

Image 2

#### Outliers!

We need to robustly compute transformation in the presence of wrong matches.

If number of outliers < 50%, then RANSAC to the rescue!





### RANdom SAmple Consensus

### General RANSAC algorithm:

- 1. Randomly choose *s* samples. Typically *s* is the minimum samples to fit the model.
- 2. Fit the model to the randomly chosen samples.
- 3. Count the number M of data points(inliers) that fit the model within a measure of error  $\varepsilon$ .
- 4. Repeat Steps 1-3 N times
- 5. Choose the model that has the largest number M of inliers.

### For homography:

s = 4 points.  $\varepsilon$  is acceptable alignment error in pixels.



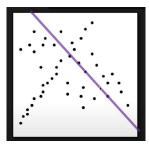


# RANSAC Example: Line Fitting

### Robust line fitting:



Least Squares Fitting Inliers:2



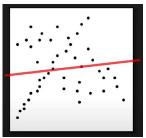
RANSAC Interation 1 Inliers:4



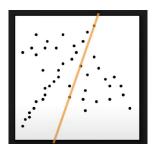


# RANSAC Example: Line Fitting

### Robust line fitting:



Least Squares Fitting Inliers:2



RANSAC Interation 2 Inliers:3



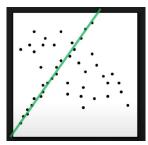


# RANSAC Example: Line Fitting

### Robust line fitting:



Least Squares Fitting Inliers:2



RANSAC Interation i Inliers:20





# Warping Images

Given a transformation T and a image f(x,y), compute the transformed image g(x,y)

$$g(x,y) = f(T(x,y))$$

$$T(x,y)$$

$$f(x,y)$$

$$g(x,y)$$





# Forward Warping

Send each pixel (x,y) in f(x,y) to its corresponding location T(x,y) in g(x,y)

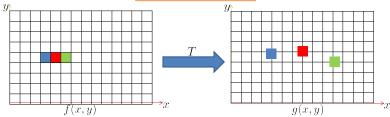




# Forward Warping

Send each pixel (x,y) in f(x,y) to its corresponding location T(x,y) in g(x,y)

$$g(x,y) = f(T(x,y))$$



What if pixel lands in between pixels? What if not all pixels in g(x,y) are filled?

Can result in holes!

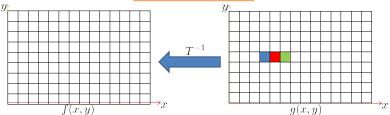




# **Backward Warping**

Get each pixel (x,y) in g(x,y) from its corresponding location  $T^{-1}(x,y)$  in f(x,y)

$$g(x,y) = f(T(x,y))$$

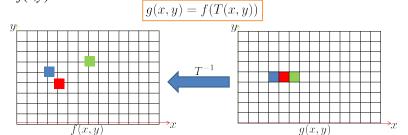






# **Backward Warping**

Get each pixel (x,y) in g(x,y) from its corresponding location  $T^{-1}(x,y)$  in f(x,y)



What if pixel lands between pixels? Use Nearest Neighbor or Interpolate







Image 1



Image 2



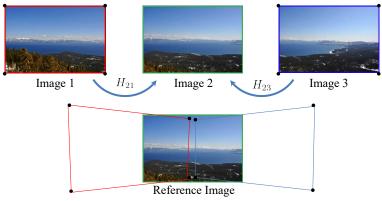
Image 3



Reference Image (Image 2)



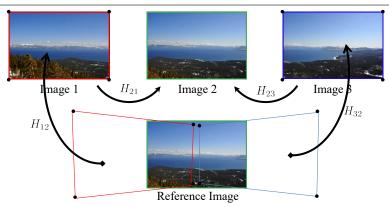




Compute the bounds of Image 1 and Image 3 in reference image space



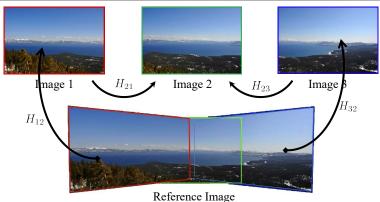




For each pixel within bounds, compute its location in captured image







For each pixel within bounds, compute its location in captured image





# Blending Images



Overlaid Aligned Images

Hard seams due to vignetting, exposure differences, etc.





# Blending Images: Averaging



Averaged Images

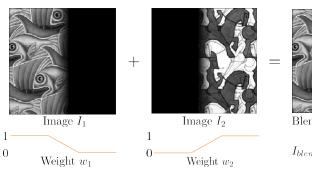
Seams still visible





# Blending Images

Say we want to blend images  $I_1$  and  $I_2$  at the center



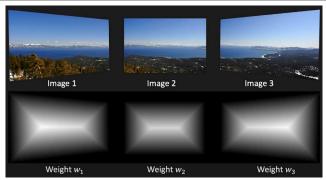


Blended Image  $I_{blend}$   $I_{blend} = \frac{w_1 I_1 + w_2 I_2}{w_1 + w_2}$ 





# **Computing Weighting Functions**



Pixels closer to the edge get a lower weight.

Ex: Distance Transform (bwdist in MATLAB)





# Weighted Blending





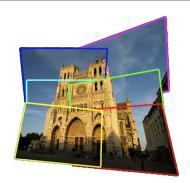




# Image Stitching Example



Source Images



Aligned Images





# Image Stitching Example



Source Images



Blended Images







# **Face Detection**

张举勇 中国科学技术大学

## What is Face Detection?

#### Locate human faces in images







# Image Stitching

Locate human faces in images.

#### Topics:

- Uses of Face Detection
- Haar Features for Face Detection
- Integral Image
- Nearest Neighbor Classifier
- Support Vector Machine







Face Detection

Finding People using Search Engines







Only faces of people named "Gates"

Finding People using Search Engines







**Intelligent Marketing** 







Biometrics, Surveillance, Monitoring





# Face Detection in Computers

Slide windows of different sizes across image. At each location match window to face model.







#### Face Detection Framework

#### For each window:



Extract

 $\stackrel{\text{Features}}{\longrightarrow} \lceil f \rceil \stackrel{\text{Model}}{\longrightarrow} \hspace{0.1cm} \text{Yes / No}$ 

Match Face

Features: Which features represent faces well?

Classifier: How to construct a face model and efficiently

classify features as face or not?





#### What are Good Features?

Interest Points (Edges, Corners, SIFT)?







Facial Components (Templates)?











### Charateristics of Good Features

#### Discriminate Face/Non-Face



Extremely Fast to Compute

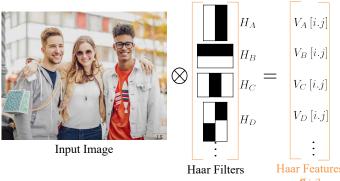
Need to evaluate millions of windows in an image





#### Haar Features

#### Set of Correlation Responses to Haar Filters









# Discriminative Ability of Haar Feature



 $V_A = 64$ 



 $V_A = 16$ 



 $V_A \approx 0$ 



 $V_A = -127$ 

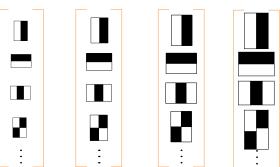
Haar Features are Sensitive to Directionality of Patterns





#### Haar Features

Compute Haar Features at different scales to detect faces of different sizes.

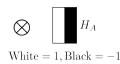






#### Haar Features





Response to Filter 
$$H_A$$
 at location  $(i, j)$ :
$$V_A[i, j] = \sum_n \sum_n I[m - i, n - j] H_A[m, n]$$

$$V_A[i, j] = \sum_n \text{(pixel intensities in white area)}$$

$$-\sum_n \text{(pixels intensities in black area)}$$





# Haar Features: Computation Cost



 $Value = \sum (pixel\ intensities\ in\ white\ area) - \sum (pixels\ intensities\ in\ black\ area)$  Computation cost = (N×M - 1) additions per pixel per filter per scale.

Can We Do Better?





# Integral Image

A table that holds the sum of all pixel values to the left and top of a given pixel, inclusive.

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

98	208	329	454	576	705
197	417	658	899	1137	1395
294	623	988	1340	1701	2093
392	833	1330	1790	2274	2799
489	1043	1687	2255	2864	3531
584	1249	2061	2751	3490	4294
680	1449	2433	3253	4118	5052

Integral Image II





# Integral Image

A table that holds the sum of all pixel values to the left and top of a given pixel, inclusive.

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

98	208	329	454	576	705
197	417	658	899	1137	1395
294	623	988	1340	1701	2093
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Integral Image II





# Integral Image

A table that holds the sum of all pixel values to the left and top of a given pixel, inclusive.

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
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98	208	329	454	576	705
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489	1043	1687	2255	2864	3531
584	1249	2061	2751	3490	4294
680	1449	2433	3253	4118	5052

Integral Image II





# Summation Within a Rectangle

#### Fast summations of arbitrary rectangles using integral images

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

98	208	329	454	576	705
197	417	658	899	1137	1395
294	623	988	1340	1701	2093
392	833	1330	1790	2274	2799
489	1043	1687	2255	2864	3531
584	1249	2061	2751	3490	4294
680	1449	2433	3253	4118	5052
	197 294 392 489 584	98 208 197 417 294 623 392 833 489 1043 584 1249	98 208 329 197 417 658 294 623 988 392 833 1330 489 1043 1687 584 1249 2061	98 208 329 454 197 417 658 899 294 623 988 1340 392 833 1330 1790 489 1043 1687 2255 584 1249 2061 2751	98 208 329 454 576 197 417 658 899 1137 294 623 988 1340 1701 392 833 1330 1790 2274 489 1043 1687 2255 2864 584 1249 2061 2751 3490

Integral Image  ${\cal II}$ 





#### Fast summations of arbitrary rectangles using integral images

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

$$Sum = II_P + \dots$$
$$= 3490 + \dots$$

98	208	329	454	576	705
197	417	658	899	1137	1395
294	623	988	1340	1701	2093
392	833	1330	1790	2274	2799
489	1043	1687	2255	2864	3531
584	1249	2061	2751	3490 4	4294
680	1449	2433	3253	4118	5052

Integral Image  ${\cal II}$ 





#### Fast summations of arbitrary rectangles using integral images

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image 
$$I$$

$$Sum = II_P - II_Q + \dots$$
  
= 3490 - 1137 + ...

329	454	576	705		
658	899	1137	1395	0	
988	1340	1701	2093	-Q	
1330	1790	2274	2799		
3 1687	2255	2864	3531		
9 2061	2751	3490 4	4294	-	
9 2433	3253	4118	5052	-P	
	658 988 1330 3 1687 9 2061	988 1340 1330 1790 3 1687 2255 9 2061 2751	658 899 1137 988 1340 1701 1330 1790 2274 3 1687 2255 2864 9 2061 2751 3490	658 899 1137 1305 988 1340 1701 2093 1330 1790 2274 2799 3 1687 2255 2864 3531 9 2061 2751 3490 4294	

Integral Image II





#### Fast summations of arbitrary rectangles using integral images

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

$$Sum = II_P - II_Q - II_S + \dots$$
  
=  $3490 - 1137 - 1249 + \dots$ 



Integral Image II





#### Fast summations of arbitrary rectangles using integral images

98	110	121	125	122	129
99	110	120	116	116	129
97	109	124	111	123	134
98	112	132	108	123	133
97	113	147	108	125	142
95	111	168	122	130	137
96	104	172	130	126	130

Image I

$$Sum = II_P - II_Q - II_S + II_R$$
  
=  $3490 - 1137 - 1249 + 417 = 1521$ 

Computation Cost: Only 3 additions





## Haar Response using Integral Image

98	110	121	125	122	129
99	110	120			129
97	109	124			134
98	112	132	108		133
97	113	147	108		142
95	111	168		130	137
96	104	172	130	126	130

98	208	329	454	576	705
197	417	658	899	1137	1395
294	623	988	1340	1701	2093
392	833	1330	1790	2274	2799
489	1043	1687	2255	2864	3531
584	1249	2061	2751	3490	4294
680	1449	2433	3253	4118	5052

Image I

Integral Image II

$$V_A = \sum (pixels \ in \ white) - \sum (pixels \ in \ black)$$





## Haar Response using Integral Image

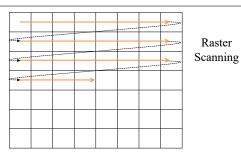
													- T
98	110	121	125	122	129	D (	98	208	329 .	454	576	705	
99	110	120	116		129	R	197	417	658	899	1137	1395	-Q
97	109	124	111		134		294	623	988	1340	1701	2093	
98	112	132	108		133		392	833	1330	1790	2274	2799	
97	113	147	108		142		489	1043	1687	2255	2864	3531	
95	111	168	122	130	137		584	1249	> 2061	2751	3490 <	4294	
96	104	172	130	126	130	$S \longrightarrow S$	680	1449	2433	3253	4118	5052	P
	Image I Integral Image II												
$V_A = \sum (pixel\ intensities\ in\ white}) - \sum (pixel\ intensities\ in\ black)$													
$= (II_O - II_T + II_R - II_S) - (II_P - II_Q + II_T - II_O)$													
-	= (2061 - 329 + 98 - 584) - (3490 - 576 + 329 - 2061) = 64												

Computation Cost: Only 7 additions





# Computing Integral Image

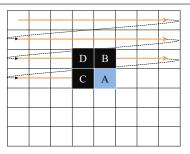






Raster

## Computing Integral Image



Raster Scanning

Let  $I_A$  and  $II_A$  be the values of Image and Integral Image, respectively, at pixel A.

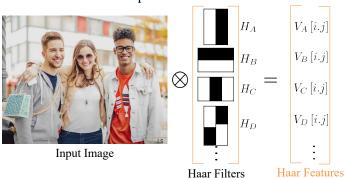
$$II_A = II_B + II_C - II_D + \underline{I}_A$$





### Haar Features using Integral Images

Integral image needs to be computed once per test image.
Allows fast computations of Haar features.

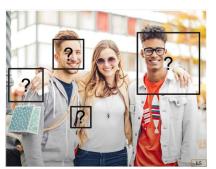






### Classifier for Face Detection?

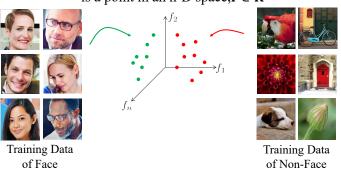
Given the features for a window, how to decide whether it contains a face or not?







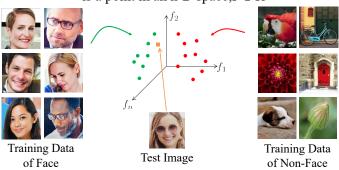
Haar Features  $\mathbf{f}$  (a vector) at a pixel is a point in an n-D space,  $\mathbf{f} \in \mathbf{R}^n$ 







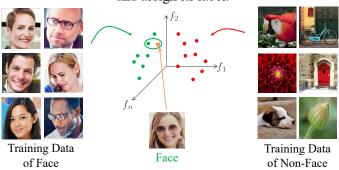
Haar Features  $\mathbf{f}$  (a vector) at a pixel is a point in an n-D space,  $\mathbf{f} \in \mathbf{R}^n$ 







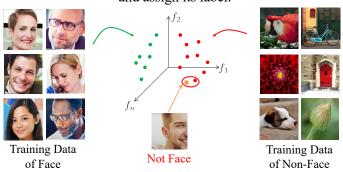
Find the Nearest training sample using L<sup>2</sup> distance and assign its label.







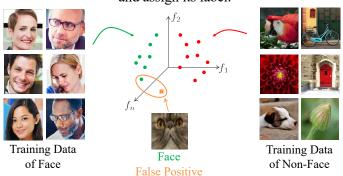
Find the Nearest training sample using L<sup>2</sup> distance and assign its label.







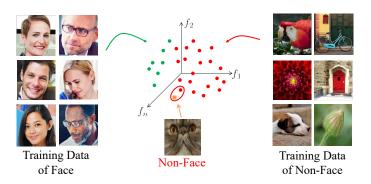
Find the Nearest training sample using L<sup>2</sup> distance and assign its label.







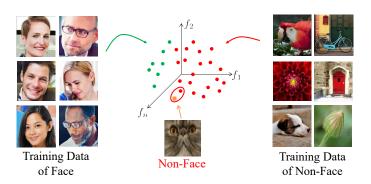
#### Larger the training set, more robust the NN classifier







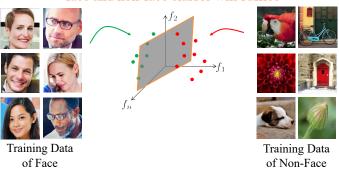
#### Larger the training set, slower the NN classifier







A simple dicision boundary separating face and non-face classes will suffice

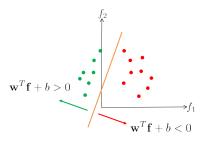






### Linear Dicision Boundaries

#### A Linear Decision Boundary in 2-D space is a 1-D Line



#### Equation of Line:

$$w_1 f_1 + w_2 f_2 + b = 0$$
$$\begin{bmatrix} w_1 & w_2 \end{bmatrix} \begin{bmatrix} f_1 \\ f_2 \end{bmatrix} + b = 0$$

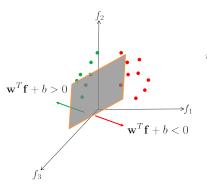
$$\mathbf{w}^T \mathbf{f} + b = 0$$





### Linear Dicision Boundaries

#### A Linear Decision Boundary in 3-D space is a 2-D Plane



#### Equation of Plane:

$$w_1 f_1 + w_2 f_2 + w_3 f_3 + b = 0$$

$$\mathbf{w}^T \mathbf{f} + b = 0$$





### Linear Dicision Boundaries

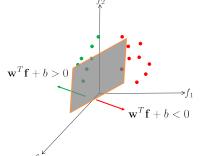
#### A Linear Decision Boundary in n-D space

is a (n-1)-D Hyperplane

#### Equation of Hyperplane:

$$w_1 f_1 + w_2 f_2 + \dots + w_n f_n + b = 0$$

$$\mathbf{w}^T \mathbf{f} + b = 0$$

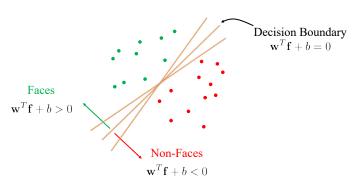






## Decision Boundary $(\mathbf{w},b)$

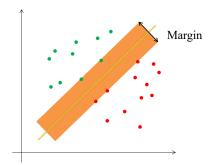
What is the optimal decision boundary?







## **Evaluating a Decision Boundary**

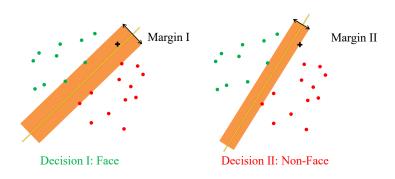


Margin or Safe Zone: The width that the boundary could be increased by, before hitting a feature point.





### **Evaluating a Decision Boundary**



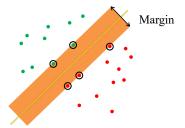
Choose Decision Boundary with Maximum Margin!





## Support Vector Machine (SVM)

Classifier optimized to Maximum Margin



Support Vectors: Closest data samples to the boundary Decision Boundary & Margin depend only on Support Vectors





## Support Vector Machine (SVM)

#### Given:

- k training images  $\{I_1, I_2, \dots, I_k\}$  and their Haar features  $\{\mathbf{f}_1, \mathbf{f}_2, \dots, \mathbf{f}_k\}$ .
- k corresponding labels  $\{\lambda_1, \lambda_2, \dots, \lambda_k\}$ , where  $\lambda_j = +1$  if  $I_j$  is a face and  $\lambda_i = -1$  if  $I_j$  is not a face.

#### Find:

Decision Boundary  $\mathbf{w}^T \mathbf{f} + b = 0$ with Maximum Margin  $\rho$ 

$$\mathbf{w}^T\mathbf{f} + b = 0$$



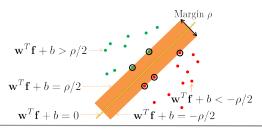


Margin  $\rho$ 

## Finding Decision Boundary $(\mathbf{w},b)$

For each training sample ( $\mathbf{f}_i$ ,  $\lambda_i$ ):

If 
$$\lambda_i = +1$$
:  $\mathbf{w}^T \mathbf{f}_i + b \ge \rho/2$   
If  $\lambda_i = -1$ :  $\mathbf{w}^T \mathbf{f}_i + b \le -\rho/2$   $\lambda_i(\mathbf{w}^T \mathbf{f}_i + b) \ge \rho/2$ 







## Finding Decision Boundary $(\mathbf{w},b)$

For each training sample ( $\mathbf{f}_i$ ,  $\lambda_i$ ):

If 
$$\lambda_i = +1$$
:  $\mathbf{w}^T \mathbf{f}_i + b \ge \rho/2$   
If  $\lambda_i = -1$ :  $\mathbf{w}^T \mathbf{f}_i + b \le -\rho/2$   $\lambda_i(\mathbf{w}^T \mathbf{f}_i + b) \ge \rho/2$ 

If **S** is the set of support vectors,

Then for every support vector  $s \in S$ :

$$\lambda_s(\mathbf{w}^T \mathbf{f}_s + b) = \rho/2$$

Numerical methods exist to find  $\mathbf{w}$ ,b and  $\mathbf{S}$  that maximize  $\rho$ 

MATLAB: symtrain





### Support Vector Machine (SVM)

Given: Haar features **f** for an image window and SVM parameters  $\mathbf{w}, b, \rho, \mathbf{S}$ 

#### Classification:

Compute 
$$d = \mathbf{w}^T \mathbf{f} + b$$
 
$$d \ge \rho/2 \quad \text{Face}$$
 
$$d > 0 \text{ and } d < \rho/2 \quad \text{Probably Face}$$
 
$$d < 0 \text{ and } d > -\rho/2 \quad \text{Probably Not-Face}$$
 
$$d \le -\rho/2 \quad \text{Not-Face}$$





### Face Detection Results







### Remarks

- Current face detection systems are mature but not perfect.
- Frontal and side poses usually require different face models.
- Successful vision technology used in cameras, surveillance, biometrics, search.
- Performance continues to improve.







# **Camera Calibration**

张举勇 中国科学技术大学

### **Camera Calibration**

• Method to find a camera's internal and external parameters.

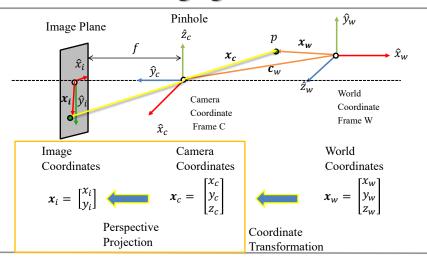
### Topics:

- (1) Linear Camera Model
- (2) Camera Calibration
- (3) Extracting Intrinsic and Extrinsic Matrices
- (4) Example Application: Simple Stereo





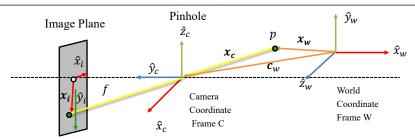
# Forward Imaging Model: 3D to 2D







# Forward Imaging Model: 3D to 2D



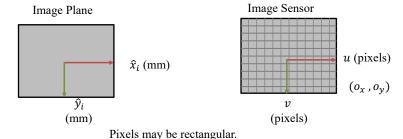
We know that 
$$\frac{x_i}{f} = \frac{x_c}{z_c}$$
 and  $\frac{y_i}{f} = \frac{y_c}{z_c}$ 

Therefore: 
$$x_i = f \frac{x_c}{z_c}$$
 and  $y_i = f \frac{y_c}{z_c}$ 





# Image Plane to Image Sensor Mapping



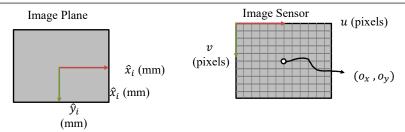
If  $m_x$  and  $m_y$  are the pixel densities (pixels/mm) in x and y directions, respectively, then pixel coordinates are:

$$u = m_x x_i = m_x f \frac{x_c}{z_c} \qquad v = m_y y_i = m_y f \frac{y_c}{z_c}$$





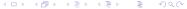
# Image Plane to Image Sensor Mapping



We usually treat the top-left corner of the image sensor as its origin (easier for indexing). If pixel  $(o_x, o_y)$  is the **Principle Point** where the optical axis pierces the sensor, then:

$$u = m_x f \frac{x_c}{z_c} + o_x \qquad \qquad v = m_y f \frac{y_c}{z_c} + o_y$$





### **Perspective Projection**

$$u = m_x f \frac{x_c}{z_c} + o_x \qquad v = m_y f \frac{y_c}{z_c} + o_y$$
$$u = f_x \frac{x_c}{z_c} + o_x \qquad v = f_y \frac{y_c}{z_c} + o_y$$

where:  $(f_x, f_y) = (m_x f, m_y f)$  are the focal lengths in pixels in the x and y directions.

 $(f_x, f_y, o_x, o_y)$ : **Intrinsic parameters** of the camera. They represent the **camera's internal geometry**.





# **Perspective Projection**

$$u = m_x f \frac{x_c}{z_c} + o_x \qquad v = m_y f \frac{y_c}{z_c} + o_y$$

$$u = f_x \frac{x_c}{z_c} + o_x \qquad \qquad v = f_y \frac{y_c}{z_c} + o_y$$

Equations for perspective projection are **Non-Linear**. It is convenient to express them as linear equations.



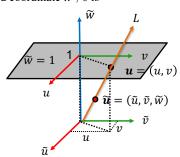


### **Homogenous Coordinates**

The homogenous representation of a 2D point  $\mathbf{u} = (u, v)$  is a 3D point  $\widetilde{\mathbf{u}} = (\widetilde{u}, \widetilde{v}, \widetilde{w})$ . The third coordinate  $\widetilde{w} \neq 0$  is fictitious such that:

$$u = \frac{\widetilde{u}}{\widetilde{w}} \qquad v = \frac{\widetilde{v}}{\widetilde{w}}$$

$$\boldsymbol{u} \equiv \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} \equiv \begin{bmatrix} \widetilde{w}u \\ \widetilde{w}v \\ \widetilde{u} \end{bmatrix} \equiv \begin{bmatrix} \widetilde{u} \\ \widetilde{v} \\ \widetilde{u} \end{bmatrix} = \widetilde{\boldsymbol{u}}$$



Every point on line L (except origin) represents the homogenous coordinate of u(u, v)





### **Homogenous Coordinates**

The **homogenous** representation of a 3D point  $x = (x, y, z) \in \mathbb{R}^3$  is a 4D point  $\widetilde{x} = (\widetilde{x}, \widetilde{y}, \widetilde{z}, \widetilde{w}) \in \mathbb{R}^4$ . The fourth coordinate  $w \neq 0$  is fictitious such that:

$$x = \frac{\widetilde{x}}{\widetilde{w}}$$
  $y = \frac{\widetilde{y}}{\widetilde{w}}$   $z = \frac{\widetilde{z}}{\widetilde{w}}$ 

$$\boldsymbol{x} \equiv \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix} \equiv \begin{bmatrix} \widetilde{w}x \\ \widetilde{w}y \\ \widetilde{w}z \\ \widetilde{w} \end{bmatrix} \equiv \begin{bmatrix} \widetilde{x} \\ \widetilde{y} \\ \widetilde{z} \\ \widetilde{w} \end{bmatrix} = \widetilde{\boldsymbol{x}}$$





# **Perspective Projection**

Perspective projection equations:

$$u = f_x \frac{x_c}{z_c} + o_x \quad v = f_y \frac{y_c}{z_c} + o_y$$

Homogenous coordinates of (u, v):

$$\begin{bmatrix} u \\ v \\ 1 \end{bmatrix} \equiv \begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} \equiv \begin{bmatrix} z_c u \\ z_c v \\ z_c \end{bmatrix} = \begin{bmatrix} f_x x_c + z_c o_x \\ f_y y_c + z_c o_y \\ z_c \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_c \\ y_c \\ z_c \\ 1 \end{bmatrix}$$

where:  $(u, v) = (\tilde{u}/\tilde{w}, \tilde{v}/\tilde{w})$ 

**Linear Model for Perspective Projection** 





### **Intrinsic Matrix**

$$\begin{bmatrix} u \\ v \\ 1 \end{bmatrix} \equiv \begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_c \\ y_c \\ z_c \\ 1 \end{bmatrix}$$

Calibration Matrix:

Intrinsic Matrix:

$$K = \begin{bmatrix} f_x & 0 & o_x \\ 0 & f_y & o_y \\ 0 & 0 & 1 \end{bmatrix}$$

$$M_{int} = [K \mid 0] = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix}$$

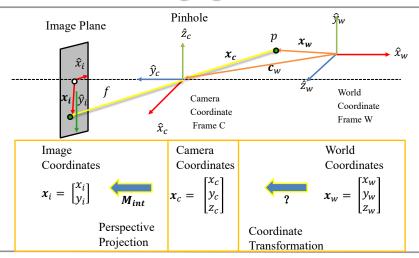
Upper Right Triangular Matrix

$$\widetilde{u} = [K \mid 0]\widetilde{\boldsymbol{x}}_c = M_{int}\widetilde{\boldsymbol{x}_c}$$





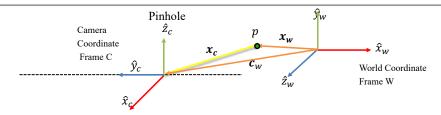
# Forward Imaging Model: 3D to 2D







#### Extrinsic Parameters



**Position**  $c_w$  and **Orientation** R of the camera in the world coordinate frame W are the camera's **Extrinsic Parameters**.

$$R = \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{31} & r_{32} & r_{33} \end{bmatrix} \xrightarrow{\hspace{0.5cm}} \text{Row 1: Direction of } \hat{x}_c \text{ in world coordinate frame}$$

$$\xrightarrow{\hspace{0.5cm}} \text{Row 2: Direction of } \hat{y}_c \text{ in world coordinate frame}$$

$$\xrightarrow{\hspace{0.5cm}} \text{Row 3: Direction of } \hat{z}_c \text{ in world coordinate frame}$$

Orientation/Rotation Matrix R is Orthonormal





### **Orthonormal Vectors and Matrices**

**Orthonormal Vectors**: Two vectors **u** and **v** are orthonormal if and only if:

$$dot(u, v) = u^T v = 0$$
 and  $u^T u = v^T v = 1$   
(Orthogonality) (Unit length)

Example: The x-, y- and z-axes of R<sup>3</sup> Euclidean space

**Orthonormal Matrix**: A square matrix R whose row (or column) vectors are orthonormal. For such a matrix:

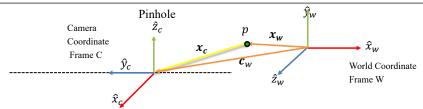
$$R^{-1} = R^T \qquad R^T R = R R^T = I$$

A Rotation Matrix is an Orthonormal Matrix





### **World-to-Camera Transformation**



Given the **extrinsic parameters**  $(R, c_w)$  of the camera, the camera-centric location of the point P in the world coordinate frame is:

$$\mathbf{x}_c = R(\mathbf{x}_w - \mathbf{c}_w) = R\mathbf{x}_w - R\mathbf{c}_w = R\mathbf{x}_w + \mathbf{t}$$

$$\mathbf{t} = -R\mathbf{c}_w$$

$$\mathbf{x}_c = \begin{bmatrix} x_c \\ y_c \\ z_c \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{21} & r_{22} & r_{23} \\ r_{22} & r_{23} & r_{23} \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \end{bmatrix} + \begin{bmatrix} t_x \\ t_y \\ t \end{bmatrix}$$





### **Extrinsic Matrix**

Rewriting using homogenous coordinates:

$$x_c = \begin{bmatrix} x_c \\ y_c \\ z_c \\ 1 \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_2 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \\ 1 \end{bmatrix}$$

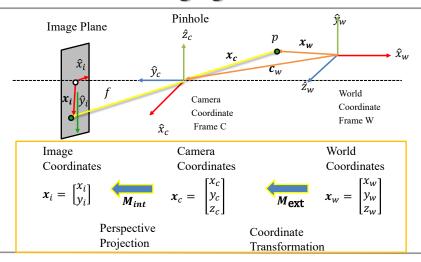
Extrinsic Matrix: 
$$M_{ext} = \begin{bmatrix} R_{3\times3} & t \\ 0_{1\times3} & 1 \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$\widetilde{x_c} = M_{\text{ext}} \, \widetilde{x_w}$$





# Forward Imaging Model: 3D to 2D







### **Projection Matrix P**

#### Camera to Pixel

$$\begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_c \\ y_c \\ z_c \\ 1 \end{bmatrix}$$

$$\widetilde{u} = \boldsymbol{M_{int}}\widetilde{x_c}$$

#### World to Camera

$$\begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_c \\ y_c \\ z_c \\ 1 \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_2 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \\ 1 \end{bmatrix}$$

$$\widetilde{x_c} = M_{\text{ext}}\widetilde{x_w}$$

Combining the above two equations, we get the full **projection matrix P**:

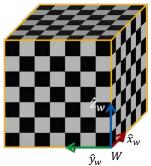
$$\widetilde{u} = M_{int} M_{ext} \widetilde{x_w} = P \widetilde{x_w}$$

$$\begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \\ 1 \end{bmatrix}$$





Step1: Capture an image of an object with known geometry.

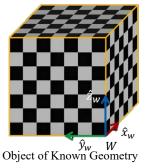


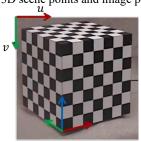
Object of Known Geometry





Step 2: Identify correspondences between 3D scene points and image points.



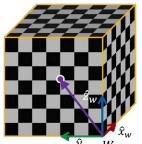


Captured Image



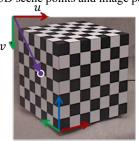


Step 2: Identify correspondences between 3D scene points and image points.



 $\hat{y}_w$  W Object of Known Geometry

• 
$$x_w = \begin{bmatrix} x_w \\ y_w \\ z_w \end{bmatrix} = \begin{bmatrix} 0 \\ 3 \\ 4 \end{bmatrix}$$
 (inches)



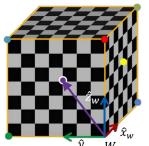
Captured Image

• 
$$u = \begin{bmatrix} u \\ v \end{bmatrix} = \begin{bmatrix} 56 \\ 115 \end{bmatrix}$$
 (pixels)



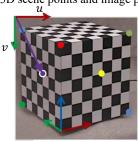


Step 2: Identify correspondences between 3D scene points and image points.



 $\hat{y}_w W$ Object of Known Geometry

• 
$$x_w = \begin{bmatrix} x_w \\ y_w \\ z_w \end{bmatrix} = \begin{bmatrix} 0 \\ 3 \\ 4 \end{bmatrix}$$
 (inches)



Captured Image

• 
$$u = \begin{bmatrix} u \\ v \end{bmatrix} = \begin{bmatrix} 56 \\ 115 \end{bmatrix}$$
 (pixels)





Step 3: For each corresponding point i in scene and image:

$$\begin{bmatrix} u^{(i)} \\ v^{(i)} \\ 1 \end{bmatrix} \equiv \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} x_w^{(i)} \\ y_w^{(i)} \\ z_w^{(i)} \\ 1 \end{bmatrix}$$
Known
Unknown
Known

Expanding the matrix as linear equations:

$$\begin{split} u^{(i)} &= \frac{p_{11}x_w^{(i)} + p_{12}y_w^{(i)} + p_{13}z_w^{(i)} + p_{14}}{p_{31}x_w^{(i)} + p_{32}y_w^{(i)} + p_{33}z_w^{(i)} + p_{34}}\\ v^{(i)} &= \frac{p_{21}x_w^{(i)} + p_{22}y_w^{(i)} + p_{23}z_w^{(i)} + p_{24}}{p_{31}x_w^{(i)} + p_{32}y_w^{(i)} + p_{33}z_w^{(i)} + p_{34}} \end{split}$$





Step4: Rearranging the terms

	1		$\mathcal{L}$	U								$\lceil P_{11} \rceil$		
$\left[x_w^{(1)}\right]$	$y_{w}^{(1)}$	$z_{w}^{(1)}$	1	0	0	0	0	$-u_1 x_w^{(1)}$	$-u_1 y_w^{(1)}$	$-u_1 z_w^{(1)}$	$-u_1$	$p_{12}$	[0]	
0	0	0	0	$x_{w}^{(1)}$	$y_{w}^{(1)}$	$z_{w}^{(1)}$	1	$-v_1 x_w^{(1)}$	$-v_1 y_w^{(1)}$	$-v_1 z_w^{(1)}$	$-v_1$	$\begin{vmatrix} p_{13} \\ p_{14} \end{vmatrix}$	0	
:	:	:	÷	:	:	:	÷	i	•	:	:	$p_{21}$	0	
$x_w^{(i)}$	$y_w^{(i)}$	$z_w^{(i)}$	1	0	0	0	0	$-u_i x_w^{(i)}$	$-u_i y_w^{(i)}$	$-u_i z_w^{(i)}$	$-u_i$	$p_{22}$	= 0	
0	0	0	0	$x_w^{(i)}$	$y_w^{(i)}$	$z_w^{(i)}$	1	$-v_ix_w^{(i)}$	$-v_i y_w^{(i)}$	$-v_i z_w^{(i)}$	$-v_i$	$\begin{vmatrix} p_{23} \\ n_{23} \end{vmatrix}$	0	
$\begin{vmatrix} \vdots \\ \mathbf{x}(n) \end{vmatrix}$	; ,(n)	; (n)	:	:	:	:	:	: ,, , (n)	: (n)	: (n)	:	$\begin{vmatrix} p_{24} \\ p_{31} \end{vmatrix}$	0	
$x_w^{(r)}$	$y_w$	$z_w^{(n)}$	1	0	0	0	0	$-u_n x_w$	$-u_n y_w$	$-u_n z_w^{(n)}$	$-u_n$	$p_{32}$	0	
L 0	0	0	0	$x_w^{(n)}$	$y_w^{(n)}$	$z_w^{(n)}$	1	$-v_n x_w^{(n)}$	$-v_n y_w^{(n)}$	$-v_n z_w^{(n)}$	$-v_n$		F07	I
							Δ					$\lfloor p_{34} \rfloor$		

Know

Unknown

Step5: Solve for P

A p = 0





### **Scale of Projection Matrix**

Projection matrix acts on homogenous coordinates.

We know that: 
$$\begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix} \equiv k \begin{bmatrix} \tilde{u} \\ \tilde{v} \\ \tilde{w} \end{bmatrix}$$
 (k \neq 0 is any constant)

That is: 
$$\begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \\ 1 \end{bmatrix} \equiv k \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{24} \\ p_{31} & p_{32} & p_{33} & p_{34} \end{bmatrix} \begin{bmatrix} x_w \\ y_w \\ z_w \\ 1 \end{bmatrix}$$

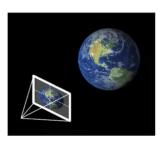
Therefore, Projection Matrices P and kP produce the same homogenous pixel coordinates.

Projection Matrix P is defined only up to a scale.

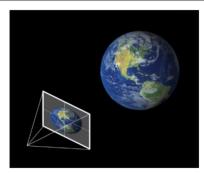




### **Scale of Projection Matrix**



Scale =  $k_1$ 



Scale =  $k_2$ 

Scaling projection matrix, implies simultaneously scaling the world and camera, which does not change the image.

Set projection matrix to some arbitrary scale!





# **Least squares Solution for P**

Option 1: Set scale so that:  $p_{34} = 1$ 

Option 2: Set scale so that:  $||p||^2 = 1$ 

We want Ap as close to 0 as possible and  $||p||^2 = 1$ :

$$\min_{\mathbf{p}} ||\mathbf{A}\mathbf{p}||^2 \text{ such that } ||\mathbf{p}||^2 = 1$$

$$\min_{\mathbf{p}} (p^T A^T A p) \text{ such that } p^T p = 1$$

Define Loss function  $L(\mathbf{p}, \lambda)$ :

$$L(\boldsymbol{p}, \lambda) = \boldsymbol{p}^T A^T A \boldsymbol{p} - \lambda (\boldsymbol{p}^T \boldsymbol{p} - 1)$$

(Similar to Solving Homography in Image Stitching)





# **Constrained Least Squares Solution**

Taking derivatives of  $L(\mathbf{p}, \lambda)$  w.r.t  $\mathbf{p}$ :  $2A^T A \mathbf{p} - 2\lambda \mathbf{p} = 0$ 

 $A^T A \boldsymbol{p} = \lambda \boldsymbol{p}$ 

**Eigenvalue Problem** 

Eigenvector p with smallest eigenvalue  $\lambda$  of matrix  $A^TA$  minimizes the loss function L(p).

Rearrange solution p to form the projection matrix P.





# **Extracting Intrinsic/Extrinsic Parameters**

We know that:

$$P = \begin{bmatrix} p_{11} & p_{12} & p_{13} \\ p_{21} & p_{22} & p_{23} \\ p_{31} & p_{32} & p_{33} \end{bmatrix} p_{14} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$M_{\text{int}}$$

That is:

$$\begin{bmatrix} p_{11} & p_{12} & p_{13} \\ p_{21} & p_{22} & p_{23} \\ p_{31} & p_{32} & p_{33} \end{bmatrix} = \begin{bmatrix} r_{1} & r_{12} & r_{13} \\ 0 & r_{2} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{31} & r_{32} & r_{33} \end{bmatrix} = KR$$

Given that K is an **Upper Right Triangular** matrix and R is an **Orthonormal** matrix, it is possible to uniquely "**decouple**" K and R from their product using "**QR factorization**".





# **Extracting Intrinsic/Extrinsic Parameters**

We know that:

$$P = \begin{bmatrix} p_{11} & p_{12} & p_{13} & p_{14} \\ p_{21} & p_{22} & p_{23} & p_{34} \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x & 0 \\ 0 & f_y & o_y & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

 $M_{\rm int}$ 

That is:

$$\begin{bmatrix} p_{14} \\ p_{24} \\ p_{34} \end{bmatrix} = \begin{bmatrix} f_x & 0 & o_x \\ 0 & f_y & o_y \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} t_x \\ t_y \\ t_z \end{bmatrix} = Kt$$

Therefore:

$$\mathbf{t} = K^{-1} \begin{bmatrix} p_{14} \\ p_{24} \\ p_{34} \end{bmatrix}$$

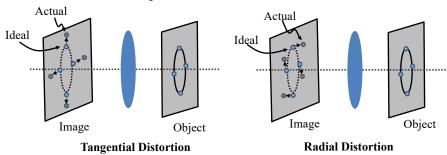




 $M_{\rm ext}$ 

#### **Camera Calibration**

Pinholes do not exhibit image distortions. But, lenses do!



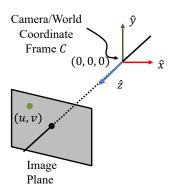
The intrinsic model of the camera will need to include the distortion coefficients. We ignore distortions here.





# **Backward Projection: From 2D to 3D**

Given a calibrated camera, can we find the 3D scene point from a single 2D image?

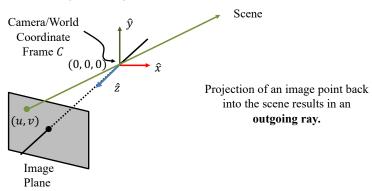






### **Backward Projection: From 2D to 3D**

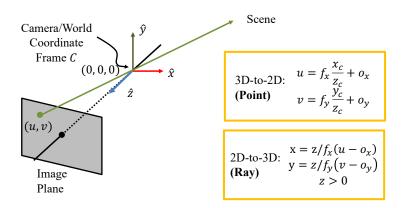
Given a calibrated camera, can we find the 3D scene point from a single 2D image?







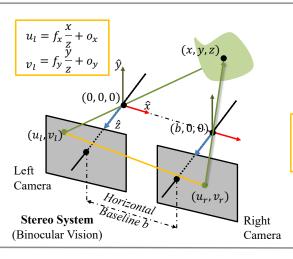
# **Computing 2D-to-3D Outgoing Ray**







### **Triangulation using Two Cameras**



$$u_r = f_x \frac{x - b}{\frac{y}{z}} + o_x$$
$$v_r = f_y \frac{y}{z} + o_y$$

 $f_x$ ,  $f_y$ , b,  $o_x$ ,  $o_y$  are known.





# Simple Stereo: Depth and Disparity

From perspective projection:

$$(u_l, v_l) = \left(f_x \frac{x}{z} + o_x, f_y \frac{y}{z} + o_y\right) \quad (u_r, v_r) = \left(f_x \frac{x - b}{z} + o_x, f_y \frac{y}{z} + o_y\right)$$

Solving for (x, y, z):

$$x = \frac{b(u_l - o_x)}{(u_l - u_r)}$$
  $y = \frac{bf_x(v_l - o_y)}{f_y(u_l - u_r)}$   $z = \frac{bf_x}{(u_l - u_r)}$ 

where  $(u_l - u_r)$  is called **Disparity**.

Depth z is inversely proportional to Disparity.

Disparity is proportional to Baseline.





### A Simple Stereo Camera



Fujifilm FinePix REAL 3D W3





# **Stereo Matching: Finding Disparities**

Goal: Find the disparity between left and right stereo pairs.



Left/Right Camera Images



Disparity Map(Ground Truth)

From perspective projection:

$$v_l = v_r = f_y \frac{y}{z} + o_y$$

Corresponding scene points lie on the same horizontal scan line.





#### Window Based Methods

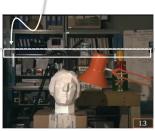
#### Determine Disparity using Template Matching

Template Window T



Left Camera Image  $E_l$ 

Search Scan Line L



Right Camera Image  $E_r$ 

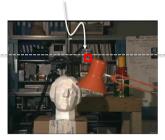




### **Window Based Methods**

#### Determine Disparity using Template Matching

Template Window T



Left Camera Image  $E_l$ 

Search Scan Line L



Right Camera Image  $E_r$ 

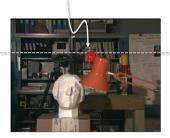




#### Window Based Methods

#### Determine Disparity using Template Matching

Template Window T



Left Camera Image  $E_l$ Disparity:  $d = u_l - u_r$ 

Search Scan Line L



Right Camera Image  $E_r$ Depth:  $z = \frac{bf_x}{(u_l - u_r)}$ 





# **Similarity Metrics for Template Matching**

Find pixel  $(k, l) \in L$  with Minimum **Sum of Absolute Differences:** 

$$SAD(k,l) = \sum\nolimits_{(i,j) \in T} |E_l(i,j) - E_r(i+k,j+l)|$$

Find pixel  $(k, l) \in L$  with Minimum Sum of Squared Differences:

$$SSD(k,l) = \sum\nolimits_{(i,j) \in T} |E_l(i,j) - E_r(i+k,j+l)|^2$$

Find pixel  $(k, l) \in L$  with Maximum Normalized Cross-Correlation:

$$NCC(k,l) = \frac{\sum_{(i,j) \in T} E_l(i,j) E_r(i+k,j+l)}{\sqrt{\sum_{(i,j) \in T} E_l(i,j)^2 \sum_{(i,j) \in T} E_r(i+k,j+l)^2}}$$





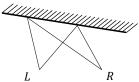
# **Issues with Stereo Matching**

• Surface must have (non-repetitive) texture





· Foreshortening effect makes matching challenging







### **How Large Should Window Be?**



Window size = 5 pixels (Sensitive to noise)



Window size = 30 pixels (Poor localization)

**Adaptive Window Method Solution:** For each point, match using windows of multiple sizes and use the disparity that is a result of the best similarity measure (minimize SSD per pixel).





### **Window Based Methods: Results**



Left Image



Right Image



Ground Truths



SSD - Adaptive Window



SD (Window size=21)



State of the Art

http://vision.middlebury.edu/stereo



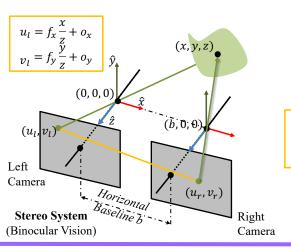




# **Uncalibrated Stereo**

张举勇 中国科学技术大学

### Simple (Calibrated) Stereo



$$u_r = f_x \frac{x - b}{\frac{z}{z}} + o_x$$
$$v_r = f_y \frac{y}{z} + o_y$$

 $f_x$ ,  $f_y$ , b,  $o_x$ ,  $o_y$  are in pixel units.





### **Depth and Disparity**

Solving for (x, y, z):

$$x = \frac{b(u_l - o_x)}{(u_l - u_r)}$$
  $y = \frac{bf_x(v_l - o_y)}{f_y(u_l - u_r)}$   $z = \frac{bf_x}{(u_l - u_r)}$ 

where  $(u_l - u_r)$  is called **Disparity**.





### **Uncalibrated stereo**

• Method to estimate 3D structure of a static scene from two arbitrary views.

#### Topics:

- (1) Problem of Uncalibrated Stereo
- (2) Epipolar Geometry
- (3) Estimating Fundamental Matrix
- (4) Finding Dense Correspondences
- (5) Computing Depth





### Uncalibrated Stereo

Compute 3D structure of static scene from two arbitrary views



Camera

**Instrinsics**  $(f_x, f_y, o_x, o_y)$  are **known** for both views/cameras.

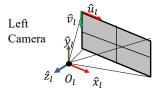
Extrinsics (relative position/orientation of cameras) are unknown.

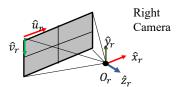




### **Uncalibrated Stereo**









- 1. Assume Camera Matrix K is known for each camera
  - 2. Find a few Reliable Corresponding Points





### **Initial Correspondence**

Find a set of corresponding features (at least 8) in left and right images (e.g. using SIFT or hand-picked).

Left image



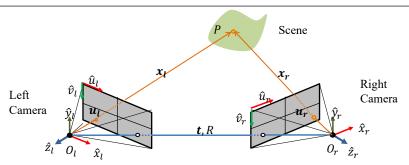
Right image







### **Uncalibrated Stereo**

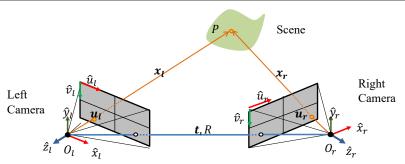


- 2 1. Assume Camera Matrix K is known for each camera
- 2. Find a few Reliable Corresponding Points
  - 3. Find Relative Camera Position t and Orientation R
  - 4. Find Dense Correspondence
  - 5. Compute Depth using Triangulation





## **Epipolar Geometry: Epipoles**



**Epipole**: Image point of origin/pinhole of one camera as viewed by the other camera.

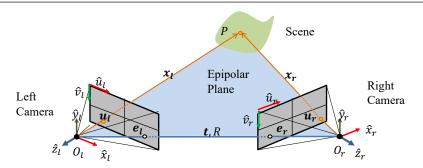
 $e_1$  and  $e_r$  are the epipoles.

 $e_l$  and  $e_r$  are unique for a given stereo pair.





# **Epipolar Geometry: Epipolar Plane**



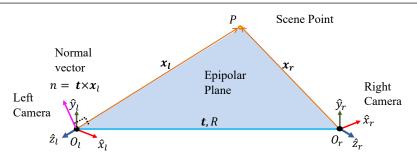
**Epipolar Plane of Scene Point P**: The plane formed by camera origins  $(O_1 \text{ and } O_r)$ , epipoles  $(e_1 \text{ and } e_r)$  and scene point P.

Every scene point lies on a unique epipolar plane.





### **Epipolar Constraint**



Vector normal to the epipolar plane:  $n = t \times x_1$ 

Dot product of n and  $x_l$  (perpendicular vectors) is zero:

$$\mathbf{x}_l \cdot (\mathbf{t} \times \mathbf{x}_l) = 0$$





### **Epipolar Constraint**

Writing the epipolar constraint in matrix form:

$$\begin{aligned} x_l \cdot (t \times x_l) &= 0 \\ [x_l \quad y_l \quad z_l] \begin{bmatrix} t_y z_l - t_z y_l \\ t_z x_l - t_x z_l \\ t_x y_l - t_y x_l \end{bmatrix} &= 0 \end{aligned}$$

Cross-product definition

$$\begin{bmatrix} x_l & y_l & z_l \end{bmatrix} \begin{bmatrix} 0 & -t_z & t_y \\ t_z & 0 & -t_x \\ -t_y & t_x & 0 \end{bmatrix} \begin{bmatrix} x_l \\ y_l \\ z_l \end{bmatrix} = 0$$

Matrix-vector form

 $T_{\times}$ t<sub>3×1</sub>: Position of Right Camera in Left Camera's Frame  $R_{3\times3}$ : Orientation of Left Camera in Right Camera's Frame

 $\begin{bmatrix} x_l \\ y_l \\ z_l \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{21} & r_{22} & r_{23} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_l \end{bmatrix} + \begin{bmatrix} t_x \\ t_y \\ t_z \end{bmatrix}$ 





### **Epipolar Constraint**

Substituting into the epipolar constraint gives:

$$\begin{bmatrix} x_{l} & y_{l} & z_{l} \end{bmatrix} \begin{pmatrix} 0 & -t_{z} & t_{y} \\ t_{z} & 0 & -t_{x} \\ -t_{y} & t_{x} & 0 \end{pmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{31} & r_{32} & r_{33} \end{bmatrix} \begin{bmatrix} x_{r} \\ y_{r} \\ z_{r} \end{bmatrix} + \begin{bmatrix} 0 & -t_{z} & t_{y} \\ t_{z} & 0 & -t_{x} \\ -t_{y} & t_{x} & 0 \end{bmatrix} \begin{bmatrix} t_{x} \\ t_{y} \\ t_{z} \end{bmatrix} = 0$$

$$\mathbf{t} \times \mathbf{t} = 0$$

$$\begin{bmatrix} x_{l} & y_{l} & z_{l} \end{bmatrix} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} \begin{bmatrix} x_{r} \\ y_{r} \\ z_{r} \end{bmatrix} = 0$$

$$\mathbf{Escential Matrix} \mathbf{E}$$

**Essential Matrix E** 

$$E = T_{\times}R$$

[Longuet-Higgins 1981]





### **Essential Matrix E: Decomposition**

$$E = T_{\times}R$$

$$\begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} = \begin{bmatrix} 0 & t_z & t_y \\ t_z & 0 & t_z \\ -t_y & t_x & 0 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} \\ r_{21} & r_{22} & r_{23} \\ r_{31} & r_{32} & r_{33} \end{bmatrix}$$

Given that Tx is a **Skew-Symmetric** matrix ( $a_{ij} = -a_{ji}$ ) and R is an **Orthonormal** matrix, it is possible to "**decouple**"  $T_{\times}$  and R from their product using "**Singular Value Decomposition**".

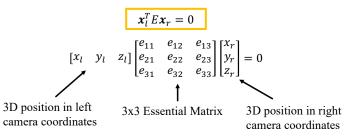
**Take Away**: If E is known, we can calculate **t** and R.





#### How do we find E?

Relates 3D position  $(x_l, y_l, z_l)$  of scene point w.r.t left camera to its 3D position  $(x_r, y_r, z_r)$  w.r.t right camera



Unfortunately, we don't have  $x_l$  and  $x_r$ .

But we do know corresponding points in image coordinates.





Perspective projection equations for left camera:

$$\begin{split} u_l &= f_x^{(l)} \frac{x_l}{z_l} + o_x^{(l)} & v_l &= f_y^{(l)} \frac{y_l}{z_l} + o_y^{(l)} \\ \\ z_l u_l &= f_x^{(l)} x_l + z_l o_x^{(l)} & z_l v_l &= f_y^{(l)} y_l + z_l o_y^{(l)} \end{split}$$

Representing in matrix form:

$$z_{l} \begin{bmatrix} u_{l} \\ v_{l} \\ 1 \end{bmatrix} = \begin{bmatrix} z_{l}u_{l} \\ z_{l}v_{l} \\ z_{l} \end{bmatrix} = \begin{bmatrix} f_{x}^{(l)}x_{l} + z_{l}o_{x}^{(l)} \\ f_{y}^{(l)}y_{l} + z_{l}o_{y}^{(l)} \\ z_{l} \end{bmatrix} = \begin{bmatrix} f_{x}^{(l)} & 0 & o_{x}^{(l)} \\ 0 & f_{y}^{(l)} & o_{y}^{(l)} \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_{l} \\ y_{l} \\ z_{l} \end{bmatrix}$$

Known Camera Matrix *K<sub>i</sub>* 





Left camera:

$$z_{l} \begin{bmatrix} u_{l} \\ v_{l} \\ 1 \end{bmatrix} = \begin{bmatrix} f_{x}^{(l)} & 0 & o_{x}^{(l)} \\ 0 & f_{y}^{(l)} & o_{y}^{(l)} \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_{l} \\ y_{l} \\ z_{l} \end{bmatrix}$$

$$K_{l}$$

$$\boldsymbol{x}_l^T = [u_l \quad v_l \quad 1] z_l K_l^{-1}$$

Right camera:

$$z_{l} \begin{bmatrix} u_{l} \\ v_{l} \\ 1 \end{bmatrix} = \underbrace{\begin{bmatrix} f_{x}^{(l)} & 0 & o_{x}^{(l)} \\ 0 & f_{y}^{(l)} & o_{y}^{(l)} \\ 0 & 0 & 1 \end{bmatrix}}_{K_{l}} \begin{bmatrix} x_{l} \\ y_{l} \\ z_{l} \end{bmatrix} \qquad z_{r} \begin{bmatrix} u_{r} \\ v_{r} \\ 1 \end{bmatrix} = \underbrace{\begin{bmatrix} f_{x}^{(r)} & 0 & o_{x}^{(r)} \\ 0 & f_{y}^{(r)} & o_{y}^{(r)} \\ 0 & 0 & 1 \end{bmatrix}}_{K_{r}} \begin{bmatrix} x_{r} \\ y_{r} \\ z_{r} \end{bmatrix}$$

$$\boldsymbol{x}_r K_r^{-1} \boldsymbol{Z}_r = \begin{bmatrix} c \boldsymbol{u}_r \\ \boldsymbol{v}_r \\ 1 \end{bmatrix}$$





Epipolar constraint:

$$\begin{bmatrix} x_l & y_l & z_l \end{bmatrix} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \end{bmatrix} = 0$$

Rewriting in terms of image coordinates:

$$\begin{bmatrix} u_l & v_l & 1 \end{bmatrix} \mathbf{z}_l^{\prime} K_l^{-1} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} K_r^{-1} \mathbf{z}_r^{\prime} \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} = 0$$

$$z_l \neq 0$$
  
$$z_r \neq 0$$





Epipolar constraint:

$$\begin{bmatrix} x_l & y_l & z_l \end{bmatrix} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \end{bmatrix} = 0$$

Rewriting in terms of image coordinates:

$$\begin{bmatrix} u_l & v_l & 1 \end{bmatrix} K_l^{-1} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} K_r^{-1} \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} = 0$$



### **Fundamental Matrix F**

Epipolar constraint:

$$\begin{bmatrix} x_l & y_l & z_l \end{bmatrix} \begin{bmatrix} e_{11} & e_{12} & e_{13} \\ e_{21} & e_{22} & e_{23} \\ e_{31} & e_{32} & e_{33} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \end{bmatrix} = 0$$

Rewriting in terms of image coordinates:

$$\begin{bmatrix} u_l & v_l & 1 \end{bmatrix} \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} = 0$$

#### Fundamental Matrix F

$$E = K_l^T F K_r$$

$$E = T_{\times}R$$

[Fagueras 1992, Luong 1992]





Find a set of corresponding features in left and right images(e.g. using SIFT or hand-picked)

Left image



#### Right image







Find a set of corresponding features in left and right images(e.g. using SIFT or hand-picked)

Left image



- $\begin{array}{c} \bullet \quad \left(u_l^{(1)}, v_l^{(1)}\right) \\ & \vdots \\ \bullet \quad \left(u_l^{(m)}, v_l^{(m)}\right) \end{array}$

Right image







**Step A**: For each correspondence i, write out epipolar constraint.

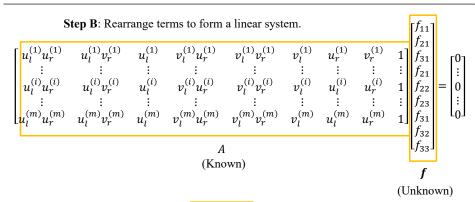
$$\frac{\begin{bmatrix} u_l^{(i)} & v_l^{(i)} & 1 \end{bmatrix} \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} u_r^{(i)} \\ v_r^{(i)} \\ 1 \end{bmatrix}}{\text{Known}} = 0$$
Known

Expand the matrix to get linear equation:

$$\left( f_{11} u_r^{(i)} + f_{12} v_r^{(i)} + f_{13} \right) u_l^{(i)} + \left( f_{21} u_r^{(i)} + f_{22} v_r^{(i)} + f_{23} \right) v_l^{(i)} + f_{31} u_r^{(i)} + f_{32} v_r^{(i)} + f_{33} \ = \ 0$$







$$A\mathbf{f} = 0$$





## The Tale of Missing Scale

Fundamental matrix acts on homogenous coordinates.

$$\begin{bmatrix} u_l & v_l & 1 \end{bmatrix} \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} = 0 = \begin{bmatrix} u_l & v_l & 1 \end{bmatrix} \begin{bmatrix} kf_{11} & kf_{12} & kf_{13} \\ kf_{21} & kf_{22} & kf_{23} \\ kf_{31} & kf_{32} & kf_{33} \end{bmatrix} \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix}$$

Fundamental Matrix F and kF describe the same epipolar geometry. That is, F is defined only up to a scale.

Set Fundamental Matrix to some arbitrary scale.

$$||f||^2 = 1$$





## Solving for *F*

**Step C**: Find least squares solution for fundamental matrix F.

We want Af as close to 0 as possible and  $||f||^2 = 1$ :

$$\min_{f} \|A\mathbf{f}\|^2 \text{ such that } \|\mathbf{f}\|^2 = 1$$

Constrained linear least squares problem

Like solving Projection Matrix during Camera Calibration. Or, Homography Matrix for Image Stitching.

Rearrange solution f to form the fundamental matrix F.





### **Extracting Rotation and Translation**

**Step D:** Compute essential matrix E from known left and right intrinsic camera matrices and fundamental matrix F.

$$E = K_l^T F K_r$$

**Step E**: Extract R and t from E.

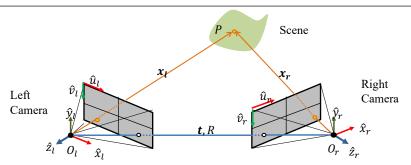
$$E = T_{\times}R$$

(Using Singular Value Decomposition)





### **Uncalibrated Stereo**



- 2 1. Assume Camera Matrix K is known for each camera
- 2. Find a few Reliable Corresponding Points
- 3. Find Relative Camera Position t and Orientation R
  - 4. Find Dense Correspondence
  - 5. Compute Depth using Triangulation





### **Simple Stereo: Finding Correspondences**

Goal: Find the disparity between left and right stereo pairs.



Left/Right Camera Images



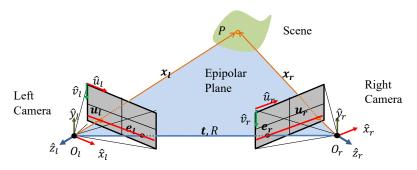
Disparity Map(Ground Truth)

Corresponding scene points lie on the **same horizontal scan-line** Finding correspondence is a **1D search**.





# **Epipolar Geometry: Epipolar Line**



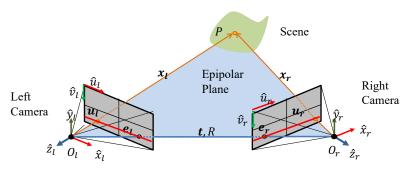
Epipolar Line: Intersection of image plane and epipolar plane.

Every scene point has **two corresponding epipolar lines**, one each on the two image planes.





# **Epipolar Geometry: Epipolar Line**



Given a point in one image, the corresponding point in the other image must lie on the epipolar line.

Finding correspondence reduces to a 1D search.





## Finding Epipolar Lines

**Given:** Fundamental matrix F and point on left image (u, v)

**Find**: Equation of Epipolar line in the right image

**Epipolar Constraint Equation:** 

$$\begin{bmatrix} u_l & v_l & 1 \end{bmatrix} \begin{bmatrix} f_{11} & f_{12} & f_{13} \\ f_{21} & f_{22} & f_{23} \\ f_{31} & f_{32} & f_{33} \end{bmatrix} \begin{bmatrix} \boldsymbol{u_r} \\ \boldsymbol{v_r} \\ \boldsymbol{1} \end{bmatrix} = 0$$

Expanding the matrix equation gives:

$$(f_{11}u_l+f_{21}v_l+f_{31})\boldsymbol{u}_r+(f_{12}u_l+f_{22}v_l+f_{32})\boldsymbol{v}_r+(f_{13}u_l+f_{23}v_l+f_{33})=0$$

Equation for **right epipolar line**:  $a_l u_r + b_l v_r + c_l = 0$ 

$$a_l \boldsymbol{u_r} + b_l \boldsymbol{v_r} + c_l = 0$$

Similarly we can calculate epipolar line in left image for a point in right image.





## **Finding Epipolar Lines: Example**

Given the Fundamental matrix,

$$F = \begin{bmatrix} -.003 & -.028 & 13.19 \\ -.003 & -.008 & -29.2 \\ 2.97 & 56.38 & -9999 \end{bmatrix}$$

Left Image



Right Image







## **Finding Epipolar Lines: Example**

Given the Fundamental matrix,

$$F = \begin{bmatrix} -.003 & -.028 & 13.19 \\ -.003 & -.008 & -29.2 \\ 2.97 & 56.38 & -9999 \end{bmatrix}$$

and the left image point

$$\widetilde{u_l} = \begin{bmatrix} 343 \\ 221 \\ 1 \end{bmatrix}$$

Left Image



Right Image



The equation for the epipolar line in the **right** image is:

$$\begin{bmatrix} u_r & v_r & 1 \end{bmatrix} \begin{bmatrix} -.003 & -.003 & 2.97 \\ -.028 & -.008 & 56.38 \\ 13.19 & -29.2 & -9999 \end{bmatrix} \begin{bmatrix} 343 \\ 221 \\ 1 \end{bmatrix} = 0$$





# **Finding Epipolar Lines: Example**

Given the Fundamental matrix,

$$F = \begin{bmatrix} -.003 & -.028 & 13.19 \\ -.003 & -.008 & -29.2 \\ 2.97 & 56.38 & -9999 \end{bmatrix}$$

and the left image point

$$\widetilde{u_l} = \begin{bmatrix} 343 \\ 221 \\ 1 \end{bmatrix}$$

Left Image



Right Image



Epipolar Line

The equation for the epipolar line in the **right** image is:

$$.03u_r + .99v_r - 265 = 0$$





### **Finding Correspondence**



Left Image



Epipolar Line

Right Image

Corresponding scene points lie on the epipolar lines. Finding correspondence is a **1D search**.





### **Finding Correspondence**



Left Image



Epipolar Line

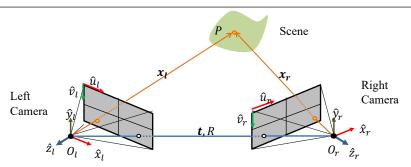
Right Image

Corresponding scene points lie on the epipolar lines. Finding correspondence is a **1D search**.





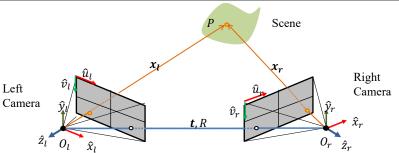
### **Uncalibrated Stereo**



- 2 1. Assume Camera Matrix K is known for each camera
- 2. Find a few Reliable Corresponding Points
- 3. Find Relative Camera Position t and Orientation R
- 4. Find Dense Correspondence
  - 5. Compute Depth using Triangulation







Given the intrinsic parameters, the projections of scene point on the two image sensors are:

$$\begin{bmatrix} u_l \\ v_l \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(l)} & 0 & o_x^{(l)} & 0 \\ 0 & f_y^{(l)} & o_y^{(l)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_l \\ y_l \\ z_l \\ 1 \end{bmatrix} \qquad \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(r)} & 0 & o_x^{(r)} & 0 \\ 0 & f_y^{(r)} & o_y^{(r)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$

$$\begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(r)} & 0 & o_x^{(r)} & 0 \\ 0 & f_y^{(r)} & o_y^{(r)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$





Left Camera Imaging Equation

$$\begin{bmatrix} u_l \\ v_l \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(l)} & 0 & o_x^{(l)} & 0 \\ 0 & f_y^{(l)} & o_y^{(l)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_l \\ y_l \\ z_l \\ 1 \end{bmatrix}$$

Right Camera Imaging Equation

$$\begin{bmatrix} u_l \\ v_l \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(l)} & 0 & o_x^{(l)} & 0 \\ 0 & f_y^{(l)} & o_y^{(l)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_l \\ y_l \\ z_l \\ 1 \end{bmatrix} \qquad \begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(r)} & 0 & o_x^{(r)} & 0 \\ 0 & f_y^{(r)} & o_y^{(r)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$

We also know the relative position and orientation between the two cameras.

$$\begin{bmatrix} x_l \\ y_l \\ z_l \\ 1 \end{bmatrix} = \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$





Left Camera Imaging Equation:

$$\begin{bmatrix} u_l \\ v_l \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(l)} & 0 & o_x^{(l)} & 0 \\ 0 & f_y^{(l)} & o_y^{(l)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} r_{11} & r_{12} & r_{13} & t_x \\ r_{21} & r_{22} & r_{23} & t_y \\ r_{31} & r_{32} & r_{33} & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$

$$\widetilde{\boldsymbol{u}_l} = P_l \widetilde{\boldsymbol{x}_r}$$

Right Camera Imaging Equation:

$$\begin{bmatrix} u_r \\ v_r \\ 1 \end{bmatrix} \equiv \begin{bmatrix} f_x^{(r)} & 0 & o_x^{(r)} & 0 \\ 0 & f_y^{(r)} & o_y^{(r)} & 0 \\ 0 & 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \\ 1 \end{bmatrix}$$

$$\widetilde{\boldsymbol{u}_r} = M_{int_r} \widetilde{\boldsymbol{x}_r}$$





The imaging equation:

$$\widetilde{u_r} = M_r \widetilde{x_r} \qquad \widetilde{u_l} = P_l \widetilde{x_r}$$
 
$$\widetilde{u_l} = P_l \widetilde{u_l}$$
 
$$\widetilde{u_l$$

Rearranging the terms:

$$\begin{bmatrix} u_r m_{31} - m_{11} & u_r m_{32} - m_{12} & u_r m_{33} - m_{13} \\ v_r m_{31} - m_{21} & v_r m_{32} - m_{22} & v_r m_{33} - m_{23} \\ u_l p_{31} - p_{11} & u_l p_{32} - p_{12} & u_l p_{33} - p_{13} \\ v_l p_{31} - p_{21} & v_l p_{32} - p_{22} & v_l p_{33} - p_{23} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \end{bmatrix} = \begin{bmatrix} m_{14} - m_{34} \\ m_{24} - m_{34} \\ p_{14} - p_{34} \\ p_{24} - p_{34} \end{bmatrix}$$





### **Computing Depth: Least Squares Solution**

$$\begin{bmatrix} u_r m_{31} - m_{11} & u_r m_{32} - m_{12} & u_r m_{33} - m_{13} \\ v_r m_{31} - m_{21} & v_r m_{32} - m_{22} & v_r m_{33} - m_{23} \\ u_l p_{31} - p_{11} & u_l p_{32} - p_{12} & u_l p_{33} - p_{13} \\ v_l p_{31} - p_{21} & v_l p_{32} - p_{22} & v_l p_{33} - p_{23} \end{bmatrix} \begin{bmatrix} x_r \\ y_r \\ z_r \end{bmatrix} = \begin{bmatrix} m_{14} - m_{34} \\ m_{24} - m_{34} \\ p_{14} - p_{34} \\ p_{24} - p_{34} \end{bmatrix}$$

$$A_{4\times3} \qquad x_r \qquad b_{4\times1}$$
(Known) (Known)

Find least squares solution using pseudo-inverse:

$$A\mathbf{x}_r = \mathbf{b}$$
 $A^T A\mathbf{x}_r = A^T \mathbf{b}$ 
 $\mathbf{x}_r = (A^T A)^{-1} A^T \mathbf{b}$ 





St. Peter's Basilica (1275 Images)



[Snavely 2006]





St. Peter's Basilica (1275 Images)



[Snavely 2006]





Piazza San Marco (13709 Images)



[Furukawa 2010]





Piazza San Marco (13709 Images)

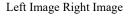


[Furukawa 2010]



### **Active Stereo Results**







3D Structure

[Zhang 2003]

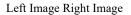




### **Active Stereo Results**









3D Structure

[Zhang 2003]







# **Optical Flow**

张举勇 中国科学技术大学

### **Overview**

Method to estimate apparent motion of scene points from a sequence of images

### **Topics:**

- (1) Motion Field and Optical Flow
- (2) Optical Flow Constraint Equation
- (3) Lucas-Kanada Method
- (4) Coarse-to-Fine Flow Estimation
- (5) Applications of Optical Flow





### **Motion Field**

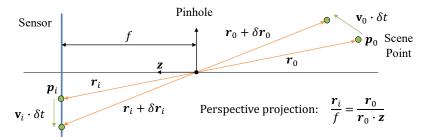


Image Point Velocity: 
$$\mathbf{v}_i = \frac{d\mathbf{r}_i}{dt} = f \frac{(\mathbf{r}_0 \cdot \mathbf{z})\mathbf{v}_0 - (\mathbf{v}_0 \cdot \mathbf{z})\mathbf{r}_0}{(\mathbf{r}_0 \cdot \mathbf{z})^2} : \mathbf{v}_0 = \frac{d\mathbf{r}_0}{dt}$$
(Motion Field)

$$\mathbf{v}_i = \frac{(\mathbf{r}_0 \times \mathbf{v}_0) \times \mathbf{z}}{(\mathbf{r}_0 \cdot \mathbf{z})^2}$$



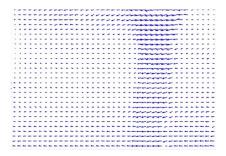


# **Optical Flow**

#### Motion of brightness patterns in the image



Image Sequence (2 frames)



Optical Flow

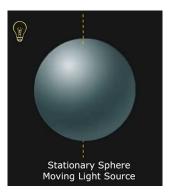




## When is Optical Flow $\neq$ Motion Field?



Motion Field exists But no Optical Flow



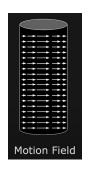
No Motion Field exists But there is Optical Flow





# When is Optical Flow $\neq$ Motion Field?



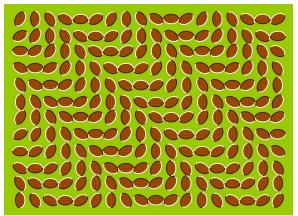








### **Motion Illusions**

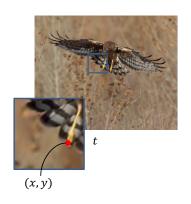


"Dongurakokko" (Donguri wave), produced by Akiyoshi Kitaoka in 2004 as an artwork of waving demonstration of the 'optimized' Fraser-Wilcox illusion Type IIa.

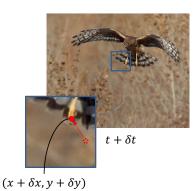




### **Optical Flow**



Displacement:  $(\delta x, \delta y)$ 

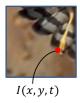


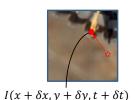
Optical Flow: 
$$(u, v) = (\frac{\delta x}{\delta t}, \frac{\delta y}{\delta t})$$





## **Optical Flow Constraint Equation**





#### **Assumption #1:**

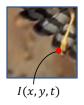
Brightness of image point remains constant over time

$$I(x + \delta x, y + \delta y, t + \delta t) = I(x, y, t)$$





### **Optical Flow Constraint Equation**





$$I(x + \delta x, y + \delta y, t + \delta t)$$

#### **Assumption #2:**

Dispacement  $(\delta x, \delta y)$  and time step  $\delta t$  are small

$$I(x+\delta x,y+\delta y,t+\delta t)=I(x,y,t)+\frac{\partial I}{\partial x}\delta x+\frac{\partial I}{\partial y}\delta y+\frac{\partial I}{\partial t}\delta t$$

$$I(x+\delta x,y+\delta y,t+\delta t)=I(x,y,t)+I_x\delta x+I_y\delta y+I_t\delta t$$





# **Optical Flow Constraint Equation**

$$I(x + \delta x, y + \delta y, t + \delta t) = I(x, y, t)$$
(1)

$$I(x + \delta x, y + \delta y, t + \delta t) = I(x, y, t) + I_x \delta x + I_y \delta y + I_t \delta t$$
.....(2)

Subtract (1) from (2): 
$$I_x \delta x + I_y \delta y + I_t \delta t = 0$$

Divide by 
$$\delta t$$
 and take limit as  $\delta t \to 0$ :  $I_x \frac{\partial x}{\partial t} + I_y \frac{\partial y}{\partial t} + I_t = 0$ 

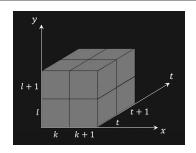
Constraint Equation: 
$$I_x u + I_y v + I_t = 0$$
 (*u, v*): Optical Flow

 $(I_x, I_y, I_t)$  can be easily computed from two frames





# Computing Partial Derivatives $I_x, I_y, I_t$



$$I_x(k,l,t)$$

$$= \frac{1}{4} [I(k+1,l,t) + I(k+1,l+1,t) + I(k+1,l,t+1) + I(k+1,l+1,t+1)]$$

$$-\frac{1}{4}[I(k,l,t) + I(k,l+1,t) + I(k,l,t+1) + I(k,l+1,t+1)]$$

Similarly find  $I_y(k, l, t)$  and  $I_t(k, l, t)$ 





## **Geometric interpretation**

For any point (x, y) in the image, its optical flow (u, v) lies on the line:

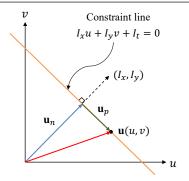
$$I_x u + I_y v + I_t = 0$$

Optical Flow can be split into two components.

$$u = u_n + u_n$$

 $\boldsymbol{u}_n$ : Normal Flow

 $u_n$ : Parallel Flow







### **Normal Flow**

#### Direction of Normal Flow:

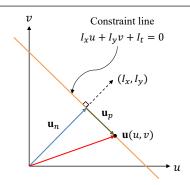
Unit vector perpendicular to the constraint line:

$$\widehat{\boldsymbol{u}}_n = \frac{(I_x, I_y)}{\sqrt{I_x^2 + I_y^2}}$$

#### Magnitude of Normal Flow:

Distance of origin from the constant line:

$$|\boldsymbol{u}_n| = \frac{|I_t|}{\sqrt{I_x^2 + I_y^2}}$$



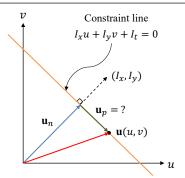
$$u_n = \frac{|I_t|}{(I_x^2 + I_y^2)} (I_x, I_y)$$





### Parallel Flow

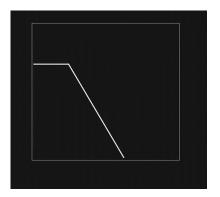
We can not determine  $u_p$ , the optical flow component parallel to the constraint line.







# **Aperture Problem**



Locally, we can only determine Normal Flow!





# **Optical Flow is Under constrained**

Constraint Equation: 
$$I_x u + I_y v + I_t = 0$$

2 unknowns, 1 equation.





### **Lucas-Kanada Solution**

Assumption: For each pixel, assume Motion Field, and hence Optical Flow (u, v), is constant within a small neighbourhood W.



That is for all points  $(k, l) \in W$ :

$$I_x(k,l)u + I_v(k,l)v + I_t(k,l) = 0$$





#### **Lucas-Kanada Solution**

For all points  $(k, l) \in W$ :  $I_x(k, l)u + I_y(k, l)v + I_t(k, l) = 0$ 

Let the size of window W be  $n \times n$ 

In matrix form:

$$\begin{bmatrix} I_{x}(1,1) & I_{y}(1,1) \\ \vdots & \vdots \\ I_{x}(k,l) & I_{y}(k,l) \\ \vdots & \vdots \\ I_{x}(n,n) & I_{y}(n,n) \end{bmatrix} \begin{bmatrix} \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} I_{t}(1,1) \\ \vdots \\ I_{t}(k,l) \\ \vdots \\ I_{t}(n,n) \end{bmatrix}$$

$$A \qquad \mathbf{u} \qquad B$$
(Known) (Unknown) (Known)
$$n^{2} \times 2 \qquad 2 \times 1 \qquad n^{2} \times 1$$

 $n^2$  Equations, 2 Unknowns: Find Least Squares Solution





#### When Dose Optical Flow Estimation Work?

$$Au = B$$

$$A^T A u = A^T B$$

- $A^T A$  must be invertible. That is  $det(A^T A) \neq 0$
- $A^TA$  must be well-conditioned.

If  $\lambda_1$  and  $\lambda_2$  are eigen values of  $A^TA$ , then

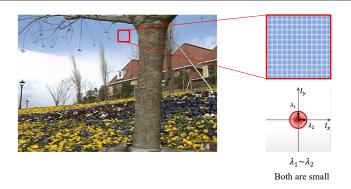
$$\lambda_1 > \epsilon$$
 and  $\lambda_2 > \epsilon$ 

$$\lambda_1 \ge \lambda_2$$
 but not  $\lambda_1 \gg \lambda_2$ 





### **Smooth Regions (Bad)**



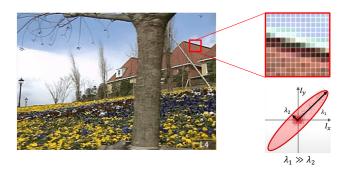
Equations for all pixels in window are both more or less the same

Cannot reliably compute flow!





# Edges (Bad)



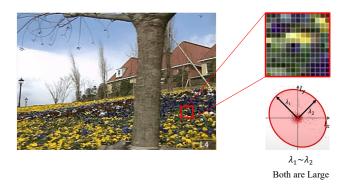
Badly conditioned. Prominent gradient in one direction.

Cannot reliably compute flow! Same as Aperture Problem.





# **Textured Regions (Good)**



Well conditioned. Large and diverse gradient magnitudes.

Can reliably compute optical flow!





# What if we have Large Motion?





Taylor Series approximation of

$$I(x + \delta x, y + \delta y, t + \delta t)$$
 is not valid

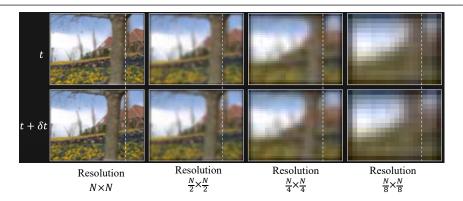
Our simple linear constraint equation not valid

$$I_x u + I_y v + I_t \neq 0$$





# **Large Motion: Coarse-to-Fine Estimation**

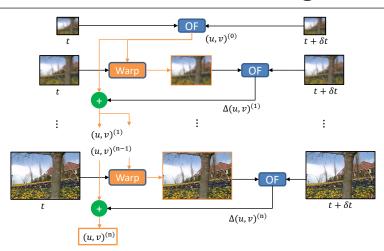


At lowest resolution, motion  $\leq 1$  pixel





## **Coarse-to-Fine Estimation Algorithm**



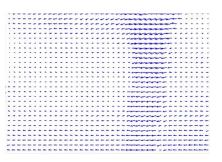




## **Results: Tree Sequence**







Optical Flow





### **Results: Rotating Ball**



Image Sequence



Optical Flow



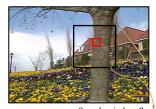


#### **Alternative Approach: Template Matching**

#### Determine Flow using Template Matching



Template window TImage  $I_1$  at time t



Search window S Image  $I_2$  at time  $t + \delta t$ 

For each template window T in image  $I_1$ , find the corresponding match in image  $I_2$ 





#### **Alternative Approach: Template Matching**

#### Determine Flow using Template Matching



Template window TImage  $I_1$  at time t



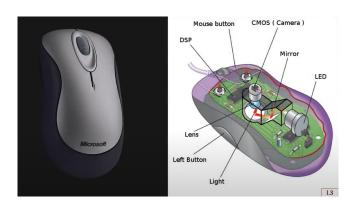
Image  $I_2$  at time  $t + \delta t$ 

- 1. Template Matching is slow when search window S is large.
- 2. Also mismatches are possible





# **Applications: Optical Mouse**



**Estimating Mouse Movements** 





# **Applications: Traffic Monitoring**



Finding Velocities of Vehicles





# **Applications: Video Retiming**





Optical Flow is used to determine the intermediate frames to produce slow-motion effect.







# **Struction from Motion**

张举勇 中国科学技术大学

# **Uncontrolled (Casual) Video**







#### **Overview**

Compute 3D scene structure and camera motion from a sequence of frames.

#### **Topics:**

- (1) Structure from Motion Problem
- (2) SFM Observation Matrix
- (3) Rank of Observation Matrix
- (4) Tomasi-Kanade Factorization





## **Feature Detection and Tracking**

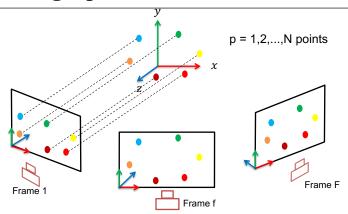
- Detect feature points: Corners, SIFT points, ...
- Track feature points: Template Matching, Optical Flow...







## **Orthographic Structure from Motion**

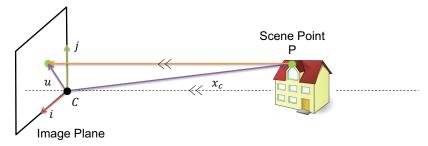


Given sets of corresponding image points (2D):  $(u_{f,p}, v_{f,p})$ Find scene points (3D)  $P_p$ , assuming orthographic camera.





# From 3D to 2D: Orthographic Projection



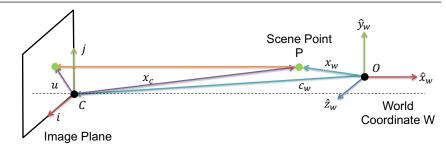
$$u = i \cdot x_c = i^T x_c$$
$$v = j \cdot x_c = j^T x_c$$

Perspective cameras exhibit orthographic projection when distance of scene from camera is large compared to depth variation within scene (magnification is nearly constant).





# From 3D to 2D: Orthographic Projection



$$u = i^{T} x_{c} = i^{T} (x_{w} - c_{w}) = i^{T} (P - C)$$

$$v = j^{T} x_{c} = j^{T} (x_{w} - c_{w}) = j^{T} (P - C)$$

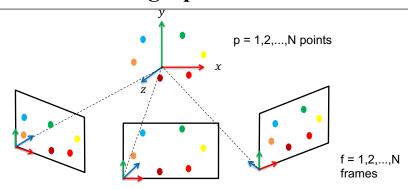
$$u = i^{T} (P - C)$$

$$v = j^{T} (P - C)$$





## **Orthographic SFM**



Given corresponding image points (2D)  $(u_{f,p}, v_{f,p})$ 

Find scene points  $\{P_p\}$ .

Camera Positions  $\{C_f\}$ , camera orientations  $\{(i_f, j_f)\}$  are unknown.





### **Orthographic SFM**

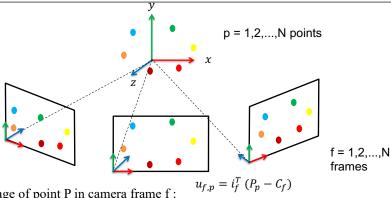


Image of point P in camera frame f:

$$v_{f,p} = j_f^T (P_p - C_f)$$

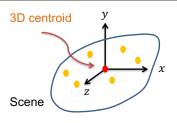
Known Unknown

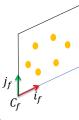
We can remove C from equations to simply SFM problem.





## **Centering Trick**





Frame f

Assume origin of world at centroid of scene points:

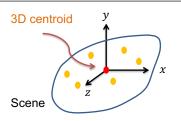
$$\frac{1}{N}\sum_{p=1}^{N}P_{p}=\bar{P}=0$$

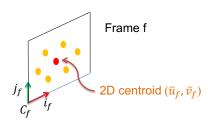
We will compute scene points w.r.t their centroid!





#### **Centering Trick**





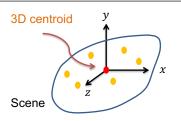
Centroid  $(\bar{u}_f, \bar{v}_f)$  of the image points in frame f:

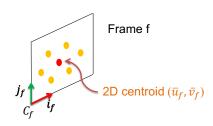
$$\begin{split} \bar{u}_f &= \frac{1}{N} \sum_{p=1}^N u_{f,p} = \frac{1}{N} \sum_{p=1}^N i_f^T \left( P_p - C_f \right) & \bar{v}_f &= \frac{1}{N} \sum_{p=1}^N v_{f,p} = \frac{1}{N} \sum_{p=1}^N j_f^T \left( P_p - C_f \right) \\ \bar{u}_f &= \frac{1}{N} i_f^T \sum_{p=1}^N P_p - \frac{1}{N} \sum_{p=1}^N i_f^T C_f & \bar{v}_f &= \frac{1}{N} j_f^T \sum_{p=1}^N P_p - \frac{1}{N} \sum_{p=1}^N j_f^T C_f \end{split}$$





#### **Centering Trick**





Shift camera origin to the centroid  $(\bar{u}_f, \bar{v}_f)$ .

Image points w.r.t.  $(\bar{u}_f, \bar{v}_f)$ :

$$\begin{split} \tilde{u}_{f,p} &= u_{f,p} - \bar{u}_f = i_f^T \left( P_p - C_f \right) + i_f^T C_f & \tilde{v}_{f,p} &= v_{f,p} - \bar{v}_f = j_f^T \left( P_p - C_f \right) + j_f^T C_f \\ \tilde{u}_{f,p} &= i_f^T P_p & \\ \end{split}$$

Camera locations  $C_f$  now removed from equations.



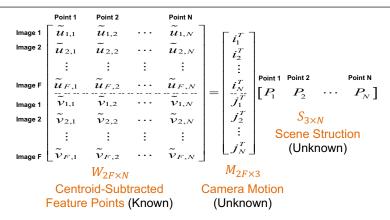


#### **Observation Matrix W**





#### **Observation Matrix W**



Can we find M and s from W?





# **Linear Independence of Vectors**

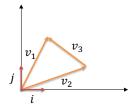
A set of vectors  $\{v_1, v_2, \dots v_n\}$  is said to be linearly independent if no vector can be represented as a weighted linear sum of the others.

 $\{i, j\}$  is linearly independent.

 $\{i, j, v_1\}$  is linearly dependent.

 $\{i, j, v_3\}$  is linearly dependent.

 $\{v_1, v_2, v_3\}$  is linearly dependent.







#### Rank of a Matrix

 Column Rank: The number of linearly independent columns of the matrix.

Row Rank: The number of linearly independent rows of the matrix.  $\begin{bmatrix} r^T \end{bmatrix}$ 

$$m\begin{bmatrix} & A & \\ & A & \\ & & n \end{bmatrix} = \begin{bmatrix} c_1 & c_2 & \cdots & c_n \end{bmatrix} = \begin{bmatrix} r_1^T \\ r_2^T \\ \vdots \\ r_n^T \end{bmatrix}$$

ColumnRank(A)≤n

 $ColumnRank(A) \leq m$ 

$$ColumnRank(A) = RowRank(A) = Rank(A)$$
  
 $Rank(A) \le min(m,n)$ 





# **Geometric Meaning of Matrix Rank**

Rank is the dimensionality of the space spanned by column or row vectors of the matrix.

$$A = \begin{bmatrix} a_1 & b_1 & c_1 \\ a_2 & b_2 & c_2 \\ a_3 & b_3 & c_3 \end{bmatrix} = \begin{bmatrix} \mathbf{a} & \mathbf{b} & \mathbf{c} \end{bmatrix}$$

$$Rank(\mathbf{A}) = 1$$





# **Geometric Meaning of Matrix Rank**

Rank is the dimensionality of the space spanned by column or row vectors of the matrix.

$$A = \begin{bmatrix} a_1 & b_1 & c_1 \\ a_2 & b_2 & c_2 \\ a_3 & b_3 & c_3 \end{bmatrix} = [\mathbf{a} \ \mathbf{b} \ \mathbf{c}]$$

$$Rank(\mathbf{A}) = 2$$





# **Geometric Meaning of Matrix Rank**

Rank is the dimensionality of the space spanned by column or row vectors of the matrix.

$$A = \begin{bmatrix} a_1 & b_1 & c_1 \\ a_2 & b_2 & c_2 \\ a_3 & b_3 & c_3 \end{bmatrix} = \begin{bmatrix} \boldsymbol{a} & \boldsymbol{b} & \boldsymbol{c} \end{bmatrix}$$

$$Rank(\mathbf{A}) = 3$$





# **Important Properties of Matrix Rank**

- $Rank(A^T) = Rank(A)$
- $Rank(A_{m \times n} | B_{n \times p}) = \min(Rank(A_{m \times n}), Rank(B_{n \times p}))$  $\leq \min(m, n, p)$
- $Rank(A A^T) = Rank(A^T A) = Rank(A) = Rank(A^T)$
- $A_{m \times m}$  is invertible iff  $Rank(A_{m \times m}) = m$





#### ...Back to Observation Matrix W





#### Rank of Observation Matrix

$$W = M \times S$$
$$2F \times N \quad 2F \times 3 \quad 3 \times N$$

#### We know:

$$Rank(MS) \le Rank(M)$$
  $Rank(MS) \le Rank(S)$ 

$$Rank(MS) \le min(3,2F)$$
  $Rank(MS) \le min(3,N)$ 

$$Rank(W) = Rank(MS) \le min(3, N, 2F)$$

Rank throem:  $Rank(W) \leq 3$ 

We can "factorize" W into M and S!





### **Singular Value Decomposition (SVD)**

For any matrix A there exists a factorization:

$$A_{M\times N} = U_{M\times M} \cdot \Sigma_{M\times N} \cdot V_{N\times N}^{T}$$

Where  $U_{M\times M}$  and  $V_{N\times N}^T$  are orthonormal and  $\Sigma_{M\times N}$  is diagonal.

$$Mathlab: [U,S,V] = svd(A)$$

$$\Sigma_{M\times N} = \begin{bmatrix} \sigma_1 & 0 & 0 & 0 & \cdots & 0 \\ 0 & \sigma_2 & 0 & 0 & \cdots & 0 \\ 0 & 0 & \sigma_3 & 0 & \cdots & 0 \\ 0 & 0 & 0 & \sigma_4 & \cdots & 0 \\ 0 & 0 & 0 & 0 & \ddots & 0 \\ 0 & 0 & 0 & 0 & \cdots & \sigma_N \\ 0 & 0 & 0 & 0 & \cdots & 0 \\ \vdots & \vdots & \vdots & \vdots & \cdots & \vdots \end{bmatrix} \qquad \sigma_1, ..., \sigma_N : Singular Values$$

If Rank(A) = r then A has r non-zero singular values.





#### Using SVD:

$$W = U \Sigma V^{T}$$

$$= \begin{bmatrix} U & \begin{bmatrix} \sigma_{1} & 0 & 0 & 0 & \cdots & 0 \\ 0 & \sigma_{2} & 0 & 0 & \cdots & 0 \\ 0 & 0 & \sigma_{3} & 0 & \cdots & 0 \\ 0 & 0 & 0 & \sigma_{4} & \cdots & 0 \\ 0 & 0 & 0 & 0 & \ddots & 0 \\ 0 & 0 & 0 & 0 & \cdots & \sigma_{N} \\ 0 & 0 & 0 & 0 & \cdots & \sigma_{N} \\ \vdots & \vdots & \vdots & \vdots & \vdots & \cdots & \vdots \end{bmatrix}$$

$$V^{T}$$

$$V^{$$

Where:  $\sigma_1 \ge \sigma_2 \ge ... \ge \sigma_N$  are the singular values of  $\Sigma$ .





#### Using SVD:

$$W = U\Sigma V^{T}$$

$$= U$$

 $2.F \times 2.F$ 

$$\begin{bmatrix} \sigma_1 & 0 & 0 & 0 & \cdots & 0 \\ 0 & \sigma_2 & 0 & 0 & \cdots & 0 \\ 0 & 0 & \sigma_3 & 0 & \cdots & 0 \\ 0 & 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & 0 & \cdots & 0 \\ 0 & 0 & 0 & 0 & \cdots & 0 \\ \vdots & \vdots & \vdots & \vdots & \cdots & \vdots \end{bmatrix}$$

,

 $N \times N$ 

Since  $Rank(W) \leq 3$ ,  $Rank(\Sigma) \leq 3$ .

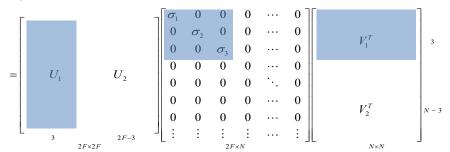
All expect first 3 diagonal elements of  $\Sigma$  must be 0.





#### Using SVD:

$$W = U\Sigma V^T$$



Since  $Rank(W) \leq 3$ ,  $Rank(\Sigma) \leq 3$ .

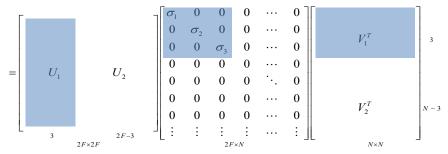
Submatrices  $U_2$  and  $V_2^T$  do not contribute to W.





#### Using SVD:

$$W = U\Sigma V^T$$



$$W = U_1 \sum_1 V_1^T$$

$$(2F \times 3)(3 \times 3)(3 \times P)$$





### Factorization (Finding M, S)

$$W = U_{1} (\Sigma_{1})^{\frac{1}{2}} (\Sigma_{1})^{\frac{1}{2}} V_{1}^{T}$$

$$(2F \times 3) \qquad (3 \times N)$$

$$= M? \qquad = S?$$

Not so fast. Decomposition not unique!

For any 3X3 non-singular matrix Q:

$$W = U_1 (\Sigma_1)^{\frac{1}{2}} Q Q^{-1} (\Sigma_1)^{\frac{1}{2}} V_1^T \text{ is also valid.}$$

$$(2F \times 3) \qquad (3 \times N)$$

$$= M \qquad = S \text{ for some } Q$$

How to find the matrix Q?





### Orthonormality of M

The Motion Matrix M:

$$M = \begin{bmatrix} \boldsymbol{i}_1^T \\ \vdots \\ \boldsymbol{i}_F^T \\ \boldsymbol{j}_1^T \\ \vdots \\ \boldsymbol{j}_F^T \end{bmatrix} = U_1(\boldsymbol{\Sigma}_1)^{1/2} Q = \begin{bmatrix} \boldsymbol{i}_1^T \\ \vdots \\ \boldsymbol{i}_F^T \\ \boldsymbol{j}_1^T \\ \vdots \\ \boldsymbol{j}_F^T \end{bmatrix} Q = \begin{bmatrix} \boldsymbol{i}_1^T Q \\ \vdots \\ \boldsymbol{i}_F^T Q \\ \vdots \\ \boldsymbol{i}_F^T Q \\ \vdots \\ \boldsymbol{j}_F^T Q \end{bmatrix}$$

**Orthonormality Constraints:** 

$$i_f \cdot i_f = i_f^T i_f = 1$$

$$j_f \cdot j_f = j_f^T j_f = 1$$

$$i_f \cdot j_f = i_f^T j_f = 0$$

$$\hat{i}_f^T Q Q^T \hat{i}_f = 1$$

$$\hat{i}_f^T Q Q^T \hat{j}_f = 1$$

$$\hat{i}_f^T Q Q^T \hat{j}_f = 0$$



$$\hat{i}_f^T Q Q^T \hat{i}_f = 1$$
$$\hat{j}_f^T Q Q^T \hat{j}_f = 1$$
$$\hat{i}_f^T Q Q^T \hat{j}_f = 0$$





### **Orthonormality of M**

• We have computed  $(\hat{i}_f^T, \hat{j}_f^T)$  for f = 1,...,F.

$$\hat{\imath}_f^T Q Q^T \hat{\imath}_f = 1$$
 
$$\hat{\jmath}_f^T Q Q^T \hat{\jmath}_f = 1$$
 Q is unknown. 
$$\hat{\imath}_f^T Q Q^T \hat{\jmath}_f = 0$$

- Q is 3x3 matrix, 9 variables, 3F quadratic equations.
- Q can be solved with 3 or more images (F≥ 3) using Newton's method.

Final Solution:

$$M = U_1 \left( \Sigma_1 \right)^{\frac{1}{2}} Q$$

Camera Motion

$$S = Q^{-1} (\Sigma_1)^{\frac{1}{2}} V_1^T$$

Scene struction





## **Summary: Orthographic SFM**

- 1. Detect and track feature points.
- 2. Create the centroid subtracted matrix w of corresponding feature points.
- 3. Compute SVD of W and enforce rank constraint.

$$W = U \Sigma V^{T} = U_{1} \Sigma_{1} V_{1}^{T}$$

$$(2F \times 3)(3 \times 3)(3 \times P)$$

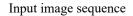
- 4. Set  $M = U_1(\Sigma_1)^{\frac{1}{2}}Q$  and  $S = Q^{-1}(\Sigma_1)^{\frac{1}{2}}V_1^T$ .
- 5. Find Q by enforcing the orthonormality constraint.





#### Result





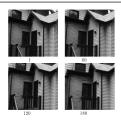


Estimated 3D points





#### Result



Input image sequence



3D reconstruction



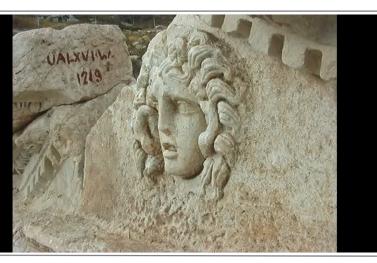
Tracked features



3D reconstruction



#### **Structure from Motion Result**





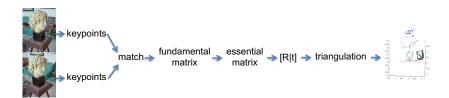




# **Bundle Adjustment**

张举勇 中国科学技术大学

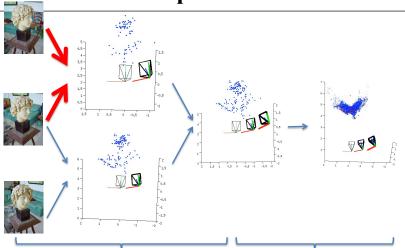
#### **Two-view Reconstruction**







# **Pipeline**



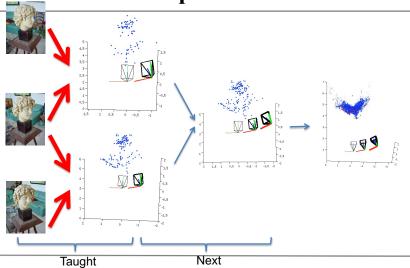
Structure from Motion (SFM)

Multi-view Stereo (MVS)

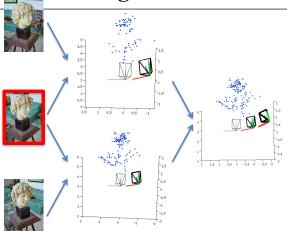




# **Pipeline**

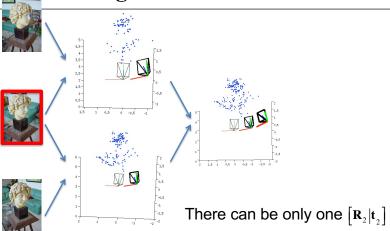
















• From the 1<sup>st</sup> and 2<sup>nd</sup> images, we have

$$\begin{bmatrix} \mathbf{R}_1 | \mathbf{t}_1 \end{bmatrix}$$
 and  $\begin{bmatrix} \mathbf{R}_2 | \mathbf{t}_2 \end{bmatrix}$ 

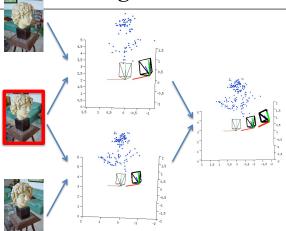
• From the 2<sup>nd</sup> and 3<sup>rd</sup> images, we have

$$\left[\mathbf{R}_{2}|\mathbf{t}_{2}\right]$$
 and  $\left[\mathbf{R}_{3}|\mathbf{t}_{3}\right]$ 

• How to transform the coordinate system of the second point cloud to align with the first point cloud so that there is only one  $[\mathbf{R}_2|\mathbf{t}_2]$ ?



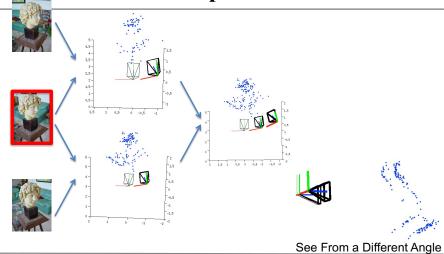








# **Oops**



# **Bundle Adjustment**











### **Rethinking the SFM problem**

• Input: Observed 2D image position

• Output:

Unknown Camera Parameters (with some guess) 
$$[\mathbf{R}_1|\mathbf{t}_1], [\mathbf{R}_2|\mathbf{t}_2], [\mathbf{R}_3|\mathbf{t}_3]$$

Unknown Point 3D coordinate (with some guess)

$$\mathbf{X}^1, \mathbf{X}^2, \mathbf{X}^3, \cdots$$





#### **Bundle Adjustment**

A valid solution  $[\mathbf{R}_1|\mathbf{t}_1]$ ,  $[\mathbf{R}_2|\mathbf{t}_2]$ ,  $[\mathbf{R}_3|\mathbf{t}_3]$  and  $[\mathbf{X}^1,\mathbf{X}^2,\mathbf{X}^3,...]$  must let

$$\begin{aligned} \text{Re-projection-} \begin{bmatrix} & \mathbf{x}_1^1 = \mathbf{K} \big[ \mathbf{R}_1 \big| \mathbf{t}_1 \big] \mathbf{X}^1 & \mathbf{x}_1^2 = \mathbf{K} \big[ \mathbf{R}_1 \big| \mathbf{t}_1 \big] \mathbf{X}^2 \\ & \mathbf{x}_2^1 = \mathbf{K} \big[ \mathbf{R}_2 \big| \mathbf{t}_2 \big] \mathbf{X}^1 & \mathbf{x}_2^2 = \mathbf{K} \big[ \mathbf{R}_2 \big| \mathbf{t}_2 \big] \mathbf{X}^2 & \mathbf{x}_2^3 = \mathbf{K} \big[ \mathbf{R}_2 \big| \mathbf{t}_2 \big] \mathbf{X}^3 \\ & \mathbf{x}_3^1 = \mathbf{K} \big[ \mathbf{R}_3 \big| \mathbf{t}_3 \big] \mathbf{X}^1 & \mathbf{x}_3^3 = \mathbf{K} \big[ \mathbf{R}_3 \big| \mathbf{t}_3 \big] \mathbf{X}^3 \end{aligned}$$

Observation 
$$\begin{bmatrix} \tilde{\mathbf{x}}_1^1 & \tilde{\mathbf{x}}_1^2 \\ \tilde{\mathbf{x}}_2^1 & \tilde{\mathbf{x}}_2^2 & \tilde{\mathbf{x}}_2^3 \\ \tilde{\mathbf{x}}_3^1 & \tilde{\mathbf{x}}_3^3 \end{bmatrix}$$





#### **Bundle Adjustment**

A valid solution  $[\mathbf{R}_1|\mathbf{t}_1], [\mathbf{R}_2|\mathbf{t}_2], [\mathbf{R}_3|\mathbf{t}_3]$  and  $\mathbf{X}^1, \mathbf{X}^2, \mathbf{X}^3, \cdots$  must let the Re-projection close to the Observation, i.e. to minimize the reprojection error

$$\min \sum_{i} \sum_{j} \left( \tilde{\mathbf{x}}_{i}^{j} - \mathbf{K} \left[ \mathbf{R}_{i} \middle| \mathbf{t}_{i} \right] \mathbf{X}^{j} \right)^{2}$$





#### **Solving This Optimization Problem**

• Theory:

The Levenberg-Marquardt algorithm

http://en.wikipedia.org/wiki/Levenberg-Marquardt algorithm

• Practice:

The Ceres-Solver from Google

http://code.google.com/p/ceres-solver/





#### **Ceres-solver: A Nonlinear Least Squares Minimizer**

## Toy problem to solve $\min(10-x)^2$

```
class SimpleCostFunction
  : public ceres::SizedCostFunction<1 /* number of residuals */,
                                    1 /* size of first parameter */> {
public:
 virtual ~SimpleCostFunction() {}
 virtual bool Evaluate(double const* const* parameters,
                        double* residuals.
                        double** jacobians) const {
   const double x = parameters[0][0];
   residuals[0] = 10 - x; // f(x) = 10 - x
   // Compute the Jacobian if asked for.
   if (jacobians != NULL && jacobians[0] != NULL) {
      iacobians[0][0] = -1;
   return true:
};
```





#### **Ceres-solver: A Nonlinear Least Squares Minimizer**

## Toy problem to solve $\min(10-x)^2$

```
int main(int argc, char** argv) {
  double x = 5.0:
  ceres::Problem problem;
  // The problem object takes ownership of the newly allocated
  // SimpleCostFunction and uses it to optimize the value of x.
  problem.AddResidualBlock(new SimpleCostFunction, NULL, &x);
  // Run the solver!
  Solver::Options options;
  options.max_num_iterations = 10;
  options.linear_solver_type = ceres::DENSE_QR;
  options.minimizer_progress_to_stdout = true;
  Solver::Summary summary;
  Solve(options, &problem, &summary);
  std::cout << summary.BriefReport() << "\n";
  std::cout << "x : 5.0 -> " << x << "\n";
  return 0:
```





#### **Ceres-solver: A Nonlinear Least Squares Minimizer**

Toy problem to solve  $\min(10-x)^2$ 

```
0: f: 1.250000e+01 d: 0.00e+00 g: 5.00e+00 h: 0.00e+00 rho: 0.00e+00 mu: 1.00e-04 li: 0
1: f: 1.249750e-07 d: 1.25e+01 g: 5.00e-04 h: 5.00e+00 rho: 1.00e+00 mu: 3.33e-05 li: 1
2: f: 1.388518e-16 d: 1.25e-07 g: 1.67e-08 h: 5.00e-04 rho: 1.00e+00 mu: 1.11e-05 li: 1
Ceres Solver Report: Iterations: 2, Initial cost: 1.250000e+01, \
Final cost: 1.388518e-16, Termination: PARAMETER_TOLERANCE.
x: 5 -> 10
```



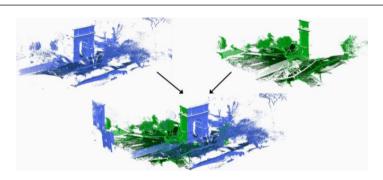




# **Geometry Registration**

张举勇 中国科学技术大学

## 什么是注册?



- 计算最佳空间变换,以使得多个几何曲面之间进行对齐。
  - 将传感器采集的多个局部测量数据拼接成一个完整的几何模型
  - 将新测量数据对齐到已知模型以估计其姿态





## 变换类型

- Same object in a different position: size and shape preserving
  - Rigid-body transformation (rotation and translation)
  - Six degrees of freedom

$$ightharpoonup$$
 translation  $\mathbf{t} = (t_x, t_y, t_z)^T$ 

$$\rightarrow$$
 rotation  $(\alpha, \beta, \gamma)$ 

$$\mathbf{T}_{\mathrm{rigid}}(\mathbf{x}) = \mathbf{R}\mathbf{x} + \mathbf{t}$$

$$\mathbf{T}_{\mathrm{rigid}}(\mathbf{x}) = \begin{bmatrix} \cos\beta\cos\gamma & \cos\alpha\sin\gamma + \sin\alpha\sin\beta\cos\gamma & \sin\alpha\sin\gamma - \cos\alpha\sin\beta\cos\gamma & t_x \\ -\cos\beta\sin\gamma & \cos\alpha\cos\gamma - \sin\alpha\sin\beta\sin\gamma & \sin\alpha\cos\gamma + \cos\alpha\sin\beta\sin\gamma & t_y \\ \sin\beta & -\sin\alpha\cos\beta & \cos\alpha\cos\beta & t_z \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$





## 变换类型

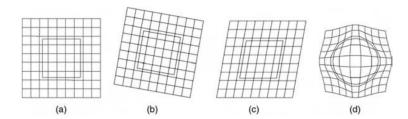
- Affine or Linear Transformation
  - Rigid-body transformation (rotation and translation)
  - Scaling and Shearing
  - Twelve degrees of freedom

$$\mathbf{T}(\mathbf{x}) = \begin{bmatrix} a_{00} & a_{01} & a_{02} & a_{03} \\ a_{10} & a_{11} & a_{12} & a_{13} \\ a_{20} & a_{21} & a_{22} & a_{23} \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$$





## 变换类型



Example of different types of transformations of a square

(a) identity transformation

(c) affine transformation

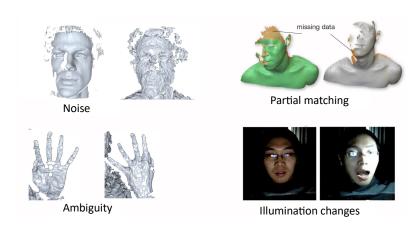
(b) rigid transformation

(d) nonrigid transformation





# 配准问题中的一些挑战







## 配准问题建模

• 将配准问题表达为能量最小化问题:

$$rgmax E_{reg}(T,P,Q)$$
 $T$ 
 $E_{reg}(T,P,Q)=E_{match}(T,P,Q)+E_{prior}(T)$ 
配准误差 变换误差 如何衡量配准结果的质量? 变换的类型与表示方式?

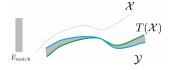




## 配准问题建模

• 配准误差

$$E_{reg}(T, P, Q) = E_{match}(T, P, Q) + E_{prior}(T)$$
 $E_{match}(T, P, Q) = \int_{X} \phi(T(p), Q) dx$ 
 $E_{match}(T, P, Q) = \sum_{X} \phi(T(p), Q) dx$ 





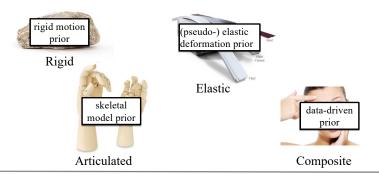




# 配准问题建模

• 变换误差

$$E_{rea}(T, P, Q) = E_{match}(T, P, Q) + E_{prior}(T)$$



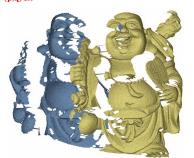




# 几何数据融合与跟踪-刚性注册

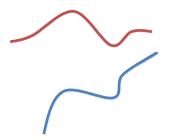
• 刚体几何建模:将不同视角点云进行刚性拼接,以获得完整几何模型

$$E(\mathbf{T}) = \sum_{(\mathbf{p},\mathbf{q}) \in K} \parallel \mathbf{p} - \mathbf{T}\mathbf{q} \parallel^2$$
  $\mathbf{T}$  是一个包含旋转与平移的刚性变换













### **Corresponding Point Set Alignment**

- Let M be a model point set.
- Let S be a scene point set.

#### We assume:

- 1.  $N_M = N_S$ .
- 2. Each point S<sub>i</sub> correspond to

M (model)





### **Corresponding Point Set Alignment**

The MSE objective function:

$$f(R,T) = \frac{1}{N_S} \sum_{i=1}^{N_S} ||m_i - Rot(s_i) - Trans||^2$$
$$f(q) = \frac{1}{N_S} \sum_{i=1}^{N_S} ||m_i - R(q_R)s_i - q_T||^2$$

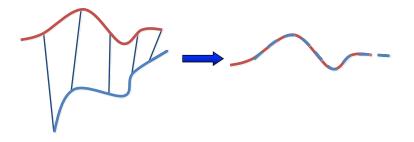
The alignment is:

$$(rot, trans, d_{mse}) = \Phi(M, S)$$





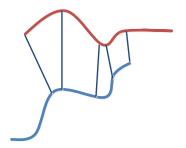
• If correct correspondences are known, can find correct relative rotation/translation







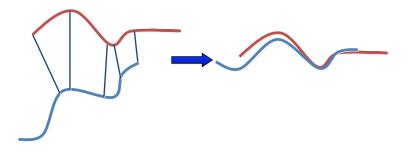
- How to find correspondences: User input? Feature detection?
   Signatures?
- Alternative: assume closest points correspond







- How to find correspondences: User input? Feature detection?
   Signatures?
- Alternative: assume closest points correspond







• Converges if starting position "close enough"







### **Closest Point**

• Given 2 points  $r_1$  and  $r_2$ , the Euclidean distance is:

$$d(r_1, r_2) = ||r_1 - r_2|| = \sqrt{(x_1 - x_2)^2 + (y_1 - y_2)^2 + (z_1 - z_2)^2}$$

• Given a point r<sub>1</sub> and set of points A, the Euclidean distance is:

$$d(r_1, A) = \min_{i \in 1..n} d(r_1, a_i)$$





### **Finding Matches**

- The scene shape S is aligned to be in the best alignment with the model shape M.
- The distance of each point s of the scene from the model is:

$$d(s,M) = \min_{m \in M} d \|m - s\|$$





### **Finding Matches**

$$d(s,M) = \min_{m \in M} d||m - s|| = d(s,y)$$
$$y \in M$$
$$Y = C(S,M)$$
$$Y \subseteq M$$

C- the closest point operator

Y - the set of closest points to S





### **Finding Matches**

- Finding each match is performed in O(N<sub>M</sub>) worst case.
- Given Y we can calculate alignment

$$(rot, trans, d) = \Phi(S, Y)$$

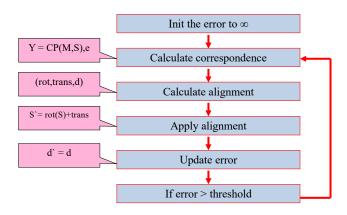
• S is updated to be:

$$S_{new} = rot(S) + trans$$





### The Algorithm







• The ICP algorithm always converges monotonically to a local minimum with respect to the MSE distance objective function.





• Correspondence error :

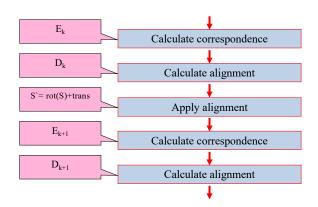
$$e_k = \frac{1}{N_S} \sum_{i=1}^{N_S} ||y_{ik} - s_{ik}||^2$$

Alignment error:

$$d_{k} = \frac{1}{N_{S}} \sum_{i=1}^{N_{S}} \|y_{ik} - Rot_{k}(s_{io}) - Trans_{k}\|^{2}$$











#### • Proof:

$$\begin{split} S_{k} &= Rot_{k}(S_{0}) + Trans_{k} \\ Y_{k} &= C(M, s_{k}) \\ e_{k} &= \frac{1}{N_{S}} \sum_{i=1}^{N_{S}} \left\| y_{ik} - s_{ik} \right\|^{2} \\ d_{k} &= \frac{1}{N_{S}} \sum_{i=1}^{N_{S}} \left\| y_{ik} - Rot_{k}(s_{io}) - Trans_{k} \right\|^{2} \end{split}$$





• Proof:  $d_k \leq e_k$ 

If not - the identity transform would yield a smaller MSE than the least square alignment.

Apply the alignment  $q_k$  on  $S_0 \rightarrow S_{k+1}$ .

Assuming the correspondences are maintained : the MSE is still  $d_k$ .

$$d_{k} = \frac{1}{N_{M}} \sum_{i=1}^{N_{M}} \left\| y_{ik} - S_{ik} \right\|^{2}$$





#### • Proof:

After the last alignment, the closest point operator is applied:  $Y_{k+1} = C(M, S_{k+1})$ It is clear that:

$$\begin{aligned} & \left\| y_{i,k+1} - S_{i,k+1} \right\| \leq \left\| y_{ik} - S_{i,k+1} \right\| \\ & e_{k+1} \leq d_k \end{aligned}$$

Thus: 
$$0 \le d_{k+1} \le e_{k+1} \le d_k \le e_k$$





### Time analysis

Each iteration includes 3 main steps

A. Finding the closest points :

O(N<sub>M</sub>) per each point

 $O(N_M*N_S)$  total.

B. Calculating the alignment:  $O(N_S)$ 

C. Updating the scene:  $O(N_S)$ 





## **Optimizing the Algorithm**

The best match/nearest neighbor problem:

Given a record, and a dissimilarity measure **D**, find the closest record from a set to the query record.



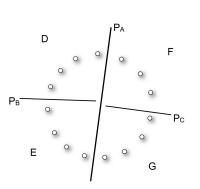


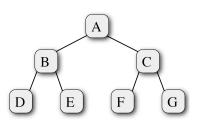
### **Closest Point Search**

- Find closest point of a query point
  - Brute force: O(n) complexity
- Use hierarchical BSP tree
  - Binary space partitioning tree (also kD-tree)
  - Recursively partition 3D space by planes
  - Tree should be balanced, put plane at median
  - $-\log(n)$  tree levels, complexity  $O(\log n)$









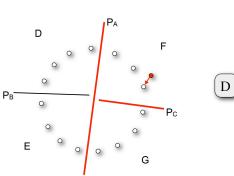


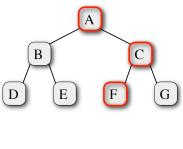


```
BSPNode::dist(Point x, Scalar& dmin)
  if (leaf node())
    for each sample point p[i]
      dmin = min(dmin, dist(x, p[i]));
  else
   d = dist to plane(x);
    if (d < 0)
      left child->dist(x, dmin);
      if (|d| < dmin) right child->dist(x, dmin);
    else
      right child->dist(x, dmin);
      if (|d| < dmin) left child->dist(x, dmin);
```



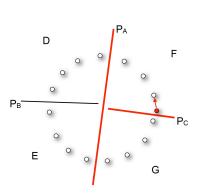


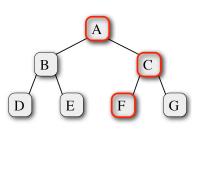






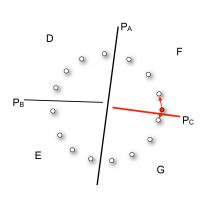


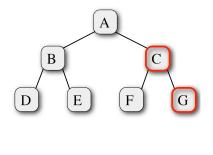






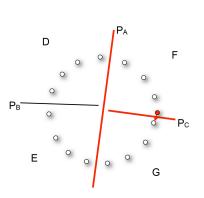


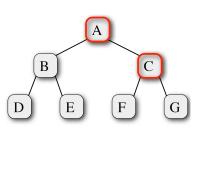
















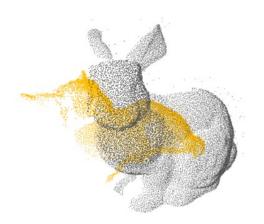
### **ICP Variants**

- Variants on the following stages of ICP have been proposed:
  - Selecting sample points (from one or both meshes)
  - Matching to points in the other mesh
  - Weighting the correspondences
  - Rejecting certain (outlier) point pairs
  - Assigning an error metric to the current transform
  - Minimizing the error metric w.r.t. transformation





### **Real Time ICP**









# **Object Tracking**

张举勇 中国科学技术大学

### **Overview**

Track the location of target objects in each frame of a video sequence

#### **Topics:**

- (1) Change Detection
- (2) Gaussian Mixture Model
- (3) Object Tracking using Templates
- (4) Tracking by Feature Detection





### **Change Detection**

Given: Static cameras observing scene (room, street, etc.) Find: Meaningful changes (moving objects, people, etc.)



Robust and real-time classification of each pixel as "foreground" (motion/change) or "background" (static).





### **Change Detection: Challenges**

#### Ignore uninteresting changes:

- Background fluctuations
- Image noise
- Rain, snow, turbulence
- Illumination changes & shadows
- Camera shake







### **Simple Frame Difference**

Label significant difference between current and previous frames as background.

$$F_t = |I_t - I_{t-1}| > T$$

*T*: threshold



Input video sequence

Frame difference

Not Robust!





## **Background Modeling: Average**

Build simple model of background before classification.



Background B median{I<sub>1</sub>, I<sub>2</sub>,..., I<sub>K</sub>} (First K frames)

Input Frame  $I_t$ 

Foreground  $F_t$  $F_t = |I_t - B| > T$ 

Cannot handle change in lighting, background, etc.





## **Background Modeling: Median**

Build simple model of background before classification.



Background 
$$B_t$$
 median $\{I_{t-1}, I_{t-2}, ..., I_{t-K}\}$  (Last K frames)

Input Frame  $I_t$ 

Foreground 
$$F_t$$
  
 $F_t = |I_t - \mathbf{B}| > T$ 

Cannot handle change in lighting, background, etc.





## **Background Modeling: Moving Median**

Build simple adaptive model of background over time.



Background  $B_t$  median $\{I_{t-1}, I_{t-2}, ..., I_{t-K}\}$  (Last K frames)

Input Frame I<sub>t</sub>

Foreground  $F_t$  $F_t = |I_t - B| > T$ 

Requires keeping the last K frames in memory. Finding median for each pixel is expensive.





# **Background Modeling: Moving Median**

Build simple adaptive model of background over time.



Background  $B_t$  median $\{I_{t-1}, I_{t-2}, ..., I_{t-K}\}$  (Last K frames)

Input Frame I<sub>t</sub>

Foreground  $F_t$  $F_t = |I_t - \mathbf{B}| > T$ 

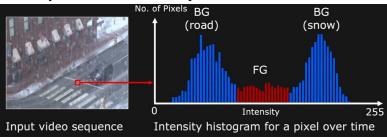
Cannot handle significant pixel fluctuations (weather shadow, shake, etc.)





#### Mixture Model

Intensity distribution at each pixel over time:



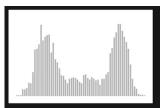
Intensity variations due to static scene (road), noise (snow), and occasional moving objects(vehicles)

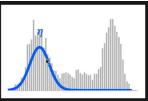
Intuition: Pixels are background most of time.





#### **Gaussian Model**





Probability Distribution P(x) (x: pixel intensity)

Gaussian  $\omega, \eta(x, \mu, \sigma)$ 

1-Dimensional Gaussian:

$$\omega \, \eta(x,\mu,\sigma) = \omega \frac{1}{\sqrt{2\pi}\sigma} \, e^{-\frac{1}{2\sigma^2}(x-\mu)^2}$$

μ: Mean

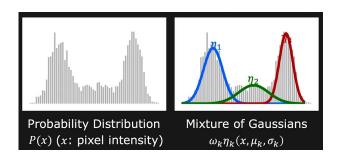
 $\sigma$ : Std. Deviation

 $\omega$ : Scale





#### **Mixture of Gaussians**

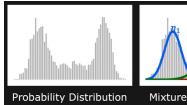


Assume P(x) is made of K different Gaussians.



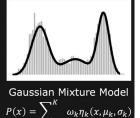


#### **Gaussian Mixture Model (GMM)**



P(x) (x: pixel intensity)

Mixture of Gaussians  $\omega_k \eta_k(x, \mu_k, \sigma_k)$ 



GMM Distribution: Weighted sum of K Gaussians

$$P(x) \cong \sum_{k=1}^K \omega_k \eta_k(x, \mu_k, \sigma_k) \qquad \text{such that } \sum_{k=1}^K \omega_k = 1$$

such that 
$$\sum_{k=1}^{K} \omega_k = 1$$





#### **High Dimensional Model**

Let  $P(\mathbf{X})$  be a probability distribution of a D-dimensional random variable  $\mathbf{X} \in \mathcal{R}^D$ . For example:  $\mathbf{X} = [r, g, b]^T$ 

GMM of P(X): Sum of K D-dimensional Gaussians

$$P(\mathbf{X}) \cong \sum_{k=1}^K \omega_k \eta_k(\mathbf{X}, \boldsymbol{\mu}_k, \Sigma_k)$$
 such that  $\sum_{k=1}^K \omega_k = 1$ 

where: 
$$\eta(\mathbf{X}, \boldsymbol{\mu}, \boldsymbol{\Sigma}) = \frac{1}{(2\pi)^{D/2} |\boldsymbol{\Sigma}|^{1/2}} e^{-\frac{1}{2} (\mathbf{X} - \boldsymbol{\mu})^T (\boldsymbol{\Sigma})^{-1} (\mathbf{X} - \boldsymbol{\mu})}$$

Mean 
$$\boldsymbol{\mu} = \begin{bmatrix} \mu_r \\ \mu_g \\ \mu_b \end{bmatrix}$$
 Covariance matrix  $\boldsymbol{\Sigma} = \begin{bmatrix} \sigma^2 & 0 & 0 \\ 0 & \sigma^2 & 0 \\ 0 & 0 & \sigma^2 \end{bmatrix}$  (can be a full matrix)

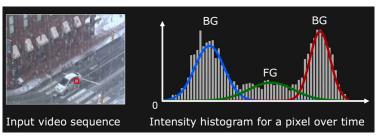
GMM can be estimated from P(X). (MATLAB: gmdistribution.fit)





#### **Background Modeling with GMM**

Given: A GMM for intensity/color variation at a pixel over time Classify: Individual Gaussians as foreground/background



Intuition: Pixels are background most of time. That is, Gaussians with large supporting evidence  $\omega$  and small  $\sigma$ .

Large 
$$\frac{\omega}{\sigma}$$
: Background

Small  $\frac{\omega}{\sigma}$ : Foreground





# **Change Detection using GMM**

#### For each pixel:

- 1. Compute pixel color histogram H using first N frames.
- 2. Normalize histogram:  $\hat{H} \leftarrow H/\parallel H \parallel$ .
- 3. Model  $\hat{H}$  as mixture of K (3 to 5) Gaussians.
- 4. For each subsequent frame:
  - a. The pixel value  $\mathbf{X}$  belongs to Gaussian k in GMM for which
  - $\|\mathbf{X} \boldsymbol{\mu}_k\|$  is minimum and  $\|\mathbf{X} \boldsymbol{\mu}_k\| < 2.5\sigma_k$
  - b. If  ${}^{\omega_k}/\sigma_k$  is large then classify pixel as background. Else classify as foreground.
  - c. Update histogram *H* using new pixel intensity.
  - d. If  $\widehat{H}$  and  $H/\parallel H \parallel$  differ a lot  $(\|\widehat{H} H/\| H \|\|$  is large),  $\widehat{H} \leftarrow H/\| H \|$  and refit GMM.





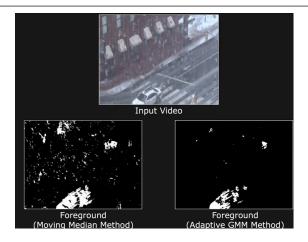
#### Adaptive GMM based change detection







### Adaptive GMM based change detection







# **Object Tracking**

Given: Location of target in initial or previous frame.

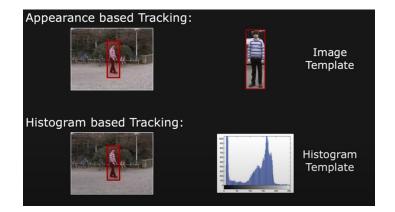
Find: Location of target in current frame.







### **Target templates for Tracking**







# **Tracking using Appearance Matching**



Given template window S in frame  $I_{t-1}$ , search neighborhood to find match in image  $I_t$ .

Simple implementation. Not robust to change in scale, viewpoint, Occlusion, etc.





# **Similarity Metrics for Template Matching**

Find pixel  $(k, l) \in S$  with Minimum Sum of Absolute Differences:

$$SAD(k \cdot l) = \sum_{(i \cdot j) \in T} |I_1(i \cdot j) - I_2(i \cdot + kj + l)|$$

Find pixel  $(k, l) \in S$  with Minimum Sum of Squared Differences:

$$SSD(k, l) = \sum_{(i,j) \in T} |I_1(i,j) - I_2(i+k, j+l)|^2$$

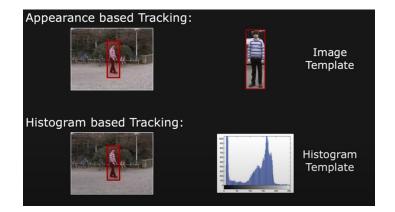
Find pixel  $(k, l) \in S$  with Minimum Normalized Cross-Correlation:

$$NCC(k,l) = \frac{\sum_{(i,j) \in T} I_1(i,j) I_2(i+k,j+l)}{\sqrt{\sum_{(i,j) \in T} I_1(i,j)^2 \sum_{(i,j) \in T} I_2(i+k,j+l)^2}}$$





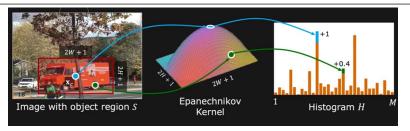
### **Target templates for Tracking**







#### **Computing Weighted Histogram**



Weighted histogram gives more importance to pixels at center.

**Epanechnikov** Kernel:

$$k(\tilde{\mathbf{x}}) = \begin{cases} 1 - \|\tilde{\mathbf{x}}\|^2, & \|\tilde{\mathbf{X}}\| < 1\\ 0, & \text{otherwise} \end{cases} \tilde{\mathbf{x}} = \begin{bmatrix} (x - x_c)/W\\ (y - y_c)/H \end{bmatrix}$$

Comparing Histograms: Correlation, Intersection, etc.





### Tracking using Histogram Matching



Given a histogram template  $H_0$  and location  $x_{t-1}$  in  $I_{t-1}$ , search neighborhood in  $I_t$  to find window in matching histogram.

More resilient to changes in object pose and/or scale





#### **Histogram Based Tracking: Results**



Robust when object appearance is unique in the environment and its size remains more or less the same.

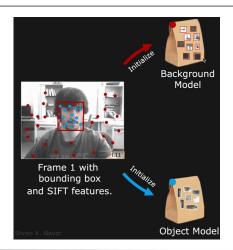






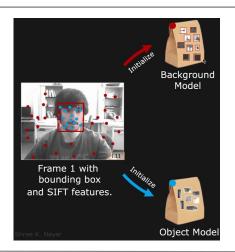








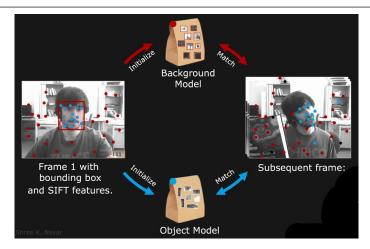






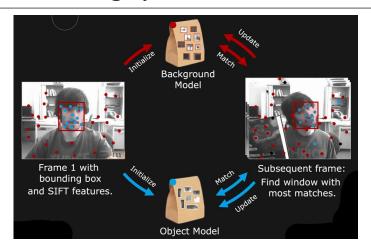
















#### **Tracking Intialization**

#### At frame 1:

- 1. User selects a bounding box  $W_1$  as object/target.
- 2. Compute SIFT (or similar) features for the frame.
- 3. Classify features within the box as object and assign them to set  $O_1$ .
- Classify remaining features as background and assign them to set B.





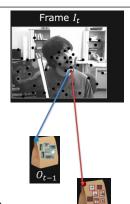


## **Object Tracking**

#### At frame t:

- 1. Compute SIFT features and SIFT descriptors  $\{\mathbf{v}_1, \dots, \mathbf{v}_K\}$  for frame  $I_t$ .
- 2. For each feature and corresponding descriptor  $\mathbf{v}_i$ :
  - a. Compute distance  $d_o$  between  $\mathbf{v}_i$  and the best match in object set  $O_{t-1}$
  - b. Compute distance  $d_B$  between  $\mathbf{v}_i$  and the best match in background set B.

c. 
$$C(\mathbf{v}_i) = \begin{cases} +1 & \text{if } d_0/d_B < 0.5(\mathbf{v}_i \text{ may belong to object }) \\ -1 & \text{otherwise } (\mathbf{v}_i \text{ does not belong to object }) \end{cases}$$







## **Object Tracking**

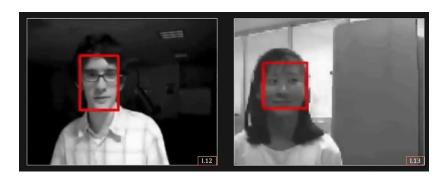
- 3. For each Search Window W:
   a. Compute  $\varphi(W) = \sum C(\mathbf{v}_i)$  for all features  $\mathbf{v}_i$  inside W.
   b. Compute a heuristic  $\tau(W, W_{t-1})$  that penalizes large deviations from previous location, size and shape  $W_{t-1}$ .
   C. Compute Match Score  $\mu(W) = \varphi(W) \tau(W, W_{t-1})$
- 4. Select window  $W_t$  with the best match score as new object location.
- 5. Update object appearance model:  $O_t = O_{t-1} \cup \{\mathbf{v}_i\} \forall \mathbf{v}_i$  inside  $W_t$  such that  $C(\mathbf{v}_i) = +1$ .







#### **Tracking Results: Scale and Orientation**



Resilient to changes in scale and orientation.



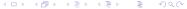


# **Tracking Results: Occlusion**

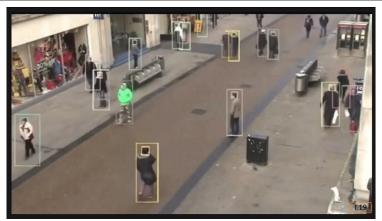


Resilient to occlusion.





# **Tracking Applications**

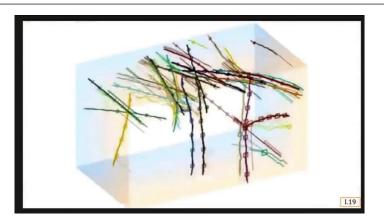


Tracking people in the wild.





# **Tracking Applications**



Tracking people in the wild.





# **Tracking Applications**



Traffic Monitoring.





#### **Tracking Applications**



Customer Behavior for In-Store Analytics.







#### **3D Face Reconstruction**

张举勇 中国科学技术大学

# 背景:数字世界

#### 真实世界



时空约束限制了工作、生活、娱乐等方面的需求

#### 数字世界



无限拓展想象力与创造力



#### 三维数字内容建模与生成

 对物理世界进行高效高保真数字化是支撑VR、AR、 元宇宙等上层应用的核心基础

微软-Fusion4D系统



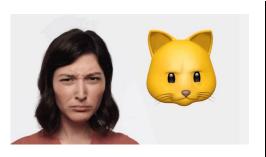
#### **Meta-Horizon Workrooms**







## 背景一人脸重建应用

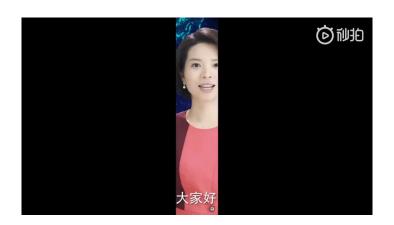








## 背景一人脸重建应用







## 背景一人脸重建应用





# 背景一数字交流







### 基于相机阵列的数字人建模





- ₩ 高精度建模效果
- 可恢复材质、光照等
- ❷ 受控的采集环境、昂贵的价格
- 会复杂的制作流程





#### 基于稀疏视角的数字人建模



Sparse view input

- 成本与便捷性得到极大提高
- 会 对于普通大众仍遥不可及





#### 愿景: 基于单目相机的数字人建模与驱动





# 基于深度学习的实时单目三维人脸重建

# 研究问题







Input Output



## 主要想法

#### 基于逆向渲染的逼真人脸图片合成





逆向 渲染



90













### 三维人脸表示



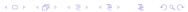












### 三维人脸参数化表示

几何主成分 1st. (+5) 2nd. (+5) 3rd. (+5)







 $\begin{aligned} \mathbf{p} &= \bar{\mathbf{p}} + \mathbf{A}_{id} \boldsymbol{\alpha}_{id} + \mathbf{A}_{exp} \overline{\boldsymbol{\alpha}_{exp}} \\ \mathbf{b} &= \bar{\mathbf{b}} + \mathbf{A}_{alb} \boldsymbol{\alpha}_{alb} \end{aligned}$ 

反照率主成分 1st. (+5) 2nd. (+5) 3rd. (+5)







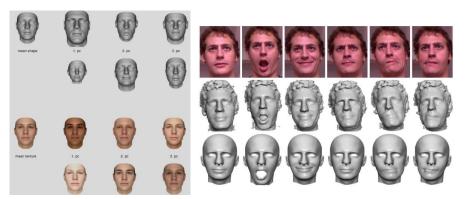








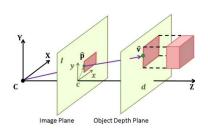
# 一些三维人脸参数化模型



3DMM FaceWarehouse

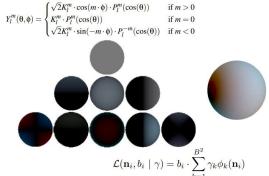


#### 逆向渲染-渲染过程



$$\mathbf{q}_i = s \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \end{pmatrix} R \mathbf{p}_i + \mathbf{t}$$

相机投影







光照模型

#### 逆向渲染-优化过程



$$E(\chi) = E_{\rm con} + w_l E_{\rm lan} + w_r E_{\rm reg}$$

$$E_{\rm con}(\chi) = \frac{1}{|\mathcal{F}|} \|I_{\rm ren} - I_{\rm in}\|^2$$

$$E_{\text{lan}}(\chi) = \frac{1}{|\mathcal{L}|} \sum_{i \in \mathcal{L}} \|\mathbf{q}_i - (\Pi R \mathbf{p}_i + \mathbf{t})\|^2$$

$$\chi = \{ \alpha_{id}, \alpha_{exp}, \alpha_{alb}, s, pitch, yaw, roll, t, r \}$$





# 逆向渲染-几何细节











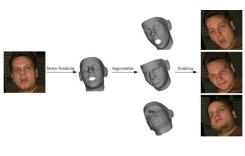
$$E(\mathbf{d}) = E_{\text{con}} + \mu_1 \|\mathbf{d}\|_2^2 + \mu_2 \|\triangle \mathbf{d}\|_1$$



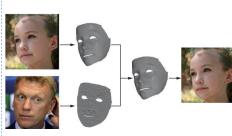


### 构造训练数据

#### CoarseData



#### FineData

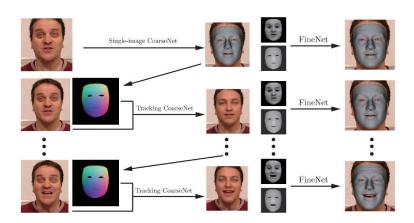


$$\begin{split} & \min_{\widetilde{\mathbf{d}}_t} & \sum_{(i,j) \in \Omega} \|\nabla \widetilde{\mathbf{d}}_t(i,j) - \mathbf{w}(i,j)\|^2, \\ & \text{s.t.} & \widetilde{\mathbf{d}}_t(i,j) = \mathbf{d}_t(i,j) \qquad (i,j) \in \partial \Omega \end{split}$$





## 神经网络流程







#### 实验结果







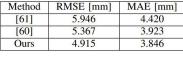


















[Jackson et al.] ICCV 17



Ours

Quantitative comparison results with other methods on FRGC dataset





# 应用-川剧变脸



输入视频

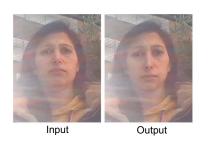


输出视频



# 基于前置摄像头的实时人脸视角矫正

# 研究问题

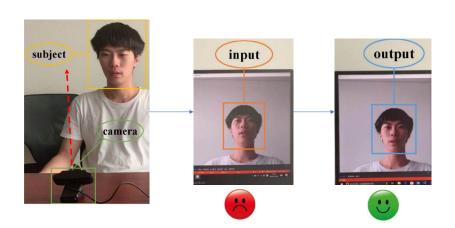




没有正对相机→ 正对相机



# 研究意义







#### 问题挑战

- 映射挑战:输入输出间的复杂映射
- 融合挑战: 如何融合原始背景和新的前景





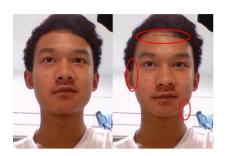
本质上为三维映射,基于二维映射的方法会失败

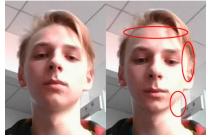




### 问题挑战

- 映射挑战: 输入输出间的复杂映射
- 融合挑战: 如何融合原始背景和新的前景









# 映射-3D



Input





Synthesized with virtual camera



# 映射-3D



Input



Dense Track



Virtual Overlay





#### 融合-割缝优化



Outside the seam: content from the original image *I* 

Inside the seam: content from the rendered face  ${\it J}$ 

$$\begin{split} E_{\text{seam}} &= \sum_{(\boldsymbol{x}, \boldsymbol{y}) \in \mathcal{P}} \alpha(\boldsymbol{x}, \boldsymbol{y}) \cdot (\|\boldsymbol{I}(\boldsymbol{x}) - \boldsymbol{J}(\boldsymbol{x})\|_2 \\ &+ \|\boldsymbol{I}(\boldsymbol{y}) - \boldsymbol{J}(\boldsymbol{y})\|_2) \end{split}$$

 ${\cal P}$  : adjacent pixels across the seam







## 融合-Laplacian融合



Input



Overlay



Seam Optimization



# 融合-Laplacian融合













Laplacian Blending



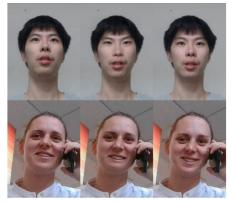
# 实验结果





(b) Giger et al. (2014)

(c) Ours



(a) Input

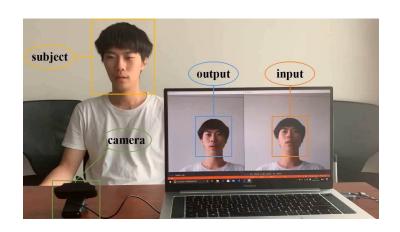
(b) Giger et al. (2014)

(c) Ours





# 应用-实际效果







#### 背景:神经隐式函数

$$F_{\theta}(\mathbf{x}) = (y_1, y_2, \cdots, y_k)$$





occupancy, SDF, color, ...



Image regression  $(x,y) \rightarrow RGB$ 



Shape regression  $(x,y,z) \rightarrow \text{occupancy}$ 

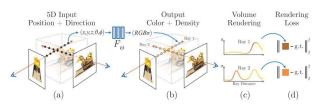


Neural rendering  $(x,y,z) \rightarrow density,RGB$ 





#### 神经辐射场



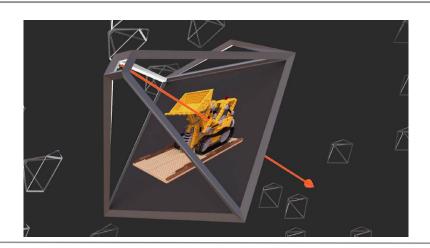
- 隐式表示: F<sub>θ</sub>: (X, D) → (c, σ)
   X:空间点, D:观测X的view direction
   c: 预测的点X的颜色, σ:预测的点X的density
- 体渲染:  $\hat{C}(\mathbf{r}) = \sum_{i=1}^{N} T_i (1 \exp(-\sigma_i \delta_i)) \mathbf{c}_i \,, \text{ where } T_i = \exp\left(-\sum_{j=1}^{i-1} \sigma_j \delta_j\right)$

NeRF: Representing Scenes as Neural Radiance Fields for View Synthesis Mildenhall et al. ECCV 2020





# 神经辐射场

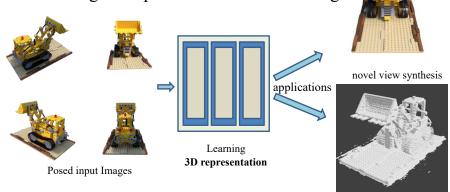






#### 神经隐式表示与逆向渲染

• Learning 3D representation from 2D images



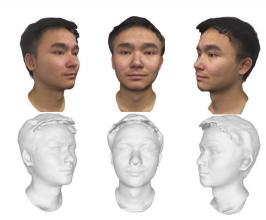




# 单手机高精度人头重建-静态



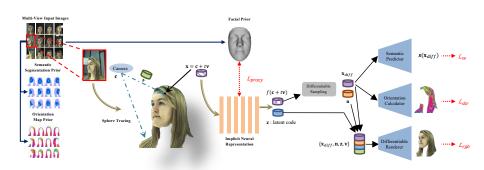
单个手机拍摄



全自动重建高保真头部三维模型



#### 先验引导的神经隐式重建

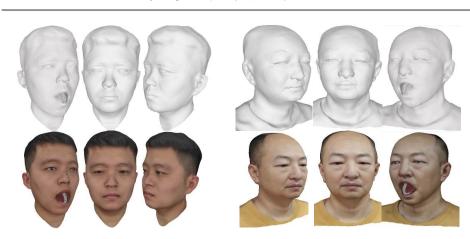


Prior-Guided Multi-View 3D Head Reconstruction IEEE Transactions on Multimedia (TMM), 2021





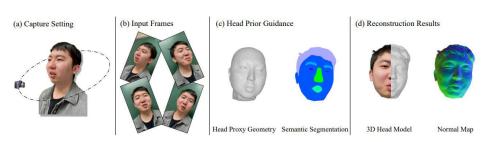
# 更多结果展示





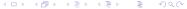


#### 单目可驱动高精度人头重建



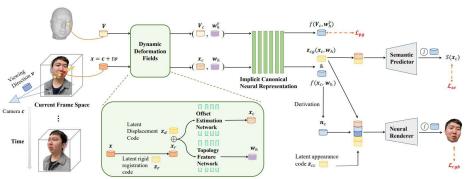
HeadRecon: High-Fidelity 3D Head Reconstruction from Monocular Video





#### 算法流程

- 输入: 单目说话视频
- 输出: 可驱动的高精度三维人头模型







# 结果展示















渲染视频

渲染法向

重建网格

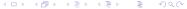


## 单目自转视频的三维人体重建



SelfRecon: Self Reconstruction Your Digital Avatar from Monocular Video CVPR, Oral Presentation, 2022





## 期望的三维人体重建方式

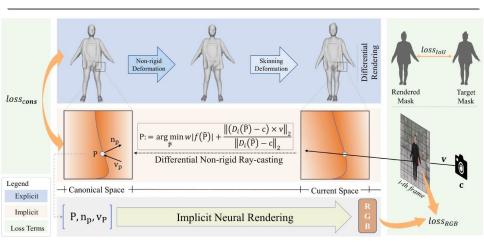


- ✓ Single camera
- ✓ Easy to capture
- ✓ High-fidelity result





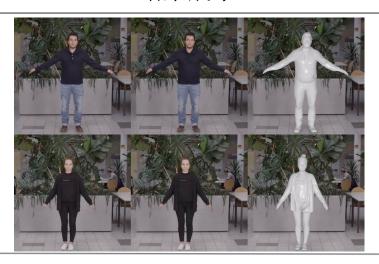
#### 算法:整体流程







# 结果展示







## 应用至任意人体动作序列







# 非刚性形变的可视化展示



Input video



Non-rigid deformation

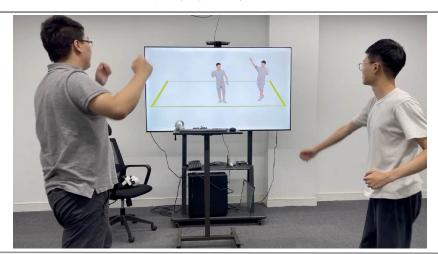


Whole deformations





# 应用展示

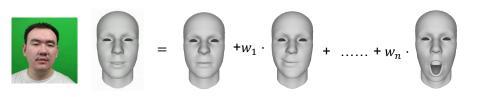






#### 我们真的需要几何模型吗?

- 现有的数字人系统通常先对头部进行高精几何建模
- 例如,采用Blendshape表示来实现人脸驱动等



Reconstructing Personalized Semantic Facial NeRF Models From Monocular Video Conditionally Accepted to the Journal Track of SIGGRAPH Asia, 2022





#### 基于NeRF表示的个性化人头参数化

• 不同于3DMM等基于网格的三维几何参数化表示,我们采用基于NeRF的三维神经渲染参数化表示







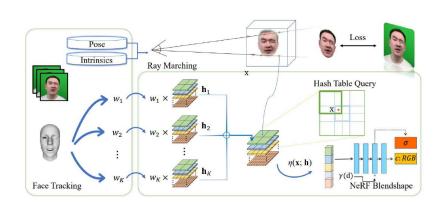
# 拍摄输入的单目视频数据







## 算法流程





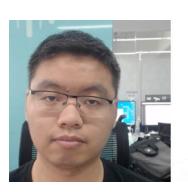


# 训练过程展示





## 跨身份表情驱动







😀 实时计算 😀 微表情迁移





## 研究问题



With audio and pose input only, AD-NeRF could synthesize high-fidelity talking head video





## 现有的语音驱动数字人

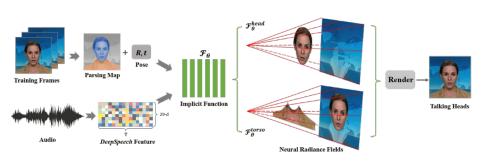






#### 主要想法

#### 利用神经辐射场直接学习语音信号到说话人视频的映射

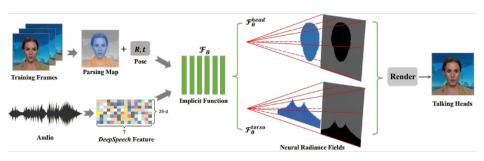


分辨率高、无需中间模态、支持姿态改变





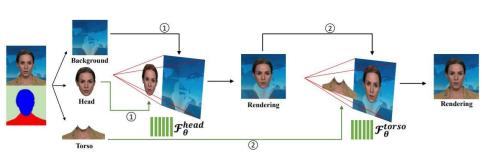
## 算法流程







## 分离的NeRF场



头部和躯干运动的不一致性





## 实验结果-对比实验

#### Comparisons with Other Methods





## 实验结果-分离NeRF场



Ground-Truth



w.o. Individual Training



w. Individual Training





#### 语音驱动高保真数字人应用







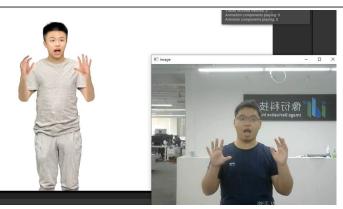
虚拟主持人/虚拟讲解员

电商虚拟主播

虚拟培训分身



## 完整数字人建模与驱动展示



相关应用: 沉浸式视频会议、远程教育等





#### 总结与展望

- 基于新的表示方式、端到端可微优化框架,数字人 建模变得更便捷、高效、高保真
- 便捷、高效、高保真仍需不断提高
  - 单目设备在采集光照、角度、表情与动作幅度的鲁棒性
  - 移动端上建模
  - 进一步提高时空信息之间的精准融合



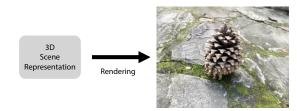




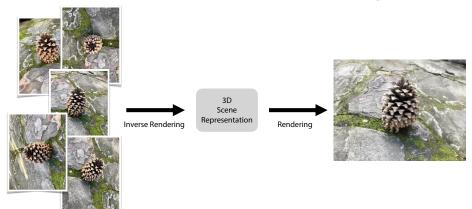
#### **Neural Radiance Fields**

张举勇 中国科学技术大学

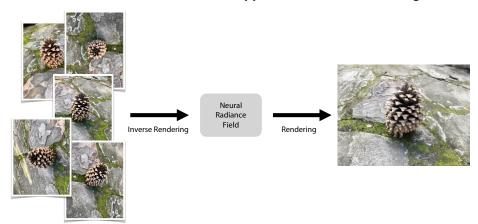
#### Computer vision as inverse rendering



#### Computer vision as inverse rendering



#### Neural Radiance Fields (NeRF) as an approach to inverse rendering



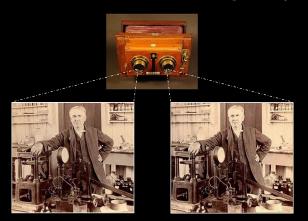
#### Deep learning for 3D reconstruction

- Previously: we reconstruct geometry by running stereo or multi-view stereo on a set of images
  - "Classical" approach
- How can we leverage powerful tools of deep learning?
  - Deep neural networks
  - GPU-accelerated stochastic gradient descent

#### NeRF and related methods – Key ideas

- We need to create a loss function and a scene representation that we can optimize using gradient descent to reconstruct the scene
- Differentiable rendering

## **Side Topic: Stereo Photography**



# **Stereo Photography**





#### Viewing Devices







## **Stereo Photography**

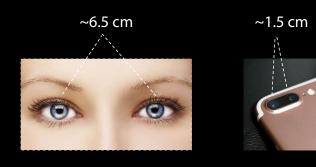


Queen Victoria at World Fair, 1851

# **Stereo Photography**



#### **Issue: Narrow Baseline**

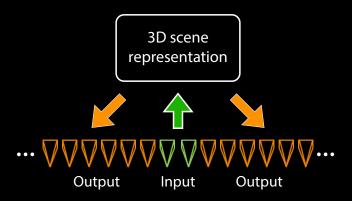




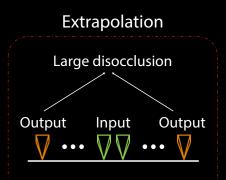




#### **Problem Statement**



#### **Challenges**

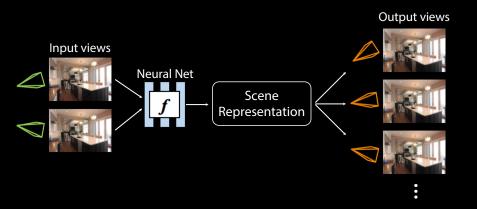


#### **Non-Lambertian Effects**

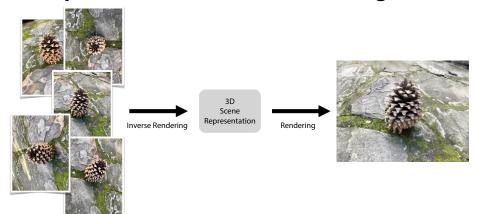
Reflections, transparencies, etc.



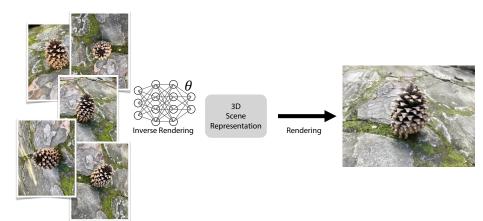
#### **Neural prediction of scene representations**



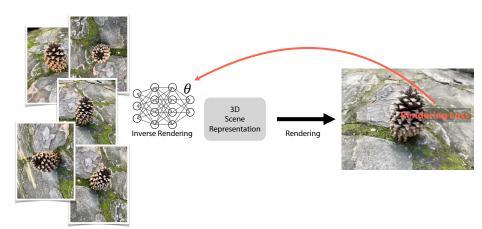
### Computer vision as inverse rendering



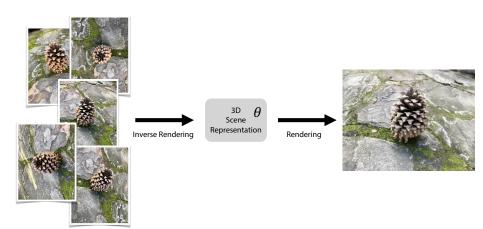
### Paradigm 1: "Feedforward" inverse rendering



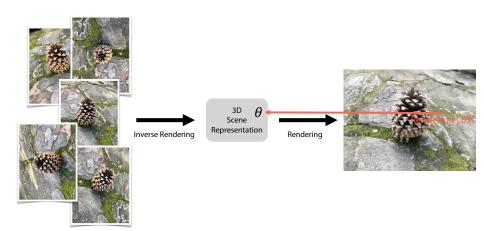
#### Paradigm 1: "Feedforward" inverse rendering



#### Paradigm 2: "Render-and-compare"



#### Paradigm 2: "Render-and-compare"



### What representation to use?

- Could use triangle meshes, but hard to differentiate during rendering
- Multiplane images (MPIs) are easy to differentiate, but only allow for rendering a small range of views



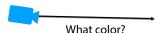


# NeRF == Differentiable Rendering with a **Neural Volumetric Representation**



Barron et al 2021, Mip-NeRF 360: Unbounded Anti-Aliased Neural Radiance Fields

querying the radiance value along rays through 3D space



continuous, differentiable rendering model without concrete ray/surface intersections



using a neural network as a scene representation, rather than a voxel grid of data

$$(x, y, z) \Rightarrow$$
 Scene properties

Multi-layer Perceptron (Neural Network)

# NeRF: Representing Scenes as Neural Radiance Fields for View Synthesis FCCV 2020



Ben Mildenhall\*



















Google

UC San Diego

**UC Berkeley** 



Given a set of sparse views of an object with known camera poses



3D reconstruction viewable from any angle

#### **NeRF Overview**

- Volumetric rendering
- Neural networks as representations for spatial data
- ► Neural Radiance Fields (NeRF)

#### **NeRF Overview**

- Volumetric rendering
- Neural networks as representations for spatial data
- Neural Radiance Fields (NeRF)



#### Traditional volumetric rendering

 Theory of volume rendering co-opted from physics in the 1980s: absorption, emission, out-scattering/in-scattering



Ray tracing simulated cumulus cloud [Kajiya]

Chandrasekhar 1950, Radiative Transfer Kajiya 1984, Ray Tracing Volume Densities

#### Traditional volumetric rendering



Medical data visualisation [Levoy]



Alpha compositing [Porter and Duff]

- Theory of volume rendering co-opted from physics in the 1980s: absorption, emission, out-scattering/in-scattering
- Adapted for visualising medical data and linked with alpha compositing

Chandrasekhar 1950, Radiative Transfer

Levoy 1988, Display of Surfaces from Volume Data Max 1995, Optical Models for Direct Volume Rendering Porter and Duff 1984. Compositing Digital Images

#### Traditional volumetric rendering



Physically-based Monte Carlo rendering [Novak et al]

- Modern path tracers use sophisticated Monte Carlo methods to render volumetric effects

Novak et al 2018, Monte Carlo methods for physically based volume rendering

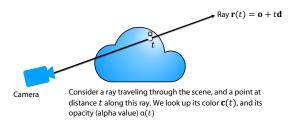
#### Volumetric formulation for NeRF



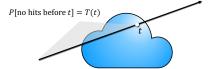
Scene is a cloud of colored fog



#### Volumetric formulation for NeRF



#### Volumetric formulation for NeRF



But t may also be blocked by earlier points along the ray. T(t): probability that the ray didn't hit any particles earlier. T(t) is called "transmittance"

#### Volume rendering estimation: integrating color along a ray

Rendering model for ray  $\mathbf{r}(t) = \mathbf{o} + t\mathbf{d}$ :

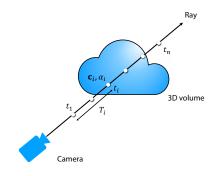
$$\mathbf{c} pprox \sum_{i=1}^{n} T_i \alpha_i \mathbf{c}_i$$
final rendered color along ray weights colors

How much light is blocked earlier along ray:

$$T_i = \prod_{j=1}^{i-1} (1 - \alpha_j)$$

Computing the color for a set of rays through the pixels of an image yields a rendered image





#### Volume rendering estimation: integrating color along a ray

Rendering model for ray  $\mathbf{r}(t) = \mathbf{o} + t\mathbf{d}$ :

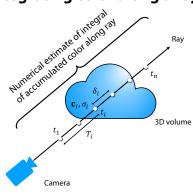
$$\mathbf{c} pprox \sum_{i=1}^{n} T_i \alpha_i \mathbf{c}_i$$
final rendered color along ray weights colors

How much light is blocked earlier along ray:

$$T_i = \prod_{j=1}^{i-1} (1 - \alpha_j)$$

Slight modification:  $\alpha$  is not directly stored in the volume, but instead is derived from a stored volume density sigma ( $\sigma$ ) that is multiplied by the distance between samples delta ( $\delta$ ):

$$\alpha_i = 1 - \exp(-\sigma_i \delta_i)$$



#### Volume rendering estimation: integrating color along a ray

Rendering model for ray  $\mathbf{r}(t) = \mathbf{o} + t\mathbf{d}$ :

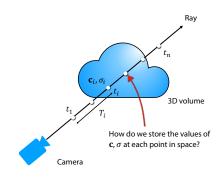
$$\mathbf{c} pprox \sum_{i=1}^{n} T_i \alpha_i \mathbf{c}_i$$
final rendered color along ray weights

How much light is blocked earlier along ray:

$$T_i = \prod_{j=1}^{i-1} (1 - \alpha_j)$$

Computing the color for a set of rays through the pixels of an image yields a rendered image

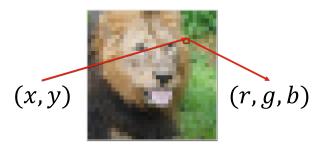




#### **NeRF** Overview

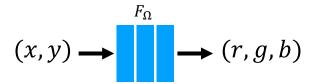
- Volumetric rendering
- Neural networks as representations for spatial data
- Neural Radiance Fields (NeRF)

#### Toy problem: storing 2D image data



Usually we store an image as a 2D grid of RGB color values

#### Toy problem: storing 2D image data



What if we train a simple fully-connected network (MLP) to do this instead?

### **Recall the TensorFlow playground**



Same concept as before, except we are computing an image, instead of a classifier!

# Naive approach fails!



Ground truth image



Neural network output fit with gradient descent

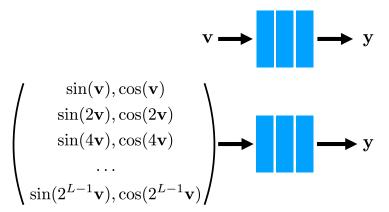
#### Problem:

"Standard" coordinate-based MLPs cannot represent high frequency functions

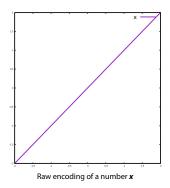
#### Solution:

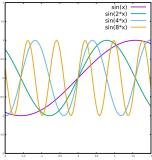
Pass input coordinates through a high frequency mapping first

#### Example mapping: "positional encoding"



# **Positional encoding**





"Positional encoding" of a number  $\boldsymbol{x}$ 

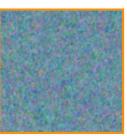
#### Problem solved!



Ground truth image

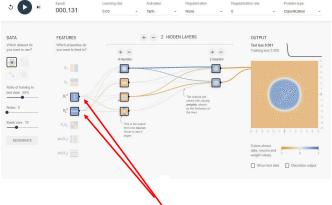


Neural network output without high frequency mapping



Neural network output with high frequency mapping

# Sometimes a better input encoding is all you need

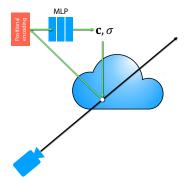


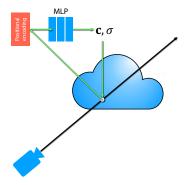
Recall "squared" encoding in TensorFlow Playground

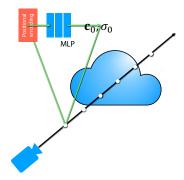
#### **NeRF** Overview

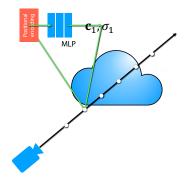
- Volumetric rendering
- Neural networks as representations for spatial data
- Neural Radiance Fields (NeRF)

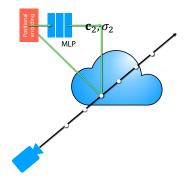
# NeRF = volume rendering + coordinate-based network

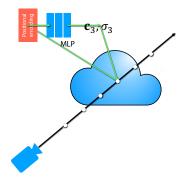


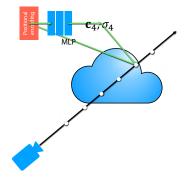


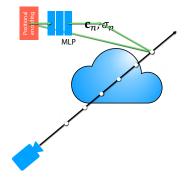




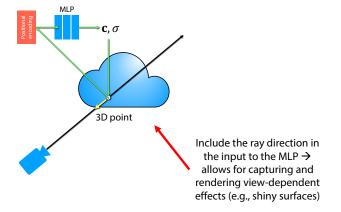




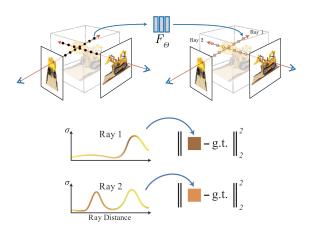




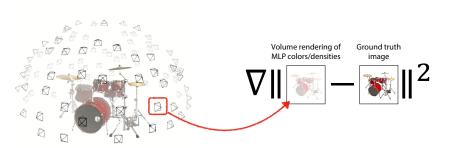
### Extension: view-dependent field



# Putting it all together



# Train network using gradient descent to reproduce all input views of scene



# Results



# NeRF encodes convincing view-dependent effects using directional dependence



# NeRF encodes convincing view-dependent effects using directional dependence



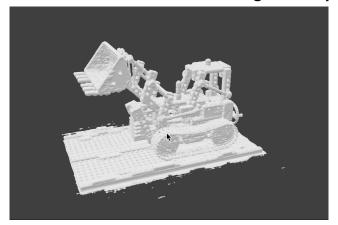
#### NeRF encodes detailed scene geometry with occlusion effects



#### NeRF encodes detailed scene geometry with occlusion effects



# NeRF encodes detailed scene geometry



### Summary

- Represent the scene as volumetric colored "fog"
- Store the fog color and density at each point as an MLP mapping 3D position (x, y, z) to color c and density σ
- Render image by shooting a ray through the fog for each pixel
- Optimize MLP parameters by rendering to a set of known viewpoints and comparing to ground truth images